

PHY 711 Classical Mechanics and Mathematical Methods
10-10:50 AM MWF Olin 103

Plan for Lecture 4:

Chapter 2 – Physics described in an accelerated coordinate frame

1. Linear acceleration
2. Angular acceleration
3. Foucault pendulum

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PHY 711 Classical Mechanics and Mathematical Methods

MWF 10 AM-10:50 PM [OPL 103] <http://www.wfu.edu/~natalie/f15phy711/>

Instructor: Natalie Holzwarth Phone:758-5510 Office:300 OPL e-mail:natalie@wfu.edu

Course schedule

(Preliminary schedule -- subject to frequent adjustment.)

Date	F&W Reading	Topic	Assignment
1 Wed, 8/26/2015	Chap. 1	Review of basic principles	#1
2 Fri, 8/28/2015	Chap. 1	Scattering theory	#2
3 Mon, 8/31/2015	Chap. 1	Scattering theory continued	#3
4 Wed, 9/02/2015	Chap. 2	Accelerated coordinate systems	#4
5 Fri, 9/04/2015	Chap. 3	Calculus of variations	
6 Mon, 9/07/2015	Chap. 3	Calculus of variations	
7 Wed, 9/09/2015	Chap. 3	Hamilton's principle	

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Department of Physics

News



Congratulations to Dr. Greg Smith, recent Ph.D. Recipient



Congratulations to Dr. Jie Liu, recent Ph.D. Recipient



Congratulations to Dr. Wei Li, recent Ph.D. Recipient

Events

Wed, Sept. 2, 2015
 Laboratory tours for students – Part I
 Meet in Olin Lobby at 4 PM
 Refreshments at 3:30 PM
 Olin Lobby

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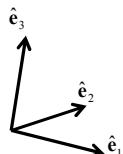
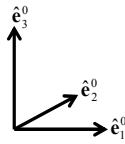
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Physical laws as described in non-inertial coordinate systems

- Newton's laws are formulated in an inertial frame of reference $\{\hat{e}_i^0\}$

- For some problems, it is convenient to transform the equations into a non-inertial coordinate system

$$\{\hat{e}_i(t)\}$$



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Comparison of analysis in “inertial frame” versus “non-inertial frame”

Denote by \hat{e}_i^0 a fixed coordinate system

Denote by \hat{e}_i a moving coordinate system

$$\mathbf{V} = \sum_{i=1}^3 V_i^0 \hat{e}_i^0 = \sum_{i=1}^3 V_i \hat{e}_i$$

$$\left(\frac{d\mathbf{V}}{dt} \right)_{inertial} = \sum_{i=1}^3 \frac{dV_i^0}{dt} \hat{e}_i^0 = \sum_{i=1}^3 \frac{dV_i}{dt} \hat{e}_i + \sum_{i=1}^3 V_i \frac{d\hat{e}_i}{dt}$$

$$\text{Define : } \left(\frac{d\mathbf{V}}{dt} \right)_{body} \equiv \sum_{i=1}^3 \frac{dV_i}{dt} \hat{e}_i$$

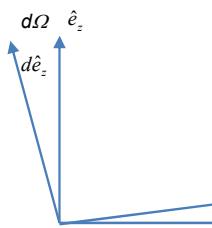
$$\Rightarrow \left(\frac{d\mathbf{V}}{dt} \right)_{inertial} = \left(\frac{d\mathbf{V}}{dt} \right)_{body} + \sum_{i=1}^3 V_i \frac{d\hat{e}_i}{dt}$$

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Properties of the frame motion (rotation only):



$$\begin{aligned} d\hat{e}_y &= d\Omega \hat{e}_z \\ d\hat{e}_z &= -d\Omega \hat{e}_y \\ \Rightarrow d\hat{e} &= d\Omega \times \hat{e} \\ \frac{d\hat{e}}{dt} &= \frac{d\Omega}{dt} \times \hat{e} \\ \frac{d\hat{e}}{dt} &= \boldsymbol{\omega} \times \hat{e} \end{aligned}$$

