

Particle Physics

The Basics

- Particle Physics arises from the combination of special relativity and quantum mechanics

Particles are described by a list of properties:

- Mass, a positive number or zero, describing the minimum energy of the particle
 - Always given in metric multiples of eV/c^2 , like MeV/c^2 and GeV/c^2
- Spin, which describes the internal angular momentum of the particle
 - Written as $s\hbar$, but we abbreviate this by just giving s , where $s > 0$
 - s is always an integer (0,1,2,3,...) or half-integer (1/2,3/2,5/2,...)
- Electric charge, which is a multiple of the fundamental charge e
 - We normally give just Q , and the charge is Qe
 - Q is an integer. It can be positive, negative, or zero
- Other properties exist, which we will discuss as they come up

$$E^2 = (\vec{p}c)^2 + (mc^2)^2$$

Fermions and Bosons

- *Fermions* are particles that obey the Pauli Exclusion Principle
 - You can't put two of the same kind in the same quantum state
 - Fermions always have half-integer spin
- *Bosons* are particles that violate the Pauli Exclusion Principle
 - They actually *prefer* being in the same quantum state
 - Bosons always have integer spin

Some Particles (masses in MeV/c²)

	<u>Name</u>	<u>Sym.</u>	<u>Spin</u>	<u>Q</u>	<u>Mass</u>
Fermions	Proton	p ⁺	1/2	+1	938.27
	Neutron	n ⁰	1/2	0	939.57
	Electron	e ⁻	1/2	-1	0.511
	Neutrino	ν	1/2	0	≲ 2×10 ⁻⁶
Bosons	Photon	γ	1	0	0
	Pi-plus	π ⁺	0	+1	139.57
	Pi-zero	π ⁰	0	0	134.98
	Pi-minus	π ⁻	0	-1	139.57

Anti-Particles

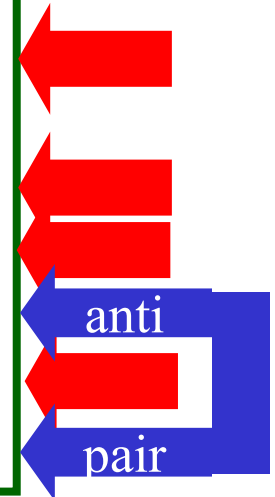
- All particles have anti-particles
- Anti-particles have the same mass and spin, but opposite charge
- Usually named by prefixing with “anti-”
- Some particles are their own anti-particles

For each of the particles below

- What is the spin, charge, and mass of the anti-particle
- Which might be their own anti-particles?
- Which might be anti-particles of each other?

<u>Name</u>	<u>Sym.</u>	<u>Spin</u>	<u>Q</u>	<u>Mass</u>
Proton	p^+	$\frac{1}{2}$	+1	938.27
Neutron	n^0	$\frac{1}{2}$	0	939.57
Electron	e^-	$\frac{1}{2}$	-1	0.511
Neutrino	ν	$\frac{1}{2}$	0	$\lesssim 2 \times 10^{-6}$
Photon	γ	1	0	0
Pi-plus	π^+	0	+1	139.57
Pi-zero	π^0	0	0	134.98
Pi-minus	π^-	0	-1	139.57

<u>Spin</u>	<u>Q</u>	<u>Mass</u>
$\frac{1}{2}$	-1	938.27
$\frac{1}{2}$	0	939.57
$\frac{1}{2}$	+1	0.511
$\frac{1}{2}$	0	$\lesssim 2 \times 10^{-6}$
1	0	0
0	-1	139.57
0	0	134.98
0	+1	139.57



Conservation Laws

Energy and Momentum

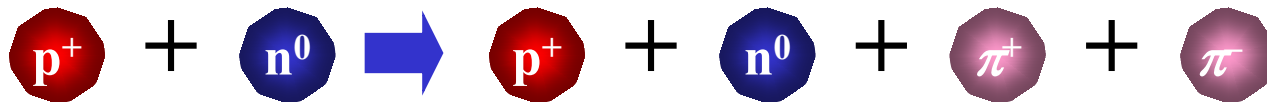
- Energy and Momentum are conserved
- We'll use only energy conservation



- Consider a frame where the initial proton is at rest

$$E_i = E_p = m_p c^2 \qquad E_f = E'_p + E_\pi \geq m_p c^2 + m_\pi c^2 > m_p c^2$$

- Is the following interaction possible?



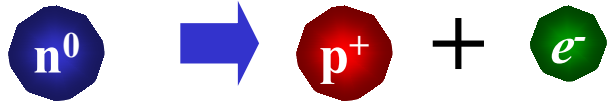
- Energy doesn't preclude it because the particles on the left can have kinetic energy in addition to their rest energy
 - There is not necessarily any frame where these particles are at rest

- Energy
- Momentum
- Angular Momentum
- Electric Charge
- Baryon Number
- Strangeness

For decays only: Mass of initial particle must exceed sum of masses of final particles

Angular Momentum

- Total angular momentum is conserved



- Consider angular momentum around z-axis

$$L_{nz} + S_{nz} = L_{pz} + S_{pz} + L_{ez} + S_{ez}$$

- All of the orbital angular momenta (L 's) are integer multiples of \hbar
- Because the neutron, proton and electron are all fermions, the internal angular momenta (S 's) are all half-integer multiples of \hbar

$$n_1 \hbar \pm \frac{1}{2} \hbar = n_2 \hbar \pm \frac{1}{2} \hbar + n_3 \hbar \pm \frac{1}{2} \hbar$$

$$n_1 \pm \frac{1}{2} = n_2 \pm \frac{1}{2} + n_3 \pm \frac{1}{2}$$

- Right side is an integer, left side is not

- Energy
- Momentum
- Angular Momentum
- Electric Charge
- Baryon Number
- Strangeness

Total number of fermions (particles with half-integer spin) on left plus right must be even

Electric Charge

- Electric charge is conserved

Charge is
conserved

Why is the electron stable?



- Energy
- Momentum
- Angular Momentum
- Electric Charge
- Baryon Number
- Strangeness

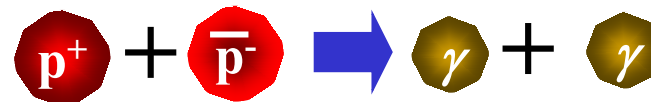
- By energy conservation, whatever is on the right must be lighter
- By charge conservation, something on the right must be charged
- No such particle exists, so electron is stable

Baryon Number

- Consider nuclear reactions
 - β^- decay: $(Z,A) \rightarrow (Z+1,A)$
 - β^+ decay: $(Z,A) \rightarrow (Z-1,A)$
 - α decay: $(Z,A) \rightarrow (Z-2,A-4) + (2,4)$
 - γ decay: $(Z,A)^* \rightarrow (Z,A)$

- Total protons plus neutrons (call this baryons) remains constant?

