

**PHY 113 A General Physics I
9:50 AM MWF Olin 101**

Plan for Lecture 25:

Review: Chapters 10-13, 15

1. Advice on how to prepare for exam
2. Review of rotational motion, angular momentum, static equilibrium, simple harmonic motion, universal gravitational force law

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|----|------------|------------------------------------|-----------|---------------------|------------|
| 14 | 10/03/2012 | Momentum and collisions | 9.5-9.9 | 9.29-9.37 | 10/05/2012 |
| | 10/05/2012 | Review | 6-9 | | |
| | 10/08/2012 | Exam | 6-9 | | |
| 15 | 10/10/2012 | Rotational motion | 10.1-10.5 | 10.6, 10.13, 10.25 | 10/12/2012 |
| 16 | 10/12/2012 | Torque | 10.6-10.9 | 10.37, 10.55 | 10/15/2012 |
| 17 | 10/15/2012 | Angular momentum | 11.1-11.5 | 11.11, 11.34 | 10/17/2012 |
| 18 | 10/17/2012 | Equilibrium | 12.1-12.4 | 12.11, 12.39 | 10/22/2012 |
| | 10/19/2012 | Fall Break | | | |
| 19 | 10/22/2012 | Simple harmonic motion | 15.1-15.3 | 15.4, 15.20 | 10/24/2012 |
| 20 | 10/24/2012 | Resonance | 15.4-15.7 | 15.43, 15.43, 15.52 | 10/26/2012 |
| 21 | 10/26/2012 | Gravitational force | 13.1-13.3 | 13.6, 13.10, 13.13 | 10/29/2012 |
| 22 | 10/29/2012 | Kepler's laws and satellite motion | 13.4-13.6 | 13.28, 13.34 | 10/31/2012 |
| | 10/31/2012 | Review | 10-13.15 | | |
| | 11/02/2012 | Exam | 10-13.15 | | |
| 23 | 11/05/2012 | Fluid mechanics | 14.1-14.4 | | 11/07/2012 |
| 24 | 11/07/2012 | Fluid mechanics | 14.5-14.7 | | 11/09/2012 |
| 25 | 11/09/2012 | Temperature | 19.1-19.5 | | 11/12/2012 |

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Format of Friday's exam

What to bring:

1. Clear, calm head
 2. Equation sheet (turn in with exam)
 3. Scientific calculator
 4. Pencil or pen
- (Note: laptops, cellphones, and other electronic equipment must be off or in sleep mode.)

Timing:

May begin as early as 8 AM; must end \leq 9:50 AM

Probable exam format

- 4-5 problems similar to homework and class examples; focus on Chapters 10-13 & 15 of your text.
- Full credit awarded on basis of analysis steps as well as final answer

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| Examples of what to include on equation sheet | |
|--|--|
| Given information on exam | Suitable for equation sheet |
| Universal or common constants (such as g, G, M _E , M _S , R _E ...) | Basic equations from material from earlier Chapters: Newton's laws, energy, momentum, center of mass |
| Particular constants (such as k, m, l, ...) | Simple derivative and integral relationships, including trigonometric functions |
| Unit conversion factors such as Hz and rad/s | Definition of moment of inertia, torque, angular momentum, rotational kinetic energy |
| | Newton's law for rotational motion; combination of rotational and center of mass motion |
| | Equations describing simple harmonic motion and driven harmonic motion |
| | Newton's universal gravitation force law and corresponding gravitational potential energy |
| | Gravitational stable circular orbits |

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Possible extra review session on Thursday:

iclicker question:
Which of the following possible times would work with your schedules (vote for one)?

- A. 2 PM
- B. 3 PM
- C. 4 PM
- D. Prefer to meet individually or in small groups in my office (Olin 300).

