

PHY 711 Classical Mechanics and Mathematical Methods 10-10:50 AM MWF online

Plan for Lecture 39

Review

- 1. Comments on presentation and exam**
- 2. Mathematical methods**
- 3. Classical mechanics concepts**

Modified schedule for holiday 11/25/2020-11/27/2020

Schedule for weekly one-on-one meetings (EST)

Nick – 11 AM Monday – 11/23

Tim – 9 AM Tuesday – 11/24

Gao – 9 PM Tuesday -- 11/24 (11/25 in China)

Jeanette – 11 AM Friday – ??

Derek – 12 PM Friday – ??



36	Mon, 11/16/2020	Chap. 11	Heat conduction
37	Wed, 11/18/2020	Chap. 12	Viscous effects
38	Fri, 11/20/2020	Chap. 13	Elasticity
39	Mon, 11/23/2020		Review
	Wed, 11/25/2020		Thanksgiving Holiday
	Fri, 11/27/2020		Thanksgiving Holiday
40	Mon, 11/30/2020		Review
	Wed, 12/02/2020		Presentations I
	Fri, 12/04/2020		Presentations II

Comments on your project and presentation.

Presumably, you have all made good progress your project. Please let me know if I can help in any way. Nevertheless, just to remind –

The purpose of this assignment is to provide an opportunity for you to study a topic of your choice in greater depth. The general guideline for your choice of project is that it should have something to do with classical mechanics, and there should be some degree of analytical or numerical computation associated with the project. The completed project will include a presentation to the class (~20 min + 5 min for questions). After completing your presentation, please turn in a copy of your presentation slides plus any notes and references.

Schedule for PHY 711 Presentations
Fall 2020

Wednesday, December 2, 2020

	Name	Topic
10:00-10:20	Nick Corak	Numerical Solutions for Heat and Wave Equations
10:25-10:45	Derek Dremann	Analyzing Soliton Equations

Friday, December 4, 2020

	Name	Topic
10:00-10:20	Tim Carlson	Modeling TeO ₂ structure
10:25-10:45	Zhi Gao	How does a plane fly in the air?

1

Comments on take-home final –

Similar to mid-term in form

Available: Mon. Dec. 7, 2020

Due before: Mon. Dec. 14, 2020 before 11 AM

DECEMBER 2020

Sunday	Monday	Tuesday	Wednesday	Thursday	Friday	Saturday
		1	2	3	4	5
6	7	8	9	10	11	12
13	14	15	16	17	18	19
20	21	22	23	24	25	26
27	28	29	30	31		

Review of mathematical methods

Some useful identities for vectors and vector operators

$$\mathbf{a} \cdot (\mathbf{b} \times \mathbf{c}) = \mathbf{b} \cdot (\mathbf{c} \times \mathbf{a}) = \mathbf{c} \cdot (\mathbf{a} \times \mathbf{b})$$

$$\mathbf{a} \times (\mathbf{b} \times \mathbf{c}) = (\mathbf{a} \cdot \mathbf{c})\mathbf{b} - (\mathbf{a} \cdot \mathbf{b})\mathbf{c}$$

$$(\mathbf{a} \times \mathbf{b}) \cdot (\mathbf{c} \times \mathbf{d}) = (\mathbf{a} \cdot \mathbf{c})(\mathbf{b} \cdot \mathbf{d}) - (\mathbf{a} \cdot \mathbf{d})(\mathbf{b} \cdot \mathbf{c})$$

$$\nabla \times \nabla \psi = 0$$

$$\nabla \cdot (\nabla \times \mathbf{a}) = 0$$

$$\nabla \times (\nabla \times \mathbf{a}) = \nabla(\nabla \cdot \mathbf{a}) - \nabla^2 \mathbf{a}$$

$$\nabla \cdot (\psi \mathbf{a}) = \mathbf{a} \cdot \nabla \psi + \psi \nabla \cdot \mathbf{a}$$

$$\nabla \times (\psi \mathbf{a}) = \nabla \psi \times \mathbf{a} + \psi \nabla \times \mathbf{a}$$

$$\nabla(\mathbf{a} \cdot \mathbf{b}) = (\mathbf{a} \cdot \nabla)\mathbf{b} + (\mathbf{b} \cdot \nabla)\mathbf{a} + \mathbf{a} \times (\nabla \times \mathbf{b}) + \mathbf{b} \times (\nabla \times \mathbf{a})$$

$$\nabla \cdot (\mathbf{a} \times \mathbf{b}) = \mathbf{b} \cdot (\nabla \times \mathbf{a}) - \mathbf{a} \cdot (\nabla \times \mathbf{b})$$

$$\nabla \times (\mathbf{a} \times \mathbf{b}) = \mathbf{a}(\nabla \cdot \mathbf{b}) - \mathbf{b}(\nabla \cdot \mathbf{a}) + (\mathbf{b} \cdot \nabla)\mathbf{a} - (\mathbf{a} \cdot \nabla)\mathbf{b}$$

Vector relations for spherical polar coordinates

$$\nabla\psi = \hat{\mathbf{r}}\frac{\partial\psi}{\partial r} + \hat{\boldsymbol{\theta}}\frac{1}{r}\frac{\partial\psi}{\partial\theta} + \hat{\boldsymbol{\phi}}\frac{1}{r\sin\theta}\frac{\partial\psi}{\partial\phi}$$

$$\nabla^2\psi = \frac{1}{r^2}\frac{\partial}{\partial r}\left(r^2\frac{\partial\psi}{\partial r}\right) + \frac{1}{r^2\sin\theta}\frac{\partial}{\partial\theta}\left(\sin\theta\frac{\partial\psi}{\partial\theta}\right) + \frac{1}{r^2\sin^2\theta}\frac{\partial^2\psi}{\partial\phi^2}$$