- Maybe anti-protons count as negative baryons?



- Maybe there are other particles that count as baryons too?
- There are a group of particles called *baryons*
 - They each have baryon number +1
- For every baryon, there is an anti-baryon
 - They each have baryon number -1

- Energy
- Momentum
- Angular Momentum
- Electric Charge
- Baryon Number**
- Strangeness

**Baryon number
is conserved**

Why is the proton stable?

- There is no lighter **baryon**

Baryons, Anti-Baryons, and Mesons

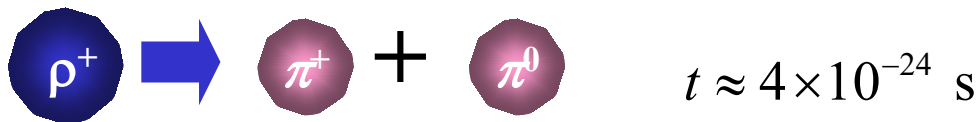
- The strong nuclear force is what holds the nucleus together
 - It must be strong and fast to do so
- Some particles (photon, electron, neutrino) do not seem to be affected by it

The particles that feel strong forces come in three categories:

- *Baryons* have baryon number +1
- *Anti-baryons* are their anti-particles and have baryon number -1
- *Mesons* have baryon number 0
 - Anti-mesons are mesons

Reactions that occur very quickly are believed to be mediated by this strong force

- The rho-mesons, for example, decay very fast



A strong interaction

- The kaons, by comparison, decay very slowly



A weak interaction

Strangeness

- Why do some reactions involving strongly interacting particles occur so slowly?
- It was speculated that some baryons and mesons had a property called *strangeness* that also had to be violated only in weak interactions

- Energy
- Momentum
- Angular Momentum
- Electric Charge
- Baryon Number
- Strangeness

**Strangeness is conserved
in all interactions except
weak interactions**

Important notes:

- Strangeness *only* applies to strongly interacting particles; other particles have $S = 0$
- Strangeness can *only* be changed by weak interactions
- The strangeness of an anti-particle is the opposite of the strangeness of the particle

<u>Symbol</u>	<u>S</u>
ρ^+, ρ^0, ρ^-	0
π^+, π^0, π^-	0
K^+, K^0	+1
K^-, \bar{K}^0	-1
p^+, n^0	0
$\Sigma^+, \Sigma^0, \Sigma^-$	-1
Λ^0	-1
Ξ^0, Ξ^-	-2

Types of Interactions

THE STRONG FORCE

- Involves only strongly interacting particles: baryons, anti-baryons, and mesons
- Conserves strangeness

ELECTROMAGNETISM

- Affects all charged particles
- Always involves photons (though this isn't always obvious)



e^-



e^+



γ



μ^+

THE WEAK FORCE

- Affects essentially all particles except photons
- The only force that affects neutrinos
- The only force that violates strangeness

Which force is at work in a given reaction?

- The stronger a force is, the more likely it is to be at work
 - Strong > Electromagnetism > Weak

Which Force?

A step-by-step procedure for determining which force is at work

- If charge conservation is violated → Impossible
- Else if baryon number is violated → Impossible
- Else if odd # fermions (left + right) → Impossible
- Else if decay and insufficient energy → Impossible
- Else if strangeness violated → Weak
- Else if all particles are strong → Strong
- Else if neutrinos are involved → Weak
- Else → Electromagnetism

Sample Problems

Classify the
reactions below:



Strong

- Charge: $(+1) + (-1) = 0 + 0$ ✓
- Baryons: $(+1) + 0 = (+1) + 0$ ✓
- Fermions: $[1+0] + [1+0] = 2 = \text{even}$ ✓
- Not a decay ✓
- Strangeness: $0 + (-1) = (-2) + (+1)$ ✓
- All particles are strong ✓



Electromagnetism

- Charge: $0 = 0 + (+1) + (-1)$ ✓
- Baryons: $(+1) = (+1) + 0 + 0$ ✓
- Fermions: $[1] + [1+1+1] = 4 = \text{even}$ ✓
- Energy: $1193 > 1116 + 0.5 + 0.5$ ✓
- Strangeness: $1 = 1 + 0 + 0$ ✓
- All particles are strong ✗
- Neutrinos are involved ✗



(the \bar{n}^0 and the \bar{p}^- are the anti-
particles of the neutron and
proton)

