

Physics 780 – General Relativity  
**Homework Set N**

34. Consider a light beam approaching a black hole with mass  $M$ . The light beam is moving in the plane  $\theta = \frac{1}{2}\pi$ . From class notes, we have

$$\frac{E^2}{J^2} - \frac{1}{r^2} + \frac{2GM}{r^3} = \frac{1}{r^4} \left( \frac{dr}{d\phi} \right)^2.$$

- (a) Find the radius at which the left side of this equation is extremized (minimum or maximum). Is it a minimum or maximum? A photon at this radius can circle endlessly, ( $dr/d\phi = 0$ ) if  $E/J$  has the right value. Will this be a stable or unstable orbit?
  - (b) What will be the value of  $E/J$  for this orbit? Keeping in mind that we showed in class that  $J/E = b$  is the impact parameter, you should be able to find the impact parameter  $b_C$  which will end up converging to a circular orbit.
  - (c) Qualitatively, what will happen to a photon that starts at a larger impact parameter, so we have  $b > b_C$ ? That is to say, will there be any radius where  $dr/d\phi = 0$ ? What if  $b < b_C$ ?
  - (d) Find the cross-section for absorption of photons by a black hole; that is, the area of incoming photons that are absorbed by the black hole.
35. Can you make a black hole in two spatial dimensions? We will work in polar coordinates  $(t, r, \phi)$ , and assume the stress-energy tensor is zero away from the origin.
- (a) First, what is the flat spacetime metric in polar coordinates? If you don't know, write it in Cartesian coordinates and rewrite it using  $x = r \cos \phi$  and  $y = r \sin \phi$ .
  - (b) Assume the black hole is rotationally invariant and time invariant, so we can choose coordinates such that  $\partial_t$  and  $\partial_\phi$  are Killing vectors. What does this tell us about the metric components, *i.e.*, what coordinates can they depend on?
  - (c) Assume the metric is invariant under reflection so that  $\phi \rightarrow -\phi$ . Argue that the metric now takes the form  $ds^2 = -f dt^2 + 2j dr dt + h dr^2 + b d\phi^2$ .
  - (d) Change time to a new coordinate  $t \rightarrow t' = t - \int (j/f) dr$ . Show that this eliminates  $j$ .  
 Once you have done so, rename any functions and variables so the metric now takes the form in part (c), but with  $j = 0$ .
  - (e) Explain why you can change variables  $r$  such that the metric is now  $ds^2 = -f dt^2 + h dr^2 + r^2 d\phi^2$ .
  - (f) Find all the components of the Ricci tensor and/or Einstein tensor. I recommend you use **gcalc** or a similar method to save your sanity.
  - (g) Since we have no source away from the origin,  $R_{\mu\nu} = G_{\mu\nu} = 0$ . Based on this, show that  $f$  and  $h$  must both be constants (I used the Einstein tensor). Argue that by rescaling your time coordinate, one of these functions can be set equal to one.
  - (h) Show that by rescaling the radial and angular coordinate  $r' = r\sqrt{h}$  and  $\phi' = \phi/\sqrt{h}$ , we can make the metric look just like the one in part (a). You might think this metric is identical to flat spacetime, but it is not. Why not? Hint: what is the range of  $\phi'$ ?