$$\nabla \cdot \mathbf{A} = \frac{1}{r^2}\frac{\partial}{\partial r}(r^2A_r) + \frac{1}{r\sin\theta}\frac{\partial}{\partial\theta}(\sin\theta A_\theta) + \frac{1}{r\sin\theta}\frac{\partial A_\phi}{\partial\phi}$$

$$\begin{aligned}\nabla \times \mathbf{A} = & \hat{\mathbf{r}}\frac{1}{r\sin\theta}\left[\frac{\partial}{\partial\theta}(\sin\theta A_\phi) - \frac{\partial A_\theta}{\partial\phi}\right] \\ & + \hat{\boldsymbol{\theta}}\left[\frac{1}{r\sin\theta}\frac{\partial A_r}{\partial\phi} - \frac{1}{r}\frac{\partial}{\partial r}(rA_\phi)\right] + \hat{\boldsymbol{\phi}}\frac{1}{r}\left[\frac{\partial}{\partial r}(rA_\theta) - \frac{\partial A_r}{\partial\theta}\right]\end{aligned}$$

$$\hat{\mathbf{x}} = \hat{\mathbf{r}}\sin\theta\cos\phi + \hat{\boldsymbol{\theta}}\cos\theta\cos\phi - \hat{\boldsymbol{\phi}}\sin\phi$$

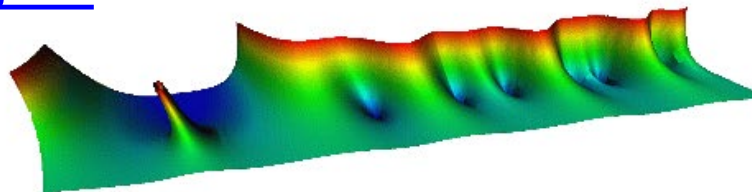
$$\hat{\mathbf{y}} = \hat{\mathbf{r}}\sin\theta\sin\phi + \hat{\boldsymbol{\theta}}\cos\theta\sin\phi + \hat{\boldsymbol{\phi}}\cos\phi$$

$$\hat{\mathbf{z}} = \hat{\mathbf{r}}\cos\theta - \hat{\boldsymbol{\theta}}\sin\theta$$

$$\frac{\partial}{\partial x} = \sin\theta\cos\phi\frac{\partial}{\partial r} + \cos\theta\cos\phi\frac{1}{r}\frac{\partial}{\partial\theta} - \frac{\sin\phi}{r\sin\theta}\frac{\partial}{\partial\phi}$$

$$\frac{\partial}{\partial y} = \sin\theta\sin\phi\frac{\partial}{\partial r} + \cos\theta\sin\phi\frac{1}{r}\frac{\partial}{\partial\theta} + \frac{\cos\phi}{r\sin\theta}\frac{\partial}{\partial\phi}$$

$$\frac{\partial}{\partial z} = \cos\theta\frac{\partial}{\partial r} - \sin\theta\frac{\partial}{\partial\theta}$$



NIST Digital Library of Mathematical Functions

Project News

2018-09-15 [DLMF Update; Version 1.0.20](#)

2018-06-22 [DLMF Update; Version 1.0.19](#)

2018-06-22 [Philip J. Davis, A&S Author, dies at age 95](#)

2018-03-27 [DLMF Update; Version 1.0.18](#)

[More news](#)

Foreword

Preface

Mathematical Introduction

1 Algebraic and Analytic Methods

2 Asymptotic Approximations

3 Numerical Methods

4 Elementary Functions

5 Gamma Function

6 Exponential, Logarithmic, Sine, and Cosine Integrals

7 Error Functions, Dawson's and Fresnel Integrals

8 Incomplete Gamma and Related Functions

9 Airy and Related Functions

10 Bessel Functions

11 Struve and Related Functions

12 Parabolic Cylinder Functions

13 Confluent Hypergeometric Functions

14 Legendre and Related Functions

15 Hypergeometric Function

16 Generalized Hypergeometric Functions & Meijer
G-Function

20 Theta Functions

21 Multidimensional Theta Functions

22 Jacobian Elliptic Functions

23 Weierstrass Elliptic and Modular Functions

24 Bernoulli and Euler Polynomials

25 Zeta and Related Functions

26 Combinatorial Analysis

27 Functions of Number Theory

28 Mathieu Functions and Hill's Equation

29 Lamé Functions

30 Spheroidal Wave Functions

31 Heun Functions

32 Painlevé Transcendents

33 Coulomb Functions

34 $3j$, $6j$, $9j$ Symbols

35 Functions of Matrix Argument

36 Integrals with Coalescing Saddles

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Notations

Example – special functions

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[10.3 Graphics](#)

§10.2 Definitions

Contents

- §10.2(i) [Bessel's Equation](#)
- §10.2(ii) [Standard Solutions](#)
- §10.2(iii) [Numerically Satisfactory Pairs of Solutions](#)

§10.2(i) Bessel's Equation

10.2.1
$$z^2 \frac{d^2 w}{dz^2} + z \frac{dw}{dz} + (z^2 - \nu^2)w = 0.$$

This differential equation has a regular singularity at $z = 0$ with indices $\pm \nu$, and an irregular singularity at $z = \infty$ of rank 1; compare §§[2.7\(i\)](#) and [2.7\(ii\)](#).

§10.2(ii) Standard Solutions

Bessel Function of the First Kind

10.2.2
$$J_\nu(z) = \left(\frac{1}{2}z\right)^\nu \sum_{k=0}^{\infty} (-1)^k \frac{\left(\frac{1}{4}z^2\right)^k}{k! \Gamma(\nu + k + 1)}.$$

This solution of [\(10.2.1\)](#) is an analytic function of $z \in \mathbb{C}$, except for a branch point at $z = 0$ when ν is not an integer. The *principal branch* of $J_\nu(z)$ corresponds to the principal value of $\left(\frac{1}{2}z\right)^\nu$ ([§4.2\(iv\)](#)) and is analytic in the z -plane cut along the interval $(-\infty, 0]$.