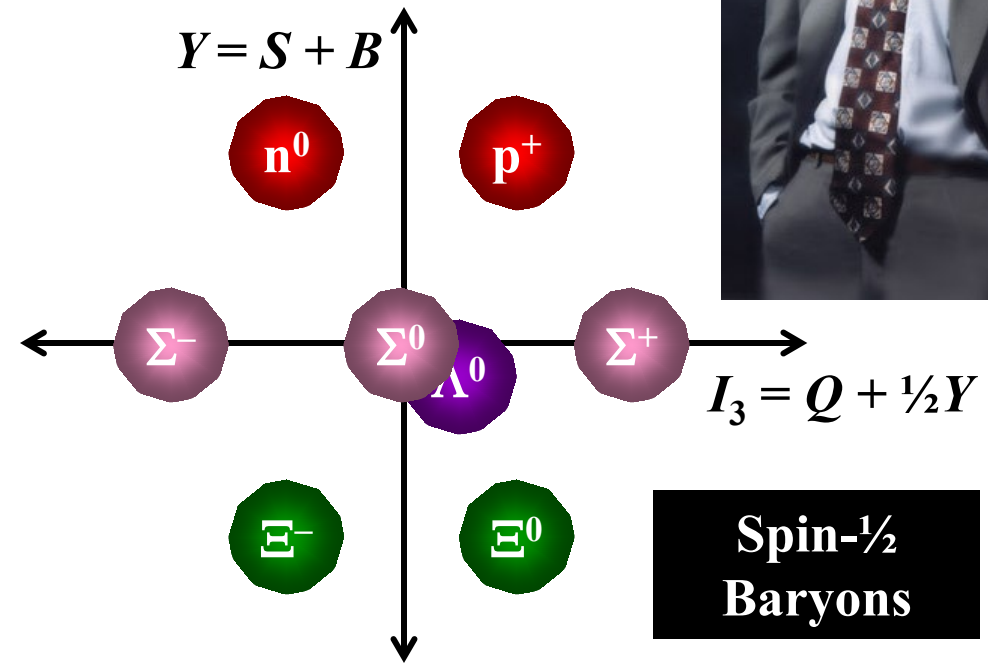
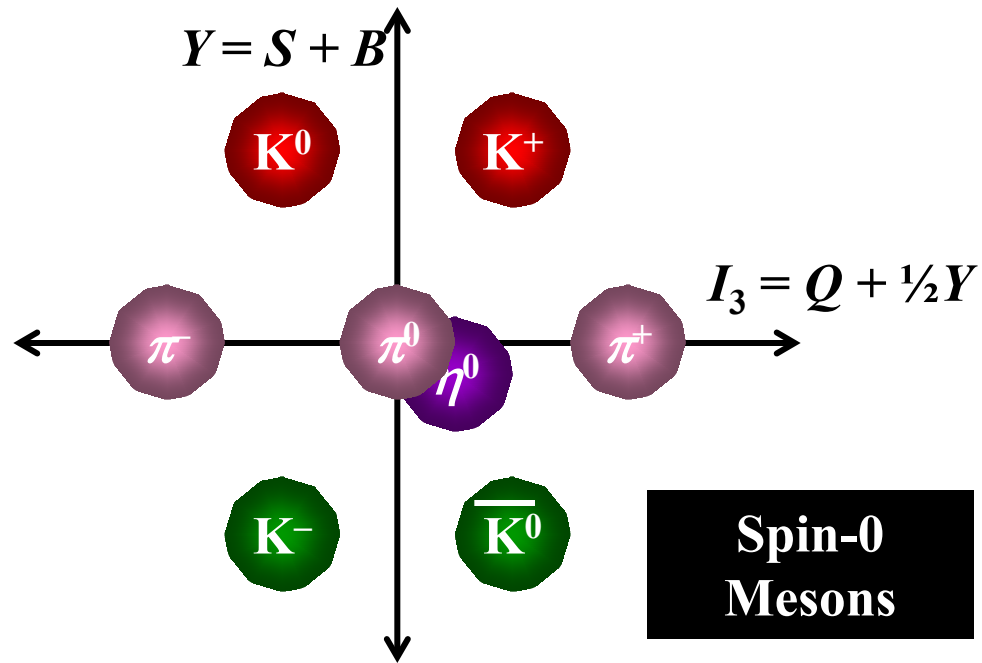
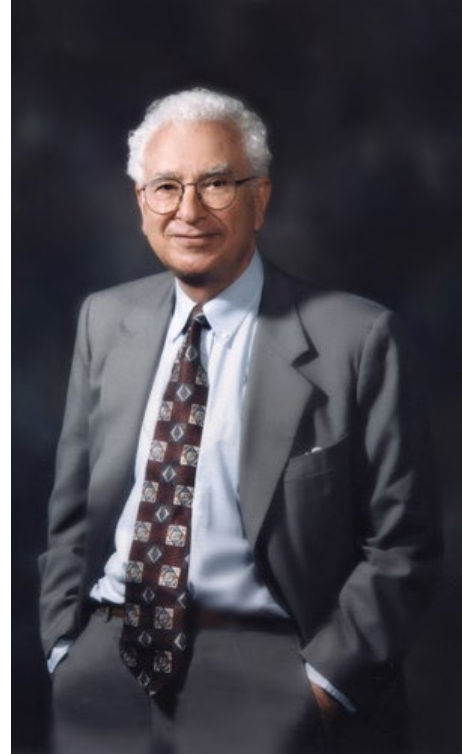
- Charge: $0 = (-1) + (+1)$ ✓
- Baryons: $(-1) = (-1) + 0$ ✓
- Fermions: $[1] + [1+1] = 3 = \text{odd}$ ✗

Impossible

The Standard Model

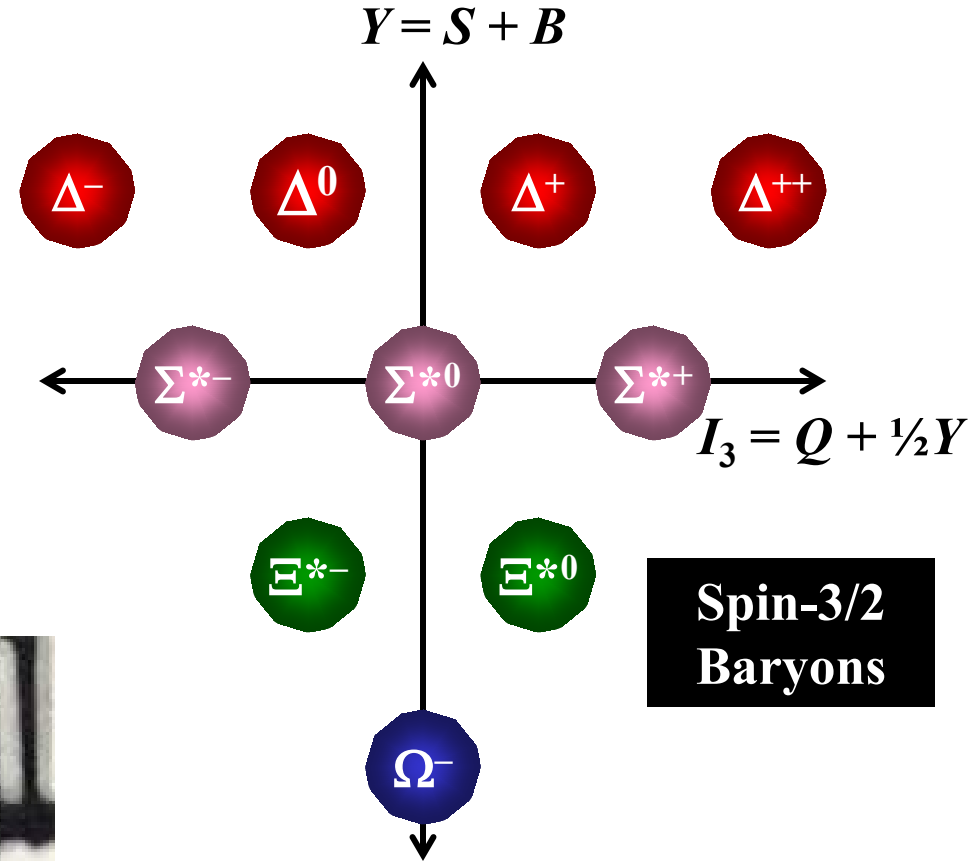
Patterns in Baryons and Mesons




- In the 50's and 60's, the number of baryons and mesons was growing out of control
 - There are currently hundreds known
- In 1961, Murray Gell-Mann noticed a series of mathematical relationships between the various particles