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Properties of the frame motion (rotation only):

$$d\Omega \hat{e}_z$$

$$d\hat{\mathbf{e}} = d\Omega \times \hat{\mathbf{e}} \quad \frac{d\hat{\mathbf{e}}}{dt} = \frac{d\Omega}{dt} \times \hat{\mathbf{e}} \quad \frac{d\hat{\mathbf{e}}}{dt} = \boldsymbol{\omega} \times \hat{\mathbf{e}}$$

Rotation about x -axis:

$$\begin{pmatrix} e_y \\ e_z \end{pmatrix} = \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix} \begin{pmatrix} e_y \\ e_z \end{pmatrix} \quad \begin{pmatrix} e_y + de_y \\ e_z + de_z \end{pmatrix} = \begin{pmatrix} \cos(d\Omega) & -\sin(d\Omega) \\ \sin(d\Omega) & \cos(d\Omega) \end{pmatrix} \begin{pmatrix} e_y \\ e_z \end{pmatrix}$$

$$\begin{pmatrix} de_y \\ de_z \end{pmatrix} = \begin{pmatrix} \cos(d\Omega) - 1 & -\sin(d\Omega) \\ \sin(d\Omega) & \cos(d\Omega) - 1 \end{pmatrix} \begin{pmatrix} e_y \\ e_z \end{pmatrix} \approx \begin{pmatrix} 0 & -d\Omega \\ d\Omega & 0 \end{pmatrix} \begin{pmatrix} e_y \\ e_z \end{pmatrix}$$

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Properties of the frame motion (rotation only):

$$d\Omega \hat{e}_z$$

$$d\hat{\mathbf{e}} = d\Omega \times \hat{\mathbf{e}} \quad \frac{d\hat{\mathbf{e}}}{dt} = \frac{d\Omega}{dt} \times \hat{\mathbf{e}} \quad \frac{d\hat{\mathbf{e}}}{dt} = \boldsymbol{\omega} \times \hat{\mathbf{e}}$$

Rotation about x -axis:

$$\begin{pmatrix} de_y \\ de_z \end{pmatrix} \approx \begin{pmatrix} 0 & -d\Omega \\ d\Omega & 0 \end{pmatrix} \begin{pmatrix} e_y \\ e_z \end{pmatrix} = d\Omega \hat{\mathbf{e}}_y - d\Omega \hat{\mathbf{e}}_z = d\Omega \hat{\mathbf{x}} \times \hat{\mathbf{e}}$$

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Properties of the frame motion (rotation only) -- continued

$$\left(\frac{d\mathbf{V}}{dt} \right)_{inertial} = \left(\frac{d\mathbf{V}}{dt} \right)_{body} + \sum_{i=1}^3 V_i \frac{d\hat{e}_i}{dt}$$

$$\left(\frac{d\mathbf{V}}{dt} \right)_{inertial} = \left(\frac{d\mathbf{V}}{dt} \right)_{body} + \boldsymbol{\omega} \times \mathbf{V}$$

Effects on acceleration (rotation only):

$$\left(\frac{d}{dt} \frac{d\mathbf{V}}{dt} \right)_{inertial} = \left\{ \left(\frac{d}{dt} \right)_{body} + \boldsymbol{\omega} \times \right\} \left(\frac{d\mathbf{V}}{dt} \right)_{body} + \boldsymbol{\omega} \times \mathbf{V}$$

$$\left(\frac{d^2\mathbf{V}}{dt^2} \right)_{inertial} = \left(\frac{d^2\mathbf{V}}{dt^2} \right)_{body} + 2\boldsymbol{\omega} \times \left(\frac{d\mathbf{V}}{dt} \right)_{body} + \frac{d\boldsymbol{\omega}}{dt} \times \mathbf{V} + \boldsymbol{\omega} \times \boldsymbol{\omega} \times \mathbf{V}$$

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Application of Newton's laws in a coordinate system which has an angular velocity ω and linear acceleration $\left(\frac{d^2\mathbf{a}}{dt^2}\right)_{inertial}$