Complex numbers

$$i \equiv \sqrt{-1} \quad i^2 = -1$$

$$\text{Define } z = x + iy$$

$$|z|^2 = zz^* = (x + iy)(x - iy) = x^2 + y^2$$

Polar representation

$$z = \rho(\cos \phi + i \sin \phi) = \rho e^{i\phi}$$

Functions of complex variables

$$f(z) = \Re(f(z)) + i\Im(f(z)) \equiv u(x, y) + iv(x, y)$$

Derivatives: Cauchy-Riemann equations

$$\frac{\partial f(z)}{\partial x} = \frac{\partial u(z)}{\partial x} + i \frac{\partial v(z)}{\partial x} \quad \frac{\partial f(z)}{i\partial y} = \frac{\partial u(z)}{i\partial y} + i \frac{\partial v(z)}{i\partial y} = \frac{\partial v(z)}{\partial y} - i \frac{\partial u(z)}{\partial y}$$

$$\text{Argue that } \frac{df}{dz} = \frac{\partial f(z)}{\partial x} = \frac{\partial f(z)}{i\partial y} \Rightarrow \frac{\partial u(z)}{\partial x} = \frac{\partial v(z)}{\partial y} \quad \text{and} \quad \frac{\partial v(z)}{\partial x} = -\frac{\partial u(z)}{\partial y}$$

Analytic function

$f(z)$ is analytic if it is:

- continuous
- single valued
- its first derivative satisfies Cauchy-Rieman conditions

→ A closed integral of an analytic function is zero.

However:

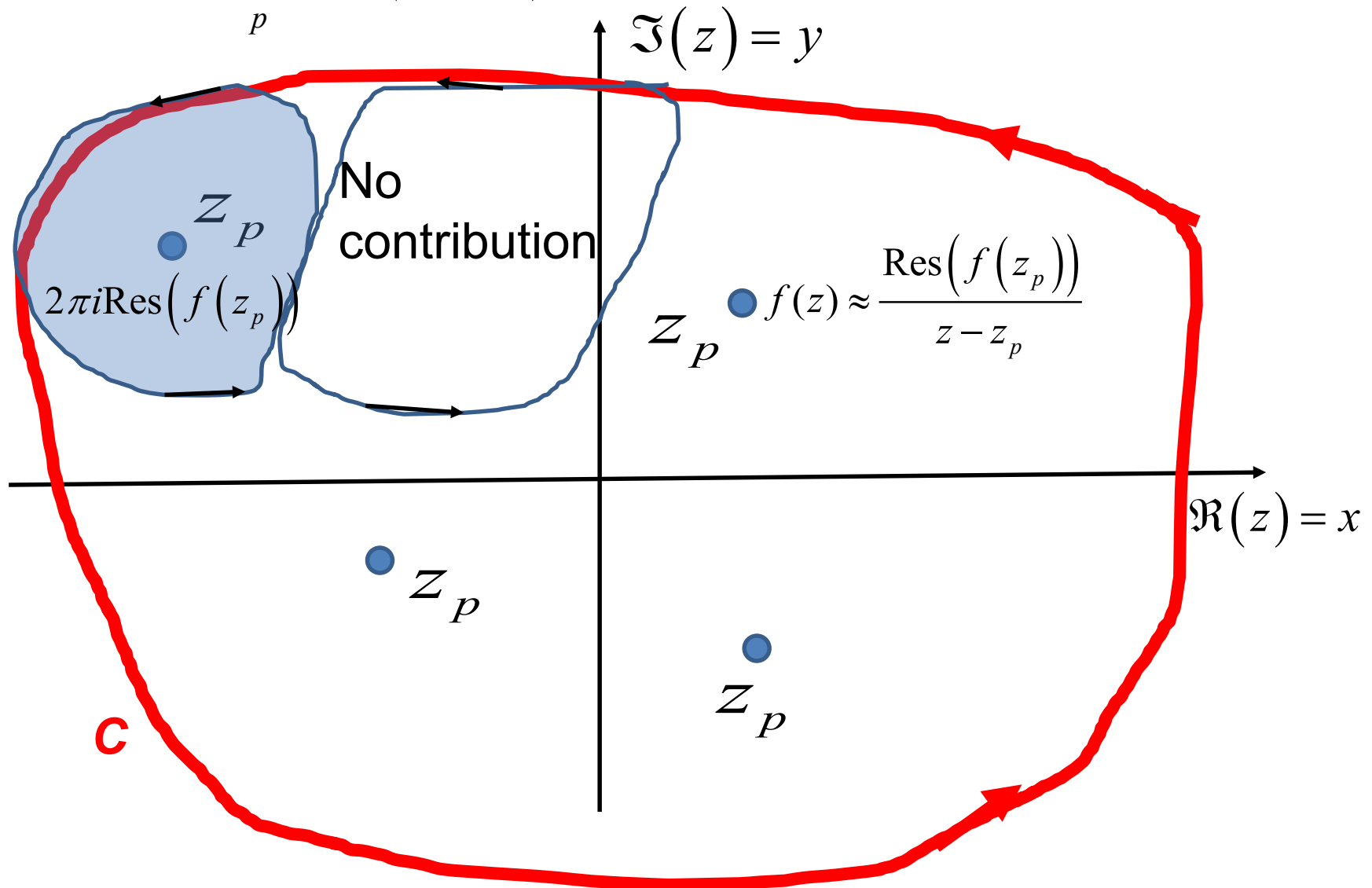
Behavior of $f(z) = \frac{1}{z^n}$ about the point $z = 0$:

For an integer n , consider

$$\oint \frac{1}{z^n} dz = \int_0^{2\pi} \frac{\rho e^{i\phi} i d\phi}{\rho^n e^{in\phi}} = \rho^{1-n} \int_0^{2\pi} e^{i(1-n)\phi} i d\phi = \begin{cases} 0 & n \neq 1 \\ 2\pi i & n = 1 \end{cases}$$

Contour integration methods --

$$\oint_C f(z) dz = 2\pi i \sum_p \text{Res}(f(z_p))$$



General formula for determining residue:

Suppose that in the neighborhood of z_p , $f(z) \approx \frac{g(z)}{(z - z_p)^m} \underset{z \rightarrow z_p}{\equiv} \frac{\text{Res}(f(z_p))}{z - z_p}$

Since $g(z)$ is analytic near z_p , we can make a Taylor expansion about z_p :

$$g(z) \approx g(z_p) + (z - z_p) \frac{dg(z_p)}{dz} + \dots + \frac{(z - z_p)^{m-1}}{(m-1)!} \frac{d^{m-1}g(z_p)}{dz^{m-1}} + \dots$$

$$\Rightarrow \text{Res}(f(z_p)) = \lim_{z \rightarrow z_p} \left\{ \frac{1}{(m-1)!} \frac{d^{m-1} \left((z - z_p)^m f(z) \right)}{dz^{m-1}} \right\}$$

Fourier transforms --

Definition of Fourier Transform for a function $f(t)$:

$$f(t) = \int_{-\infty}^{\infty} d\omega F(\omega) e^{-i\omega t}$$

Backward transform :

$$F(\omega) = \frac{1}{2\pi} \int_{-\infty}^{\infty} dt f(t) e^{i\omega t}$$

Check :

$$f(t) = \int_{-\infty}^{\infty} d\omega \left(\frac{1}{2\pi} \int_{-\infty}^{\infty} dt' f(t') e^{i\omega t'} \right) e^{-i\omega t}$$

$$f(t) = \int_{-\infty}^{\infty} dt' f(t') \left(\frac{1}{2\pi} \int_{-\infty}^{\infty} d\omega e^{i\omega(t'-t)} \right) = \int_{-\infty}^{\infty} dt' f(t') \delta(t'-t)$$

Note: The location of the 2π factor varies among texts.

Properties of Fourier transforms -- Parseval's theorem:

$$\int_{-\infty}^{\infty} dt (f(t))^* f(t) = 2\pi \int_{-\infty}^{\infty} d\omega (F(\omega))^* F(\omega)$$

Check:

$$\begin{aligned} \int_{-\infty}^{\infty} dt (f(t))^* f(t) &= \int_{-\infty}^{\infty} dt \left(\left(\int_{-\infty}^{\infty} d\omega F(\omega) e^{i\omega t} \right)^* \int_{-\infty}^{\infty} d\omega' F(\omega') e^{i\omega' t} \right) \\ &= \int_{-\infty}^{\infty} d\omega F^*(\omega) \int_{-\infty}^{\infty} d\omega' F(\omega') \int_{-\infty}^{\infty} dt e^{i(\omega' - \omega)t} \\ &= \int_{-\infty}^{\infty} d\omega F^*(\omega) \int_{-\infty}^{\infty} d\omega' F(\omega') 2\pi \delta(\omega' - \omega) \\ &= 2\pi \int_{-\infty}^{\infty} d\omega F^*(\omega) F(\omega) \end{aligned}$$

Doubly discrete Fourier Transforms

Doubly periodic functions

$$\omega \rightarrow \frac{2\pi\nu}{T} \quad t \rightarrow \frac{\mu T}{2N+1} \quad (N, \nu, \text{ and } \mu \text{ integers})$$

$$\tilde{f}_\mu = \frac{1}{2N+1} \sum_{\nu=-N}^N \tilde{F}_\nu e^{-i2\pi\nu\mu/(2N+1)}$$

$$\tilde{F}_\nu = \sum_{\mu=-N}^N \tilde{f}_\mu e^{i2\pi\nu\mu/(2N+1)}$$

➔ Fast Fourier Transforms (FFT)

Notions of eigenvalues and eigenvectors

In the context of linear algebra --

Eigenvalue properties of matrices

$$\mathbf{M}\mathbf{y}_\alpha = \lambda_\alpha \mathbf{y}_\alpha$$

Hermitian matrix: $\mathbf{H}\mathbf{y}_\alpha = \lambda_\alpha \mathbf{y}_\alpha$

$$H_{ij} = H_{ji}^*$$

Theorem for Hermitian matrices:

$$\lambda_\alpha \text{ have real values and } \mathbf{y}_\alpha^H \cdot \mathbf{y}_\beta = \delta_{\alpha\beta}$$

Unitary matrix: $\mathbf{U}\mathbf{y}_\alpha = \lambda_\alpha \mathbf{y}_\alpha$ $\mathbf{U}^H \mathbf{U} = \mathbf{I}$

$$|\lambda_\alpha| = 1 \quad \text{and} \quad \mathbf{y}_\alpha^H \cdot \mathbf{y}_\beta = \delta_{\alpha\beta}$$

In the context of Sturm-Liouville differential equations --

Notions of eigenvalues and eigenvectors -- continued

Sturm Liouville differential equations, in terms of given functions $\tau(x)$, $v(x)$, and $\sigma(x)$

Eigenfunctions:

$$\left(-\frac{d}{dx} \tau(x) \frac{d}{dx} + v(x) \right) f_n(x) = \lambda_n \sigma(x) f_n(x)$$

Orthogonality of eigenfunctions: $\int_a^b \sigma(x) f_n(x) f_m(x) dx = \delta_{nm} N_n$,

$$\text{where } N_n \equiv \int_a^b \sigma(x) (f_n(x))^2 dx.$$

Calculus of variation – a method to find a function ($y(x)$) which optimizes a particular integral relationship.