Quarks

- In 1962, based on the patterns, Gell-Mann predicted a new particle, the Ω^-
- In 1964, Gell-Mann and George Zweig independently proposed that all these particles could be explained if there were underlying particles called *quarks*
 - There were three of them, and in baryons, they always come in threes
- There are also anti-quarks for every quark



Quark	Spin	charge	\underline{S}
 Up	$\frac{1}{2}$	$+\frac{2}{3}$	0
 Down	$\frac{1}{2}$	$-\frac{1}{3}$	0
 Strange	$\frac{1}{2}$	$-\frac{1}{3}$	-1

anti-Quark	Spin	charge	\underline{S}
 anti-Up	$\frac{1}{2}$	$-\frac{2}{3}$	0
 anti-Down	$\frac{1}{2}$	$+\frac{1}{3}$	0
 anti-Strange	$\frac{1}{2}$	$+\frac{1}{3}$	+1

Baryons and Mesons from Quarks

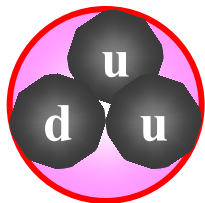
- To make a baryon, combine three quarks
- To make an anti-baryon, combine three anti-quarks
- To make a meson, combine a quark and an anti-quark
- Composition can generally be determined from strangeness and charge

What is a proton made from?

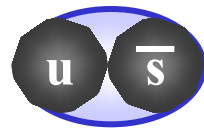
- It is a baryon: three quarks
- It has strangeness 0, so no strange quarks
- It has charge +1, so to get this, must take:




$$+\frac{2}{3} + \frac{2}{3} - \frac{1}{3} = +1$$

$$p^+ = [uud]$$



$$K^+ = [u\bar{s}]$$



	<u>Quark</u>	<u>Spin</u>	<u>charge</u>	<u>S</u>
	Up	$\frac{1}{2}$	$+\frac{2}{3}$	0
	Down	$\frac{1}{2}$	$-\frac{1}{3}$	0
	Strange	$\frac{1}{2}$	$-\frac{1}{3}$	-1

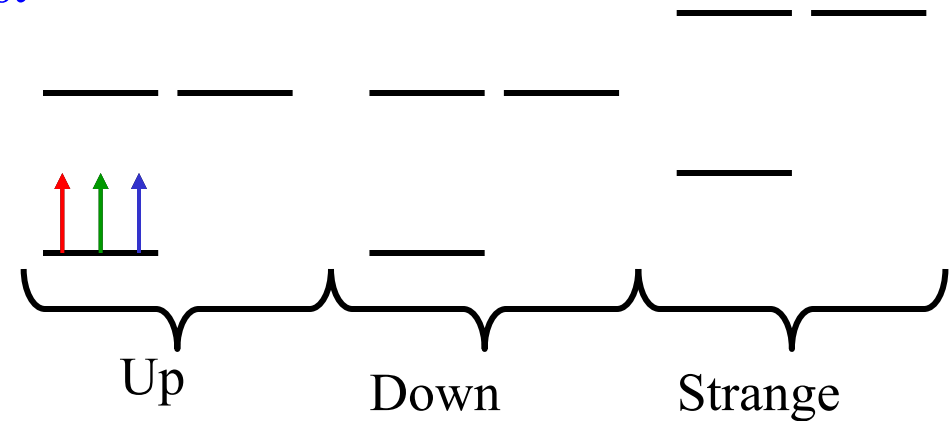
What is a K^+ made from?

- It is a meson: quark + anti-quark
- It has strangeness +1, so must have an anti-strange quark
- To get charge +1, the other quark must have charge $+\frac{2}{3}$

$$-(-\frac{1}{3}) + \frac{2}{3} = +1$$

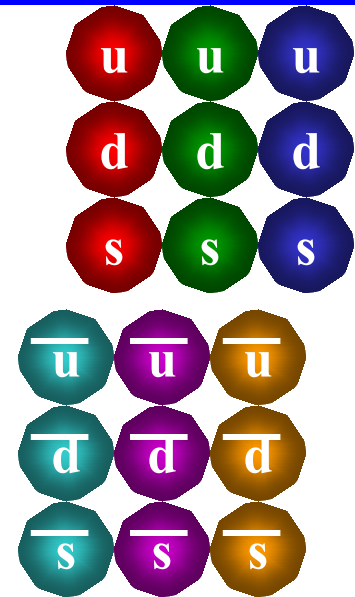
The Problem with Quarks

- Can we “predict” which baryons and mesons are lightest from the quark model?
- How about, say, the Δ^{++} ?
 - Spin 3/2, three up quarks
- Seems to violate Pauli principle
- Some people abandoned the quark model, others, in desperation, dreamed up color



Color

- Maybe there is another property, call it color, that describes an individual quark
- Need three colors: typically called red, green, and blue
- Every type of quark comes in three colors
- You must always combine quarks in colorless combinations
- Anti-quarks come in anti-red, anti-green, and anti-blue
- Everything worked fine, but looked awfully arbitrary



The Secret of the Strong Force

- Where does the arbitrary rule, “make it colorless” come from?
- Consider, by analogy, atoms:
 - Electrons and nuclei have a property called “charge”
 - However, atoms are *almost always* neutral, they are “charge-free”
- This is because there is a force (electromagnetism) mediated by a particle (the photon) that prefers when charges cancel out
- Maybe color is associated with a new force that also prefers colorless combinations



- Particles called “gluons” carry the real strong force back and forth between the three quarks in a baryon or quark and anti-quark in a meson

More About the Strong Force

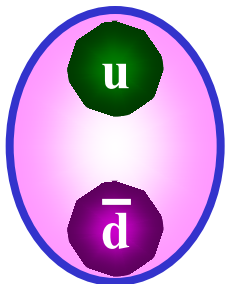
- The real strong force is this force between quarks, mediated by gluons
 - There are eight different gluons in all (don't worry about it)
- The force we have been calling the “strong” force is just a weaker version of it
 - Analogy – nuclear force : strong force :: chemistry : electric force
- All calculations are very difficult involving the strong force
 - “Perturbation theory,” the usual technique, fails
- A few conclusions have been drawn from the theory
 - Quark confinement – quarks never escape
 - The force gets weaker – slowly – at very high energies
 - Only colorless combinations – baryons, mesons, anti-baryons – are possible
 - The type of quark – like strange quarks – weren't changed; this is why strangeness is conserved
- With the advent of modern supercomputers, we are getting good match of theory and experiment