Newton's laws; Let \mathbf{r} denote the position of particle of mass m :

$$\begin{aligned} m\left(\frac{d^2\mathbf{r}}{dt^2}\right)_{inertial} &= \mathbf{F}_{ext} \\ m\left(\frac{d^2\mathbf{r}}{dt^2}\right)_{body} &= \mathbf{F}_{ext} - m\left(\frac{d^2\mathbf{a}}{dt^2}\right)_{inertial} - 2m\omega \times \left(\frac{d\mathbf{r}}{dt}\right)_{body} - m \frac{d\omega}{dt} \times \mathbf{r} - m\omega \times \omega \times \mathbf{r} \end{aligned}$$

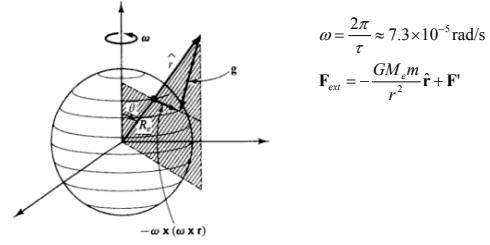
Coriolis force Centrifugal force

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Motion on the surface of the Earth:



$$\begin{aligned} \omega &= \frac{2\pi}{T} \approx 7.3 \times 10^{-5} \text{ rad/s} \\ \mathbf{F}_{ext} &= -\frac{GM_e m}{r^2} \hat{\mathbf{r}} + \mathbf{F}' \end{aligned}$$

Main contributions :

$$m\left(\frac{d^2\mathbf{r}}{dt^2}\right)_{earth} = -\frac{GM_e m}{r^2} \hat{\mathbf{r}} + \mathbf{F}' - 2m\omega \times \left(\frac{d\mathbf{r}}{dt}\right)_{earth} - m \frac{d\omega}{dt} \times \mathbf{r} - m\omega \times \omega \times \mathbf{r}$$

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Non-inertial effects on effective gravitational “constant”

$$\begin{aligned} m\left(\frac{d^2\mathbf{r}}{dt^2}\right)_{earth} &= -\frac{GM_e m}{r^2} \hat{\mathbf{r}} + \mathbf{F}' - 2m\omega \times \left(\frac{d\mathbf{r}}{dt}\right)_{earth} - m \frac{d\omega}{dt} \times \mathbf{r} - m\omega \times \omega \times \mathbf{r} \\ \text{For } \left(\frac{d\mathbf{r}}{dt}\right)_{earth} &= 0 \quad \text{and} \quad \left(\frac{d^2\mathbf{r}}{dt^2}\right)_{earth} = 0, \\ 0 &= -\frac{GM_e m}{r^2} \hat{\mathbf{r}} + \mathbf{F}' - m\omega \times \omega \times \mathbf{r} \\ \mathbf{F}' &= -mg \\ \Rightarrow \mathbf{g} &= -\frac{GM_e}{r^2} \hat{\mathbf{r}} - \omega \times \omega \times \mathbf{r} \Big|_{r=R_e} \\ &= \left(-\frac{GM_e}{R_e^2} + \omega^2 R_e \sin^2 \theta \right) \hat{\mathbf{r}} + \sin \theta \cos \theta \omega^2 R_e \hat{\theta} \end{aligned}$$

0.03 m/s²
9.80 m/s²

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Foucault pendulum http://www.si.edu/Encyclopedia_SI/nmah/pendulum.htm

The Foucault pendulum was displayed for many years in the Smithsonian's National Museum of American History. It is named for the French physicist Jean Foucault who first used it in 1851 to demonstrate the rotation of the earth.

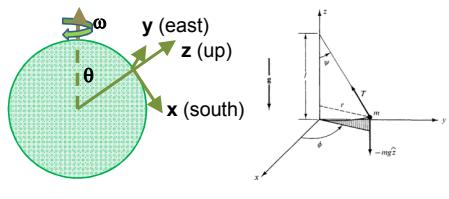
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Equation of motion on Earth's surface

$$m \left(\frac{d^2 \mathbf{r}}{dt^2} \right)_{\text{earth}} = -\frac{GM_e m}{r^2} \hat{\mathbf{r}} + \mathbf{F}' - 2m\boldsymbol{\omega} \times \left(\frac{d\mathbf{r}}{dt} \right)_{\text{earth}} - m \frac{d\boldsymbol{\omega}}{dt} \times \mathbf{r} - m\boldsymbol{\omega} \times \boldsymbol{\omega} \times \mathbf{r}$$