For $f\left(\left\{y(x), \frac{dy}{dx}\right\}, x\right),$

a necessary condition to extremize $\int_{x_i}^{x_f} f\left(\left\{y(x), \frac{dy}{dx}\right\}, x\right) dx :$

$$\left(\frac{\partial f}{\partial y}\right)_{x, \frac{dy}{dx}} - \frac{d}{dx} \left[\left(\frac{\partial f}{\partial (dy/dx)} \right)_{x, y} \right] = 0$$



Euler-Lagrange equation

Mechanics topics

- Scattering theory
- Lagrangian mechanics
- Hamiltonian mechanics
- Liouville theorem
- Rigid body motion
- Normal modes of oscillation about equilibrium
- Wave motion
- Fluid mechanics (ideal or including viscosity; linear and nonlinear)
- Heat conduction
- Elasticity

Note: The following review slides are necessarily brief. Please refer to the original “Extra” lecture slides for details. Please email: natalie@wfu.edu with any corrections/suggestions

Scattering theory

Note: The notion of cross section is common to many areas of physics including classical mechanics, quantum mechanics, optics, etc. Only in the **classical mechanics** can we calculate it from a knowledge of the particle trajectory as it relates to the scattering geometry.

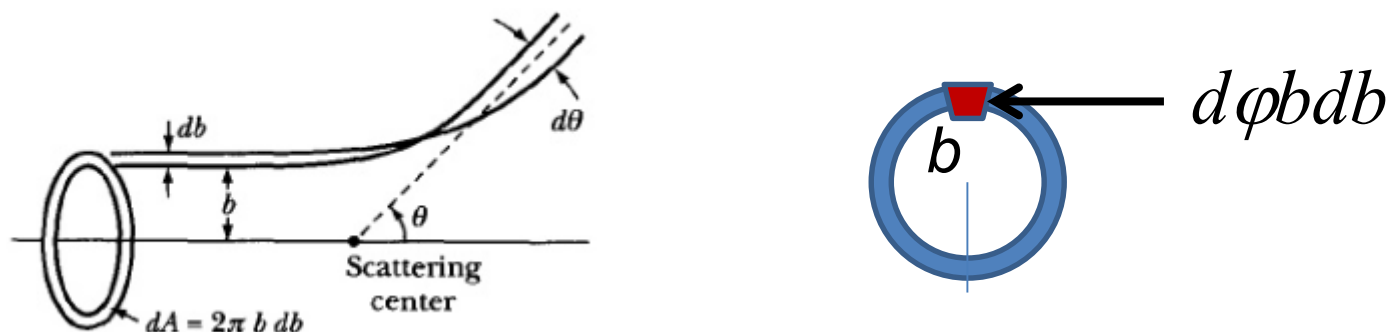


Figure from Marion & Thorton, Classical Dynamics

$$\left(\frac{d\sigma}{d\Omega} \right) = \frac{d\phi b db}{d\phi \sin\theta d\theta} = \frac{b}{\sin\theta} \left| \frac{db}{d\theta} \right|$$

Note: We are assuming that the process is isotropic in ϕ

Lagrangian mechanics

Given the Lagrangian function: $L = L(\{q_\sigma\}, \{\dot{q}_\sigma\}, t) \equiv T - U$,

The physical trajectories of the generalized coordinates $\{q_\sigma(t)\}$

Are those which minimize the action: $S = \int L(\{q_\sigma\}, \{\dot{q}_\sigma\}, t) dt$

Euler-Lagrange equations:

$$\sum_{\sigma} \left(\frac{d}{dt} \frac{\partial L}{\partial \dot{q}_{\sigma}} - \frac{\partial L}{\partial q_{\sigma}} \right) \delta q_{\sigma} = 0 \quad \Rightarrow \text{for each } \sigma : \quad \left(\frac{d}{dt} \frac{\partial L}{\partial \dot{q}_{\sigma}} - \frac{\partial L}{\partial q_{\sigma}} \right) = 0$$

For the case that there both mechanical and

electromagnetic contributions in terms of electric and magnetic fields:

$$\mathbf{E}(\mathbf{r}, t) = -\nabla \Phi(\mathbf{r}, t) - \frac{1}{c} \frac{\partial \mathbf{A}(\mathbf{r}, t)}{\partial t} \quad \mathbf{B}(\mathbf{r}, t) = \nabla \times \mathbf{A}(\mathbf{r}, t)$$

$$L = T - U_{\text{mech}} - q\Phi(\mathbf{r}, t) + \frac{q}{c} \dot{\mathbf{r}} \cdot \mathbf{A}(\mathbf{r}, t)$$

Recipe for constructing the Hamiltonian and analyzing the equations of motion

1. Construct Lagrangian function: $L = L(\{q_\sigma(t)\}, \{\dot{q}_\sigma(t)\}, t)$
2. Compute generalized momenta: $p_\sigma \equiv \frac{\partial L}{\partial \dot{q}_\sigma}$
3. Construct Hamiltonian expression: $H = \sum_\sigma \dot{q}_\sigma p_\sigma - L$
4. Form Hamiltonian function: $H = H(\{q_\sigma(t)\}, \{p_\sigma(t)\}, t)$
5. Analyze canonical equations of motion:

$$\frac{dq_\sigma}{dt} = \frac{\partial H}{\partial p_\sigma} \qquad \frac{dp_\sigma}{dt} = -\frac{\partial H}{\partial q_\sigma}$$

Question – When can you bypass the 5 step derivation process and directly write the Hamiltonian of the system as

$$H\left(\{q_{\sigma}(t)\},\{p_{\sigma}(t)\},t\right)=\sum_{\sigma}\frac{p_{\sigma}^2}{2m_{\sigma}}+V(\{q_{\sigma}\})$$

1. Only when Natalie Holzwarth is not looking
2. When you have a simple system that has no explicit velocity and/or time dependence
3. Usually

Important tool for analyzing Lagrangian and/or Hamiltonian systems -- finding constants of the motion