Two Down, One to Go

- The electromagnetic force was the first to be described quantum mechanically
 - Quantum Electrodynamics (QED) is the most accurately tested theory, ever



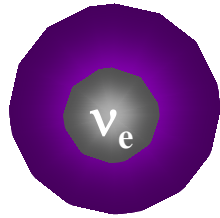
- The strong force was successfully described in terms of colors and gluons
 - Quantum Chromodynamics (QCD) is now pretty well tested



- The weak force was still being worked on
 - Actually, much of this work was simultaneous with strong force

Clues to the Weak Force

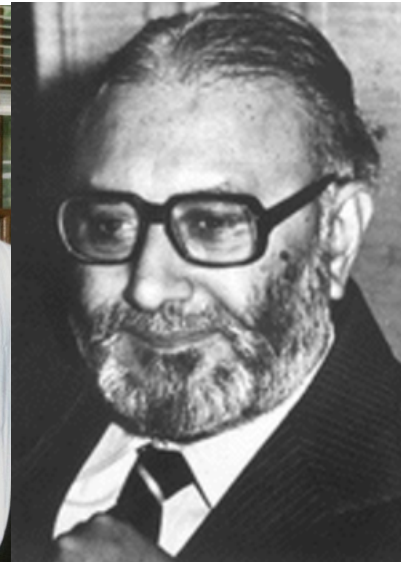
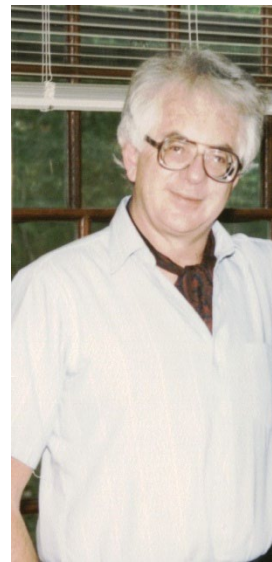
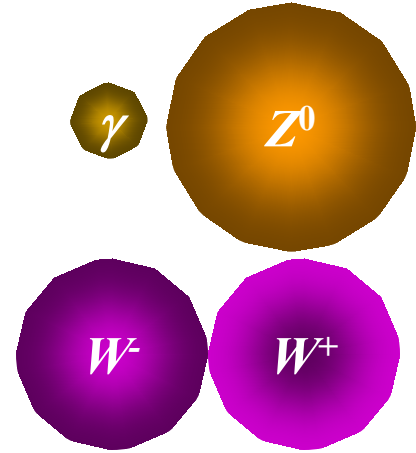
- The weak force changed the nature of particles in a more fundamental way than did the strong or the electromagnetic force
- It had a very short range, which is why it was so weak
- It was guessed it also involved exchange of a particle
 - Called the W^\pm , it was apparently very massive
 - It changed particles into ones with different identities



- Weak interactions changed the electron into an electron neutrino
 - This worked fine
 - These two particles were called leptons
- Another pair of leptons had also been discovered
 - The muon and muon neutrino were two more leptons
 - Just like the electron, but heavier

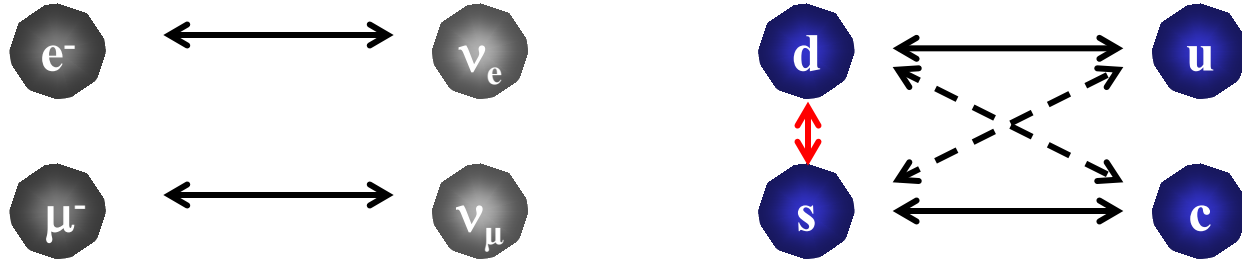
The Electroweak Theory

- During the 1960's, the modern electroweak theory was developed
 - It is a partial unification of the electromagnetic and weak forces
- In 1960 Sheldon Glashow proposed that the theory could be understood if there were also another neutral massive particle called the Z
- There were theoretical problems with this approach
 - The W 's and Z 's were not massless like the photon
 - The W 's were connecting particles of different masses
- In 1967, Steven Weinberg and Abdus Salam independently proposed a solution to these problems
- The mass of the W and the Z , as well as all the quarks and leptons, had to come from a background field that pervades the universe, now called the Higgs field



Weak Interactions in the Quark Sector

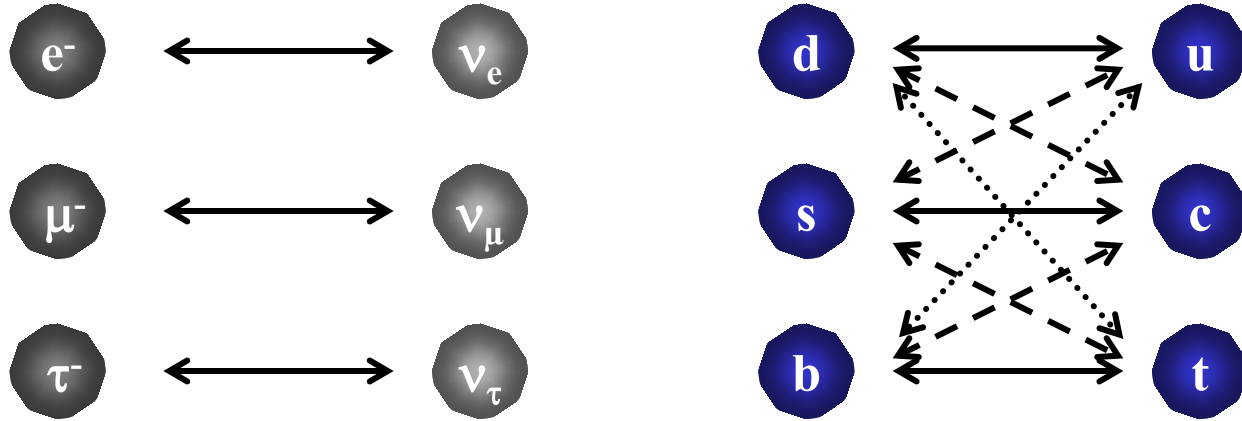
- In the leptons, we had two charged leptons and two neutrinos
- Emission or absorption of a W^\pm could convert them back and forth