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Rotations: Angular variables

angular "displacement" $\rightarrow \theta(t)$
 angular "velocity" $\rightarrow \omega(t) = \frac{d\theta}{dt}$
 angular "acceleration" $\rightarrow \alpha(t) = \frac{d\omega}{dt}$

"natural" angular unit = radian

Relation to linear variables:

$$s_\theta = r(\theta_f - \theta_i)$$

$$v_\theta = r\omega$$

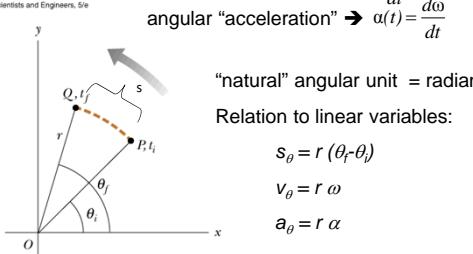
$$a_\theta = r\alpha$$


Figure 10.2: A diagram illustrating angular variables. A point P at position vector r from the origin O rotates clockwise through an angle theta_i to a new position P at time t_f. The arc length traveled by P is s. The angle theta is measured between the initial radius r_i and the final radius r_f.

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Object rotating with constant angular velocity ($\alpha = 0$)

Kinetic energy associated with rotation:

$$K = \sum_i \frac{1}{2} m_i v_i^2 = \sum_i \frac{1}{2} m_i r_i^2 \omega^2 \equiv \frac{1}{2} I \omega^2;$$

where : $I \equiv \sum_i m_i r_i^2$ "moment of inertia"

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Moment of inertia: $I \equiv \sum m_i r_i^2$

$I = 2Ma^2$ $I = 2Ma^2 + 2mb^2$

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Total kinetic energy of rolling object :

$$K_{total} = K_{rolling} + K_{CM}$$

$$= \frac{1}{2} I \omega^2 + \frac{1}{2} M v_{CM}^2$$

Note that :

$$\omega = \frac{d\theta}{dt}$$

$$\frac{ds}{dt} = R \frac{d\theta}{dt} = R\omega = v_{CM}$$

$$= \frac{1}{2} \frac{I}{R^2} (R\omega)^2 + \frac{1}{2} M v_{CM}^2$$

$$= \frac{1}{2} \left(\frac{I}{R^2} + M \right) v_{CM}^2$$

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How to make objects rotate.

Define torque:

$$\tau = \mathbf{r} \times \mathbf{F}$$

$$\tau = rF \sin \theta$$

$\mathbf{F} = m\mathbf{a}$

$\mathbf{r} \times \mathbf{F} \equiv \tau = \mathbf{r} \times m\mathbf{a} = I\mathbf{a}$

Note: We will define and use the “vector cross product” next time. For now, we focus on the fact that the direction of the torque determines the direction of rotation.

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Vector cross product; right hand rule

The direction of \vec{C} is perpendicular to the plane formed by \vec{A} and \vec{B} ; choose which perpendicular direction using the right-hand rule shown by the hand.

$\vec{C} = \vec{A} \times \vec{B}$

$|\vec{C}| = |\vec{A}| |\vec{B}| \sin \theta$

$\hat{i} \times \hat{i} = \hat{j} \times \hat{j} = \hat{k} \times \hat{k} = 0$

$\hat{i} \times \hat{j} = -\hat{j} \times \hat{i} = \hat{k}$

$\hat{j} \times \hat{k} = -\hat{k} \times \hat{j} = \hat{i}$

$\hat{k} \times \hat{i} = -\hat{i} \times \hat{k} = \hat{j}$

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Note that rolling motion is caused by the torque of friction:

Newton's law for torque:

$$\tau_{total} = I \frac{d\omega}{dt} = I\alpha$$

Diagram of a cylinder of radius R and mass M, with a force F applied at the center of the top face, causing it to roll to the right. Friction force f_s acts at the bottom center.

$$F - f_s = Ma_{CM}$$

$$f_s R = I\alpha = Ia_{CM} / R \Rightarrow a_{CM} = \frac{f_s R^2}{I}$$

$$f_s = F \left(\frac{1}{1 + (MR^2)/I} \right)$$

For a solid cylinder, $I = \frac{1}{2}MR^2 \Rightarrow f_s = \frac{1}{3}F$

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Torque and Newton's second law:

$$\mathbf{F} = m\mathbf{a}$$

$$\mathbf{r} \times \mathbf{F} \equiv \tau = \mathbf{r} \times m\mathbf{a} = \mathbf{r} \times m \frac{d\mathbf{v}}{dt} = \mathbf{r} \times \frac{d(m\mathbf{v})}{dt} = \frac{d}{dt}(\mathbf{r} \times \mathbf{p})$$

$$\mathbf{r} \times \mathbf{F} = \tau = \frac{d}{dt}(\mathbf{r} \times \mathbf{p})$$