In Lagrangian formulation --

For independent generalized coordinates $q_\sigma(t)$:

$$L = L(\{q_\sigma(t)\}, \{\dot{q}_\sigma(t)\}, t) \quad \Rightarrow \quad \frac{d}{dt} \frac{\partial L}{\partial \dot{q}_\sigma} - \frac{\partial L}{\partial q_\sigma} = 0$$

$$\text{Note that if } \frac{\partial L}{\partial q_\sigma} = 0, \text{ then } \frac{d}{dt} \frac{\partial L}{\partial \dot{q}_\sigma} = 0 \quad \Rightarrow \quad \frac{\partial L}{\partial \dot{q}_\sigma} = (\text{constant})$$

$$\text{Additionally: } \frac{d}{dt} \left(L - \sum_\sigma \frac{\partial L}{\partial \dot{q}_\sigma} \dot{q}_\sigma \right) = \frac{\partial L}{\partial t}$$

$$\text{For } \frac{\partial L}{\partial t} = 0 \quad \Rightarrow \quad L - \sum_\sigma \frac{\partial L}{\partial \dot{q}_\sigma} \dot{q}_\sigma = -E \quad (\text{constant})$$

Constants of the motion in the Hamiltonian formulation

$$H = H(\{q_\sigma(t)\}, \{p_\sigma(t)\}, t)$$

$$\frac{dq_\sigma}{dt} = \frac{\partial H}{\partial p_\sigma} \Rightarrow \text{constant } q_\sigma \text{ if } \frac{\partial H}{\partial p_\sigma} = 0$$

$$\frac{dp_\sigma}{dt} = -\frac{\partial H}{\partial q_\sigma} \Rightarrow \text{constant } p_\sigma \text{ if } \frac{\partial H}{\partial q_\sigma} = 0$$

$$\frac{dH}{dt} = \sum_\sigma \left(\frac{\partial H}{\partial q_\sigma} \dot{q}_\sigma + \frac{\partial H}{\partial p_\sigma} \dot{p}_\sigma \right) + \frac{\partial H}{\partial t}$$

$$\frac{dH}{dt} = \sum_\sigma (-\dot{p}_\sigma \dot{q}_\sigma + \dot{q}_\sigma \dot{p}_\sigma) + \frac{\partial H}{\partial t} = \frac{\partial H}{\partial t}$$

$$\Rightarrow \text{constant } H \text{ if } \frac{\partial H}{\partial t} = 0$$

Question – Why use this fancy formalism when simple conservation of energy or momentum intuitively apply?

- a. You should use your intuition whenever possible.
- b. You should never trust your intuition.
- c. The equations should be consistent with correct intuitive solutions and also reveal additional solutions (perhaps beyond intuition)

Liouville's Theorem (1838)

The density of representative points in phase space corresponding to the motion of a system of particles remains constant during the motion.

Denote the density of particles in phase space: $D = D(\{q_\sigma(t)\}, \{p_\sigma(t)\}, t)$

$$\frac{dD}{dt} = \sum_{\sigma} \left(\frac{\partial D}{\partial q_{\sigma}} \dot{q}_{\sigma} + \frac{\partial D}{\partial p_{\sigma}} \dot{p}_{\sigma} \right) + \frac{\partial D}{\partial t}$$

According to Liouville's theorem: $\frac{dD}{dt} = 0$

Rigid body motion

Moment of inertia tensor :

$$\vec{\mathbf{I}} \equiv \sum_p m_p (\mathbf{1} r_p^2 - \mathbf{r}_p \mathbf{r}_p) \quad (\text{dyad notation})$$

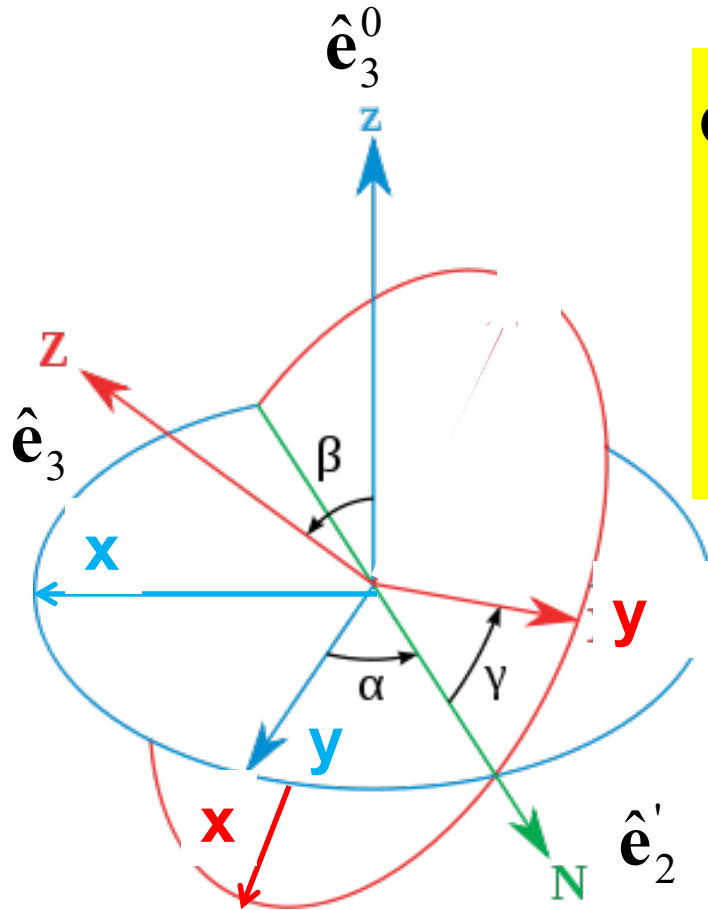
In a reference frame attached to the object, there are 3 moments of inertia and 3 distinct principal axes

Representation of rotational kinetic energy:

$$\begin{aligned} T(\alpha, \beta, \gamma, \dot{\alpha}, \dot{\beta}, \dot{\gamma}) &= \frac{1}{2} I_1 \tilde{\omega}_1^2 + \frac{1}{2} I_2 \tilde{\omega}_2^2 + \frac{1}{2} I_3 \tilde{\omega}_3^2 \\ &= \frac{1}{2} I_1 [\dot{\alpha}(-\sin \beta \cos \gamma) + \dot{\beta} \sin \gamma]^2 \\ &\quad + \frac{1}{2} I_2 [\dot{\alpha}(\sin \beta \sin \gamma) + \dot{\beta} \cos \gamma]^2 \\ &\quad + \frac{1}{2} I_3 [\dot{\alpha} \cos \beta + \dot{\gamma}]^2 \end{aligned}$$