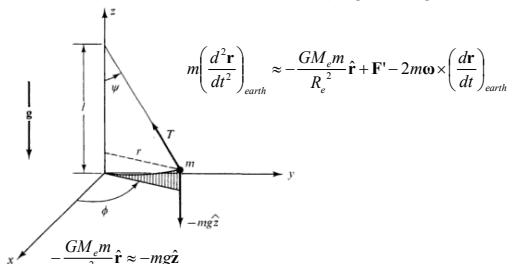
$$\boldsymbol{\omega} \approx -\omega \sin \theta \hat{x} + \omega \cos \theta \hat{z}$$

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Foucault pendulum continued – keeping leading terms:



$$\mathbf{F}' \approx -T \sin \psi \cos \phi \hat{x} - T \sin \psi \sin \phi \hat{y} + T \cos \psi \hat{z}$$

$$\boldsymbol{\omega} \approx -\omega \sin \theta \hat{x} + \omega \cos \theta \hat{z}$$

$$\boldsymbol{\omega} \times \left(\frac{d\mathbf{r}}{dt} \right)_{\text{earth}} \approx \boldsymbol{\omega} (-\dot{y} \cos \theta \hat{x} + (\dot{x} \cos \theta + \dot{z} \sin \theta) \hat{y} - \dot{y} \sin \theta \hat{z})$$

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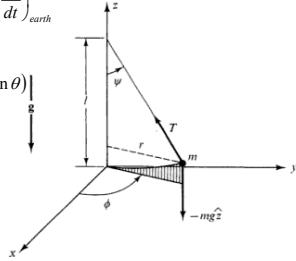
Foucault pendulum continued – keeping leading terms:

$$m \left(\frac{d^2 \mathbf{r}}{dt^2} \right)_{\text{earth}} \approx -\frac{GM_e m}{R_e^2} \hat{\mathbf{r}} + \mathbf{F}' - 2m\omega \times \left(\frac{d\mathbf{r}}{dt} \right)_{\text{earth}}$$

$$m\ddot{x} \approx -T \sin \psi \cos \phi + 2m\omega \dot{y} \cos \theta$$

$$m\ddot{y} \approx -T \sin \psi \sin \phi - 2m\omega (\dot{x} \cos \theta + \dot{z} \sin \theta)$$

$$m\ddot{z} \approx T \cos \psi - mg + 2m\omega \dot{y} \sin \theta$$



Further approximation :

$$\psi \ll 1; \quad \dot{z} \approx 0; \quad T \approx mg$$

$$m\ddot{x} \approx -mg \sin \psi \cos \phi + 2m\omega \dot{y} \cos \theta$$

$$m\ddot{y} \approx -mg \sin \psi \sin \phi - 2m\omega \dot{x} \cos \theta$$

Also note that :

$$x \approx \ell \sin \psi \cos \phi$$

$$y \approx \ell \sin \psi \sin \phi$$

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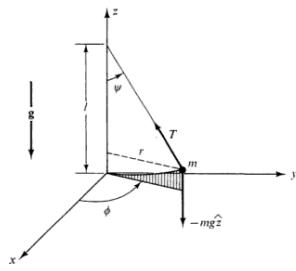
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Foucault pendulum continued – coupled equations:

$$\ddot{x} \approx -\frac{g}{\ell} x + 2\omega \cos \theta \dot{y}$$

$$\ddot{y} \approx -\frac{g}{\ell} y - 2\omega \cos \theta \dot{x}$$



Try to find a solution of the form :

$$x(t) = X e^{-iqt}, \quad y(t) = Y e^{-iqt}$$

Denote $\omega_{\perp} \equiv \omega \cos \theta$

$$\begin{pmatrix} -q^2 + \frac{g}{\ell} & i2\omega_{\perp} q \\ -i2\omega_{\perp} q & -q^2 + \frac{g}{\ell} \end{pmatrix} \begin{pmatrix} X \\ Y \end{pmatrix} = 0$$