- In the quarks, we had three quarks
- Emission or absorption of a W^\pm could convert them back and forth, but not equally
- The Z particle should also cause transitions that don't change the charge
 - This should cause $d \leftrightarrow s$ transitions
 - But they weren't observed
- In 1970, Glashow, Iliopoulos and Maiani found a solution
 - They had to assume there was a fourth quark, called charm
- In 1974, the charmed quark was discovered by Richter and Ting

The List Grows . . . But Not Forever

- In 1975, a new lepton was discovered, named the τ
 - It is just like the electron and muon, only heavier



- There was associated evidence for a new neutrino
 - Finally proven in 2001

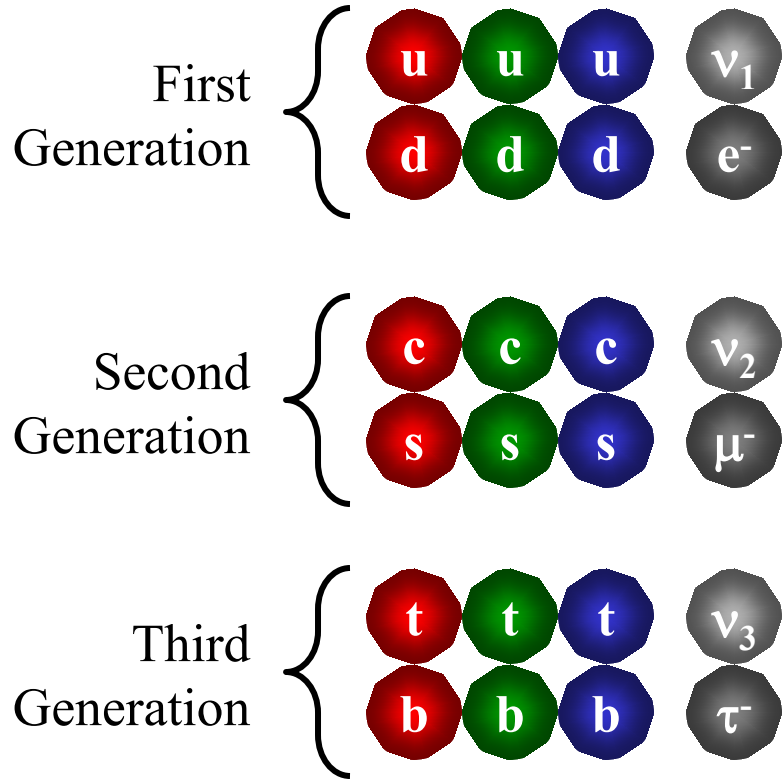
- Complicated arguments suggested it was likely there was another pair of quarks, too
 - Bottom quark (originally beauty) discovered in 1977
 - Top quark (originally truth) discovered in 1995
- Around 1989, measurements of the Z established that there were no new neutrinos
 - We now think this means we didn't miss anything

All Standard Model Particles

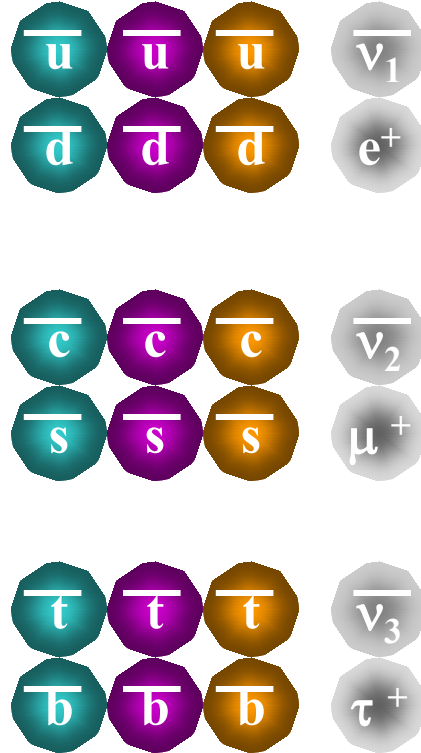
		<u>Particle</u>	<u>symbols</u>	<u>spin</u>	<u>charge</u>	<u>Mass (MeV/c²)</u>
Fermions (add anti-particles)	leptons	Electron	e^-	$\frac{1}{2}$	-1	0.511
		Neutrino 1	ν_1	$\frac{1}{2}$	0	0?
		Muon	μ^-	$\frac{1}{2}$	-1	105.7
		Neutrino 2	ν_2	$\frac{1}{2}$	0	9×10^{-9} ?
		Tau	τ^-	$\frac{1}{2}$	-1	1777
		Neutrino 3	ν_3	$\frac{1}{2}$	0	5×10^{-8} ?
	quarks	Up quark	u	$\frac{1}{2}$	$+\frac{2}{3}$	3
		Down quark	d	$\frac{1}{2}$	$-\frac{1}{3}$	5
		Charm quark	c	$\frac{1}{2}$	$+\frac{2}{3}$	1,300
		Strange quark	s	$\frac{1}{2}$	$-\frac{1}{3}$	120
		Top quark	t	$\frac{1}{2}$	$+\frac{2}{3}$	174,200
		Bottom quark	b	$\frac{1}{2}$	$-\frac{1}{3}$	4,300
	force carriers	Photon	γ	1	0	0
		Gluon	g	1	0	0
		W-boson	W^\pm	1	± 1	80,400
		Z-boson	Z^0	1	0	91,188
		Higgs	H	0	0	125,100

All Standard Model Particles

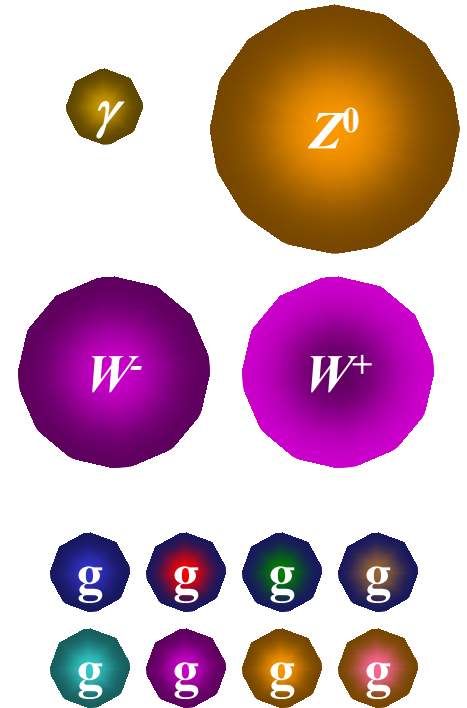
Spin $\frac{1}{2}$ Particles



Spin $\frac{1}{2}$ Anti-Particles



Spin 1 Force Carriers



Spin 0 Higgs



The Standard Model Lagrangian:

What part of

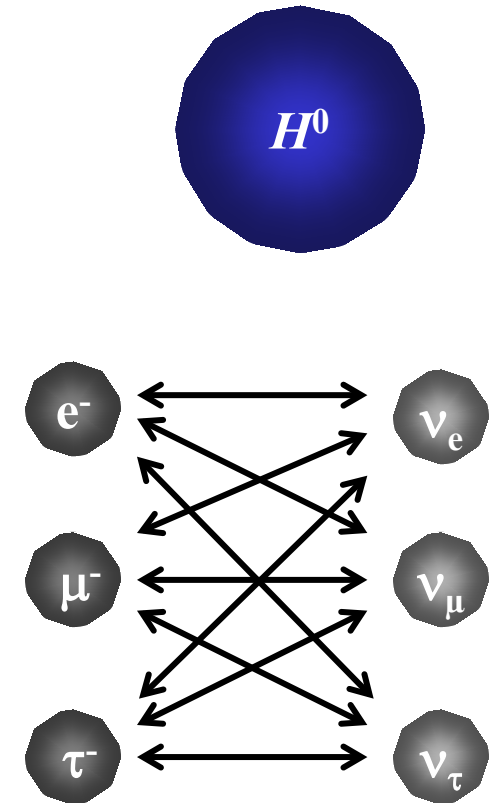
$$\begin{aligned}\mathcal{L} = & -\frac{1}{4}F_{\mu\nu}F^{\mu\nu} - \frac{1}{4}Z_{\mu\nu}Z^{\mu\nu} - \frac{1}{2}W_{\mu\nu}^+W^{-\mu\nu} - \frac{1}{4}G_{\mu\nu}^{(a)}G^{(a)\mu\nu} + \sum_f \bar{\psi}_f \left(i\gamma^\mu D_\mu - m_f \left(1 + \frac{H}{v} \right) \right) \psi_f \\ & + \left(m_W^2 W_\mu^+ W^{-\mu} + \frac{1}{2} m_Z^2 Z_\mu Z^\mu \right) \left(1 + \frac{H}{v} \right)^2 + \frac{1}{2} \partial_\mu H \partial^\mu H - \frac{1}{2} m_H^2 H^2 \left(1 + \frac{H}{2v} \right)^2\end{aligned}$$

don't you understand?

What's Missing?

- There are 18 numbers in this theory that must be put in by hand
 - 9 quark and lepton masses
 - 3 strengths of the forces (strong, weak, electromagnetic)
 - 4 describing the mixings in weak interactions
 - 2 describe the mass and strength of the Higgs field
- The Higgs particle: discovery announced July 4, 2012
- The three neutrinos are massless in the standard model
 - Experimental evidence for masses and mixing
- It is easy to fix this – too easy

Gravity



Outline of History of Universe

<u>Time</u>	<u>T or $k_B T$</u>	<u>Events</u>
10^{-43} s	10^{18} GeV	Planck Era; time becomes meaningless?
10^{-39} s	10^{16} GeV	Inflation begins; forces unified
10^{-35} s	10^{15} GeV	Inflation ends; reheating; forces separate; baryosynthesis (?)
10^{-13} s	1500 GeV	Supersymmetry breaking, LSP (dark matter)
10^{-11} s	160 GeV	Electroweak symmetry breaking
14 μ s	150 MeV	Quark Confinement
0.4 s	1.5 MeV	Neutrino Decoupling
1.5 s	0.7 MeV	Neutron/Proton freezeout
20 s	170 keV	Electron/Positron annihilation
200 s	80 keV	Nucleosynthesis
57 ky	0.76 eV	Matter-Radiation equality
370 ky	0.26 eV	Recombination
600 My	30 K	First Structure/First Stars
13.8 Gy	2.725 K	Today

Known Particle Physics

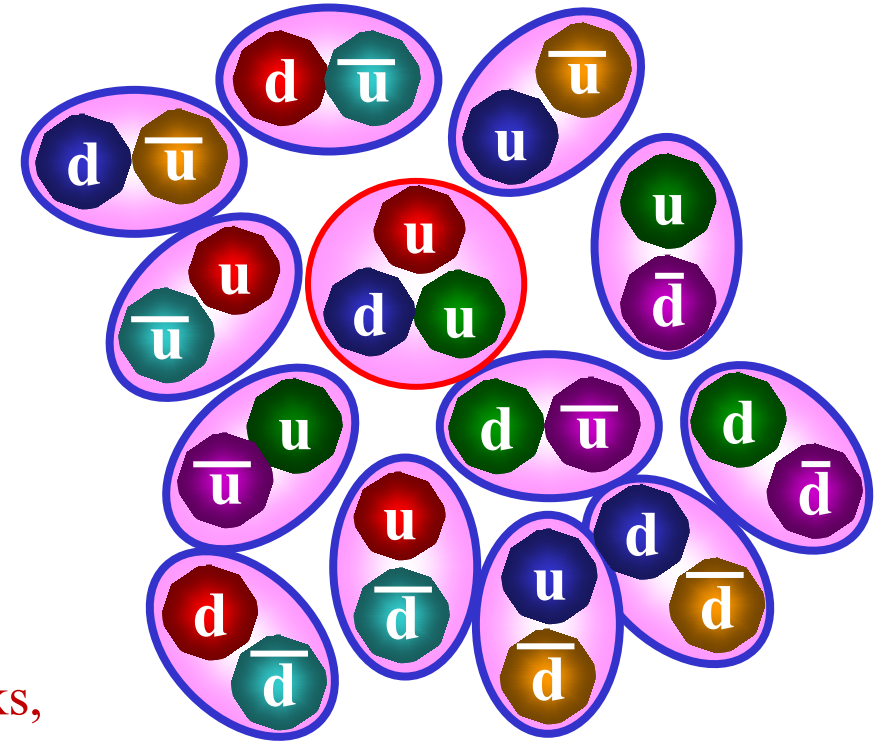
The Era of Particle Physics

More Particles Become Relevant

- At temperatures of 0.2 – 30 MeV, photons, neutrinos, and electrons (and their anti-particles) are effectively massless, and appear in high numbers
- Above 35 – 50 MeV (~ 0.3 ms) the muons and pions are relevant
- The pions are strongly interacting, and start to affect how all strongly interacting particles appear
- Theory says that the strong force becomes weaker at higher energy
- At 150 MeV quarks shift from being trapped in baryons and mesons to being free
 - Universe is filled with “quark soup”