Euler's transformation between body fixed and inertial reference frames

$$\tilde{\boldsymbol{\omega}} = \dot{\alpha} \hat{\mathbf{e}}_3^0 + \dot{\beta} \hat{\mathbf{e}}_2' + \dot{\gamma} \hat{\mathbf{e}}_3$$

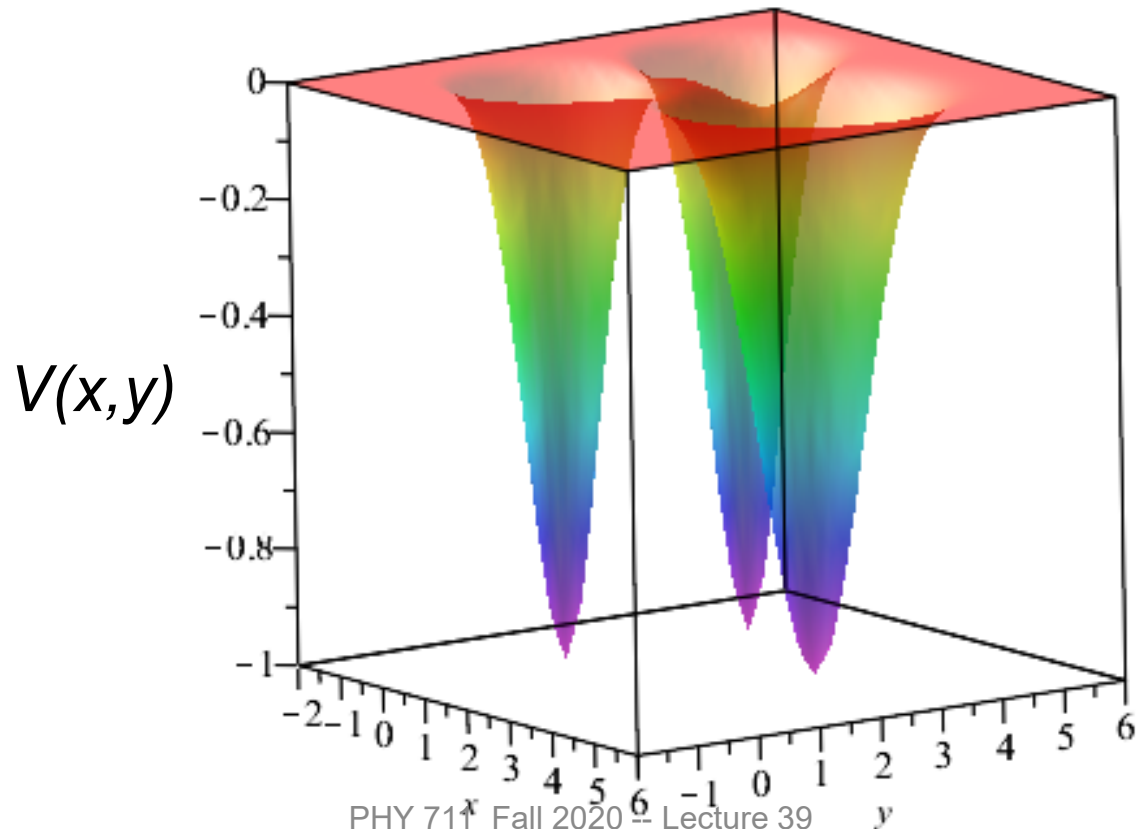


$$\begin{aligned} \tilde{\boldsymbol{\omega}} = & \left[\dot{\alpha} (-\sin \beta \cos \gamma) + \dot{\beta} \sin \gamma \right] \hat{\mathbf{e}}_1 \\ & + \left[\dot{\alpha} (\sin \beta \sin \gamma) + \dot{\beta} \cos \gamma \right] \hat{\mathbf{e}}_2 \\ & + \left[\dot{\alpha} \cos \beta + \dot{\gamma} \right] \hat{\mathbf{e}}_3 \end{aligned}$$

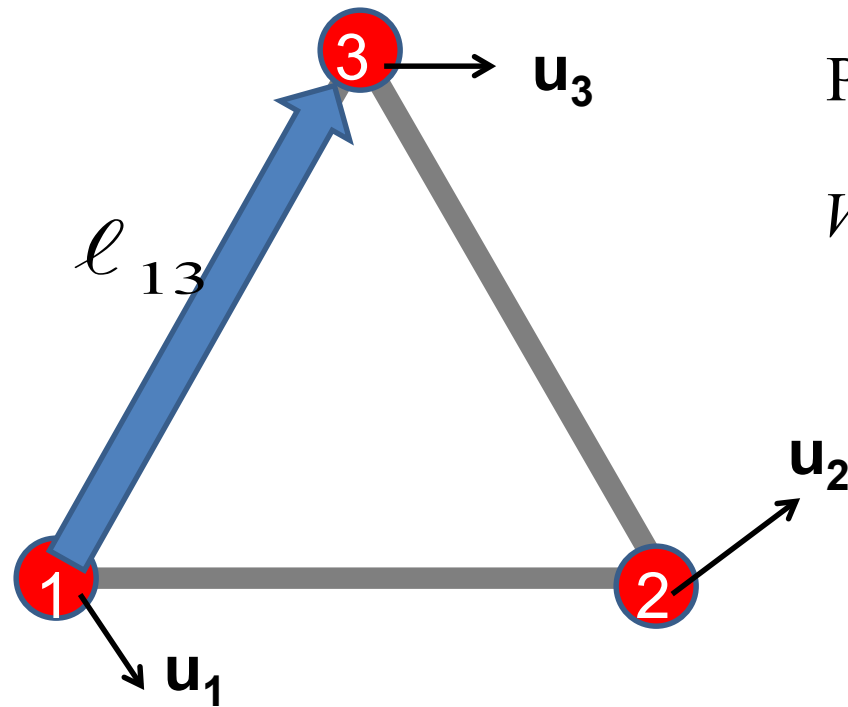
Normal modes of vibration -- potential in 2 and more dimensions

$$V(x, y) \approx V(x_{eq}, y_{eq}) + \frac{1}{2} \left(x - x_{eq} \right)^2 \frac{\partial^2 V}{\partial x^2} \bigg|_{x_{eq}, y_{eq}}$$

$$+ \frac{1}{2} \left(y - y_{eq} \right)^2 \frac{\partial^2 V}{\partial y^2} \bigg|_{x_{eq}, y_{eq}} + \left(x - x_{eq} \right) \left(y - y_{eq} \right) \frac{\partial^2 V}{\partial x \partial y} \bigg|_{x_{eq}, y_{eq}}$$