Define : $\mathbf{L} \equiv \mathbf{r} \times \mathbf{p}$

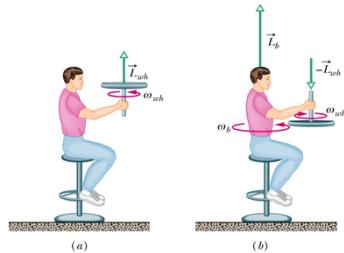
Note that : if $\tau = 0$ $\frac{d\mathbf{L}}{dt} = 0$

Then : $\mathbf{L} = (\text{constant})$

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Example of conservation of angular momentum:

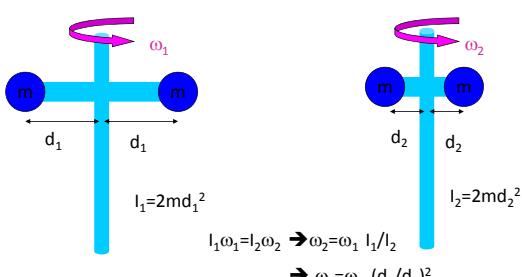
$$\begin{array}{c} \text{Initial} \\ \uparrow \vec{L}_{wh} \\ \downarrow \end{array} = \begin{array}{c} \text{Final} \\ \uparrow \vec{L}_b \\ \downarrow -\vec{L}_{wh} \end{array}$$

$$\begin{aligned} L_{bf} + L_{wheelf} &= L_{bi} + L_{wheeli} \\ L_{bf} - L_{wheel} &= 0 + L_{wheel} \\ L_{bf} &= 2L_{wheel} \end{aligned}$$

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Another example of conservation of angular momentum

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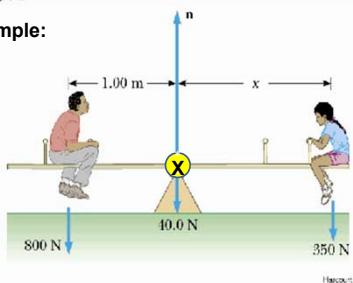
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Conditions for stable equilibriumBalance of force: $\sum_i \mathbf{F}_i = 0$ Balance of torque: $\sum_i \tau_i = 0$

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Example:

Forces : $n - M_D g - m_c g - m_p g = 0$

Torques : $M_D g (1m) - m_c g x = 0$

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Simple harmonic motion:

Newton's law for mass - spring system : $ma = m \frac{d^2x}{dt^2} = F = -kx$

$$\frac{d^2x}{dt^2} = -\frac{k}{m}x$$

Guess that solution for $x(t)$ has the form :

$x(t) = A \cos(\omega t + \phi)$ where A and ϕ and ω are unknown constants

Condition that guess satisfies the equation :

$$\frac{d^2[A \cos(\omega t + \phi)]}{dt^2} = -\omega^2 [A \cos(\omega t + \phi)] = -\frac{k}{m} [A \cos(\omega t + \phi)]$$

$$\Rightarrow \omega^2 = \frac{k}{m}$$
 (determines ω -- "natural frequency")

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Summary --

Simple harmonic motion:

$$F = -kx = m \frac{d^2x}{dt^2}$$

$$\frac{d^2x}{dt^2} = -\frac{k}{m}x$$

$$x(t) = A \cos(\omega t + \varphi); \quad \omega = \sqrt{\frac{k}{m}}$$

Note that:

$$v(t) = \frac{dx}{dt} = -A\omega \sin(\omega t + \varphi)$$

$$a(t) = \frac{dv}{dt} = -A\omega^2 \cos(\omega t + \varphi)$$

Conveniently evaluated in radians

Constants

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Energy associated with simple harmonic motion

Energy:

$$E = \frac{1}{2}mv^2 + \frac{1}{2}kx^2$$

$$E = \frac{1}{2}m\omega^2 A^2 (\sin(\omega t + \varphi))^2 + \frac{1}{2}kA^2 (\cos(\omega t + \varphi))^2$$

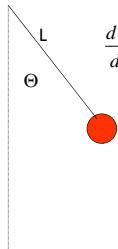
$$\text{But } \omega^2 = \frac{k}{m}$$

$$\Rightarrow E = \frac{1}{2}kA^2 [(\sin(\omega t + \varphi))^2 + (\cos(\omega t + \varphi))^2] = \frac{1}{2}kA^2$$