Non-trivial solutions :

$$q_{\pm} = \alpha \pm \beta \equiv \omega_{\perp} \pm \sqrt{\omega_{\perp}^2 + \frac{g}{\ell}}$$

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Foucault pendulum continued – coupled equations:

Solution continued :

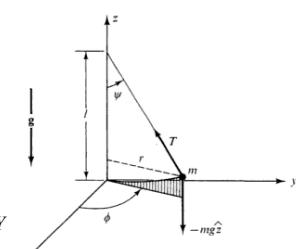
$$x(t) = X e^{-iqt}, \quad y(t) = Y e^{-iqt}$$

$$\begin{pmatrix} -q^2 + \frac{g}{\ell} & i2\omega_{\perp} q \\ -i2\omega_{\perp} q & -q^2 + \frac{g}{\ell} \end{pmatrix} \begin{pmatrix} X \\ Y \end{pmatrix} = 0$$

Non-trivial solutions :

$$q_{\pm} = \alpha \pm \beta \equiv \omega_{\perp} \pm \sqrt{\omega_{\perp}^2 + \frac{g}{\ell}}$$

Amplitude relationship : $X = iY$



General solution with complex amplitudes C and D :

$$x(t) = \operatorname{Re} \{ C e^{-i(\alpha+\beta)t} + iD e^{-i(\alpha-\beta)t} \}$$

$$y(t) = \operatorname{Re} \{ C e^{-i(\alpha+\beta)t} + D e^{-i(\alpha-\beta)t} \}$$

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General solution with complex amplitudes C and D :

$$x(t) = \operatorname{Re} \{ iCe^{-i(\alpha+\beta)t} + iDe^{-i(\alpha-\beta)t} \}$$

$$y(t) = \operatorname{Re} \{ Ce^{-i(\alpha+\beta)t} + De^{-i(\alpha-\beta)t} \}$$

$$q_{\pm} = \alpha \pm \beta \equiv \omega_{\perp} \pm \sqrt{\omega_{\perp}^2 + \frac{g}{\ell}} \approx \omega_{\perp} \pm \sqrt{\frac{g}{\ell}}$$

$$\text{since } \omega_{\perp} \approx 7 \times 10^{-5} \cos \theta \text{ rad/s} \ll \sqrt{\frac{g}{\ell}}$$

$$\text{Suppose: } x(0) = X_0 \quad y(0) = 0$$

$$x(t) = X_0 \cos \left(\sqrt{\frac{g}{\ell}} t \right) \cos(\omega_{\perp} t)$$

$$y(t) = -X_0 \cos \left(\sqrt{\frac{g}{\ell}} t \right) \sin(\omega_{\perp} t)$$

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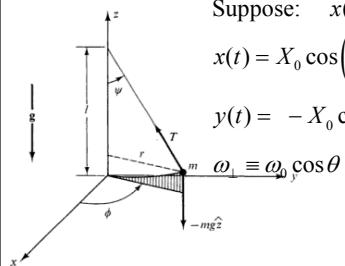
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Summary of approximate solution for Foucault pendulum:

$$\text{Suppose: } x(0) = X_0 \quad y(0) = 0$$

$$x(t) = X_0 \cos \left(\sqrt{\frac{g}{\ell}} t \right) \cos(\omega_{\perp} t)$$

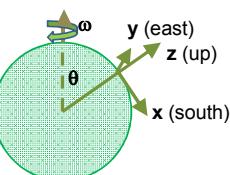
$$y(t) = -X_0 \cos \left(\sqrt{\frac{g}{\ell}} t \right) \sin(\omega_{\perp} t)$$



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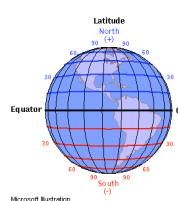
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$$\omega_{\perp} \equiv \omega_0 \cos \theta$$

$$x(t) = X_0 \cos \left(\sqrt{\frac{g}{\ell}} t \right) \cos(\omega_{\perp} t)$$

$$y(t) = -X_0 \cos \left(\sqrt{\frac{g}{\ell}} t \right) \sin(\omega_{\perp} t)$$



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