Quark Confinement

- At low temperatures (< 40 MeV) just an occasional baryon
- At 45 MeV, pions start to appear
- At 150 MeV, the pions are so thick that they start overlapping with each other
- Quarks can jump from one pion to the next
- Strong force gets weaker at higher energy
 - Effectively confinement doesn't apply
- Quarks go from free to confined at 150 MeV
- At this temperature there are up, down, and strange quarks, and gluons
- In addition to the photons, neutrinos, electrons, and muons



Time	T or $k_B T$	Events
14 μ s	150 MeV	Quark Confinement

$$t = \frac{2.42 \text{ s}}{\sqrt{g_{\text{eff}}}} \left(\frac{\text{MeV}}{k_B T} \right)^2 = \frac{2.42 \text{ s}}{\sqrt{61.75}} \left(\frac{\text{MeV}}{150 \text{ MeV}} \right)^2$$

$$= 1.4 \times 10^{-5} \text{ s}$$

Particles in the Early Universe

$$g_{\text{eff}} = g_b + \frac{7}{8} g_f$$

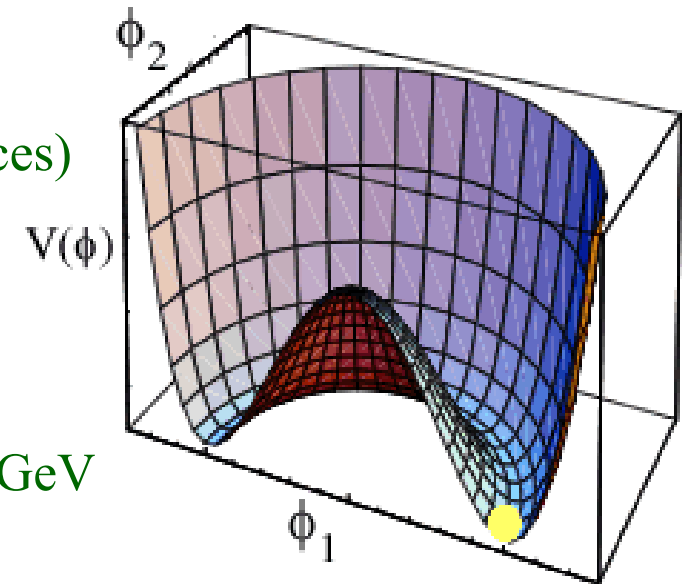
- For $k_B T < 150$ MeV, quarks are bound and it's complicated
- Above 150 MeV, all particles are in thermal equilibrium
- As temperature rises, particles get included when $3k_B T > mc^2$
- For 150 – 400 MeV, include $e, \mu, u, d, s, \nu_1, \nu_2, \nu_3, \gamma$ and g
- At 400 MeV, add c ; at 600 add τ
- At 1.5 GeV add b
- At 30 GeV, add W and Z
- At 40 GeV, add H ; at 60 add t
- Above 60 GeV, we have

$$g_{\text{eff}} = 28 + \frac{7}{8}(90) = 106.75$$

Particle	symbols	spin	g	mc^2 (MeV)
Electron	e	$\frac{1}{2}$	4	0.511
Neutrino 1	ν_1	$\frac{1}{2}$	2	~ 0
Up quark	uuu	$\frac{1}{2}$	12	~ 5
Down quark	ddd	$\frac{1}{2}$	12	~ 10
Muon	μ	$\frac{1}{2}$	4	105.7
Neutrino 2	ν_2	$\frac{1}{2}$	2	~ 0
Charm quark	ccc	$\frac{1}{2}$	12	1,270
Strange quark	sss	$\frac{1}{2}$	12	~ 100
Tau	τ	$\frac{1}{2}$	4	1,777
Neutrino 3	ν_3	$\frac{1}{2}$	2	~ 0
Top quark	ttt	$\frac{1}{2}$	12	173,000
Bottom quark	bbb	$\frac{1}{2}$	12	4,700
Photon	γ	1	2	0
Gluon	$gggggggg$	1	16	0
W-boson	W	1	6	80,400
Z-boson	Z	1	3	91,200
Higgs	H	0	1	125,100

Electroweak Phase Transition

- There are three forces that particle physicists understand:
 - Strong, electromagnetic, and weak
- Electromagnetic and weak forces affected by a field called the *Higgs field*
- The shape of the Higgs potential is interesting:
 - Sometimes called a Mexican Hat potential
- At low temperatures (us), one direction is easy to move (EM forces) and one is very hard (weak forces)
- At high temperatures, (early universe) you naturally move to the middle of the potential
- All directions are created equal
- Electroweak unification becomes apparent at perhaps $k_B T = 160 \text{ GeV}$



$$t = \frac{2.42 \text{ s}}{\sqrt{g_{\text{eff}}}} \left(\frac{\text{MeV}}{k_B T} \right)^2 = \frac{2.42 \text{ s}}{\sqrt{106.75}} \left(\frac{\text{MeV}}{160,000 \text{ MeV}} \right)^2 \approx 10^{-11} \text{ s}$$

Time

10^{-11} s

T or $k_B T$

160 GeV

Events

Electroweak Symmetry Breaking

Outline of History of Universe

<u>Time</u>	<u>T or $k_B T$</u>	<u>Events</u>
10^{-43} s	10^{18} GeV	Planck Era; time becomes meaningless?
10^{-39} s	10^{16} GeV	Inflation begins; forces unified
10^{-35} s	10^{15} GeV	Inflation ends; reheating; forces separate; baryosynthesis (?)
10^{-13} s	1500 GeV	Supersymmetry breaking, LSP (dark matter)
10^{-11} s	160 GeV	Electroweak symmetry breaking
14 μ s	150 MeV	Quark Confinement
0.4 s	1.5 MeV	Neutrino Decoupling
1.5 s	0.7 MeV	Neutron/Proton freezeout
20 s	170 keV	Electron/Positron annihilation
200 s	80 keV	Nucleosynthesis
57 ky	0.76 eV	Matter-Radiation equality
370 ky	0.26 eV	Recombination
600 My	30 K	First Structure/First Stars
13.8 Gy	2.725 K	Today

Unknown Particle Physics