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Simple harmonic motion for a pendulum:

$$\tau = mgL \sin \Theta = -I\alpha = -I \frac{d^2\Theta}{dt^2}$$

$$\frac{d^2\Theta}{dt^2} = -\frac{mgL}{I} \sin \Theta = -\frac{g}{L} \sin \Theta \quad (\text{since } I = mL^2)$$

Approximation for small Θ :

$$\sin \Theta \approx \Theta$$

$$\Rightarrow \frac{d^2\Theta}{dt^2} = -\frac{g}{L} \Theta$$

Solution :

$$\Theta(t) = A \cos(\omega t + \varphi); \quad \omega = \sqrt{\frac{g}{L}}$$

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The notion of resonance:

$$\text{Suppose } F = -kx + F_0 \sin(\Omega t)$$

According to Newton:

$$-kx + F_0 \sin(\Omega t) = m \frac{d^2x}{dt^2}$$

Differential equation ("inhomogeneous"):

$$\frac{d^2x}{dt^2} = -\frac{k}{m}x + \frac{F_0}{m} \sin(\Omega t)$$

Solution:

$$x(t) = \frac{F_0/m}{k/m - \Omega^2} \sin(\Omega t) \equiv \frac{F_0/m}{\omega^2 - \Omega^2} \sin(\Omega t)$$

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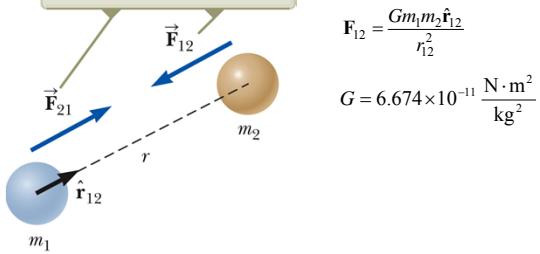
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Universal law of gravitation

→ Newton (with help from Galileo, Kepler, etc.) 1687

Consistent with Newton's third law, $\vec{F}_{21} = -\vec{F}_{12}$.



$$\vec{F}_{12} = \frac{G m_1 m_2 \hat{r}_{12}}{r_{12}^2}$$

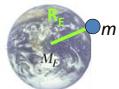
$$G = 6.674 \times 10^{-11} \frac{\text{N} \cdot \text{m}^2}{\text{kg}^2}$$

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Review: Gravitational force of the Earth



$$F = \frac{GM_E m}{R_E^2}$$

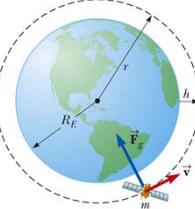
$$\Rightarrow g = \frac{GM_E}{R_E^2} = \frac{6.67 \times 10^{-11} \cdot 5.98 \times 10^{24}}{(6.37 \times 10^6)^2} \text{ m/s}^2 = 9.8 \text{ m/s}^2$$

Note: Earth's gravity acts as a point mass located at the Earth's center.

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Example: Satellite in circular Earth orbit

$$\text{For } m \ll M_E \quad r \equiv R_E + h$$

$$m \frac{v^2}{r} = m \frac{(2\pi r/T)^2}{r} = m \left(\frac{2\pi}{T} \right)^2 r = \frac{GM_E m}{r^2}$$

$$T = 2\pi \sqrt{\frac{(R_E + h)^3}{GM_E}}$$

If $h = 35.83 \times 10^6 \text{ m}$
 $T = 8.53 \times 10^4 \text{ s} = 1 \text{ day}$
 (geosynchronous)

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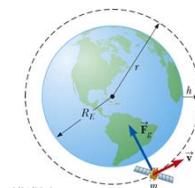
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Gravitational potential energy

$$U_{\text{gravity}}(r) = - \int_{r_{\text{ref}}}^r \mathbf{F} \cdot d\mathbf{r} \quad \mathbf{F} = -\frac{Gm_1 m_2 \hat{\mathbf{r}}}{r^2}$$

$$U_{\text{gravity}}(r) = - \int_{\infty}^r \frac{-Gm_1 m_2}{r'^2} dr' = -\frac{Gm_1 m_2}{r}$$

Example:

$$U_{\text{gravity}}(r = R_E + h) = -\frac{GM_E m}{R_E + h}$$

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