Example – normal modes of a system with the symmetry of an equilateral triangle -- continued



Potential contribution for spring 13:

$$\begin{aligned}
 V_{13} &= \frac{1}{2}k \left(|\ell_{13} + \mathbf{u}_3 - \mathbf{u}_1| - |\ell_{13}| \right)^2 \\
 &\approx \frac{1}{2}k \left(\frac{\ell_{13} \cdot (\mathbf{u}_3 - \mathbf{u}_1)}{|\ell_{13}|} \right)^2 \\
 &\approx \frac{1}{2}k \left(\frac{1}{2}(u_{x3} - u_{x1}) + \frac{\sqrt{3}}{2}(u_{y3} - u_{y1}) \right)^2
 \end{aligned}$$

$$\ell_{13} = |\ell_{13}| \left(\frac{1}{2}\hat{\mathbf{x}} + \frac{\sqrt{3}}{2}\hat{\mathbf{y}} \right)$$

Example – normal modes of a system with the symmetry of an equilateral triangle -- continued

Potential contributions: $V = V_{12} + V_{13} + V_{23}$

$$\approx \frac{1}{2}k \left(\frac{\ell_{12} \cdot (\mathbf{u}_2 - \mathbf{u}_1)}{|\ell_{12}|} \right)^2 + \frac{1}{2}k \left(\frac{\ell_{13} \cdot (\mathbf{u}_3 - \mathbf{u}_1)}{|\ell_{13}|} \right)^2$$

$$+ \frac{1}{2}k \left(\frac{\ell_{23} \cdot (\mathbf{u}_3 - \mathbf{u}_2)}{|\ell_{23}|} \right)^2$$

$$\approx \frac{1}{2}k (u_{x2} - u_{x1})^2$$

$$+ \frac{1}{2}k \left(\frac{1}{2}(u_{x3} - u_{x1}) + \frac{\sqrt{3}}{2}(u_{y3} - u_{y1}) \right)^2$$

$$+ \frac{1}{2}k \left(\frac{1}{2}(u_{x2} - u_{x3}) - \frac{\sqrt{3}}{2}(u_{y2} - u_{y3}) \right)^2$$

Example – normal modes of a system with the symmetry of an equilateral triangle -- continued

$$\frac{k}{m} \begin{bmatrix} \frac{5}{4} & -1 & -\frac{1}{4} & \frac{1}{4}\sqrt{3} & 0 & -\frac{1}{4}\sqrt{3} \\ -1 & \frac{5}{4} & -\frac{1}{4} & 0 & -\frac{1}{4}\sqrt{3} & \frac{1}{4}\sqrt{3} \\ -\frac{1}{4} & -\frac{1}{4} & \frac{1}{2} & -\frac{1}{4}\sqrt{3} & \frac{1}{4}\sqrt{3} & 0 \\ \frac{1}{4}\sqrt{3} & 0 & -\frac{1}{4}\sqrt{3} & \frac{3}{4} & 0 & -\frac{3}{4} \\ 0 & -\frac{1}{4}\sqrt{3} & \frac{1}{4}\sqrt{3} & 0 & \frac{3}{4} & -\frac{3}{4} \\ -\frac{1}{4}\sqrt{3} & \frac{1}{4}\sqrt{3} & 0 & -\frac{3}{4} & -\frac{3}{4} & \frac{3}{2} \end{bmatrix} \begin{bmatrix} u_{x1} \\ u_{x2} \\ u_{x3} \\ u_{y1} \\ u_{y2} \\ u_{y3} \end{bmatrix} = \omega^2 \begin{bmatrix} u_{x1} \\ u_{x2} \\ u_{x3} \\ u_{y1} \\ u_{y2} \\ u_{y3} \end{bmatrix}$$

Discrete particle interactions $\rightarrow$ continuous media $\rightarrow$
The wave equation

Initial value solutions $\mu(x, t)$ to the wave equation;
attributed to D'Alembert:

$$\frac{\partial^2 \mu}{\partial t^2} - c^2 \frac{\partial^2 \mu}{\partial x^2} = 0 \quad \text{where } \mu(x, 0) = \varphi(x) \text{ and } \frac{\partial \mu}{\partial t}(x, 0) = \psi(x)$$

$$\Rightarrow \mu(x, t) = \frac{1}{2} (\varphi(x - ct) + \varphi(x + ct)) + \frac{1}{2c} \int_{x-ct}^{x+ct} \psi(x') dx'$$

Mechanical motion of fluids

Newton's equations for fluids

Use Euler formulation; following “particles” of fluid

Variables: Density $\rho(x,y,z,t)$

 Pressure $p(x,y,z,t)$

 Velocity $\mathbf{v}(x,y,z,t)$

Navier-Stokes equation

$$\frac{\partial \mathbf{v}}{\partial t} + (\mathbf{v} \cdot \nabla) \mathbf{v} = \mathbf{f} - \frac{1}{\rho} \nabla p + \underbrace{\frac{\eta}{\rho} \nabla^2 \mathbf{v} + \frac{1}{\rho} \left(\zeta + \frac{1}{3} \eta \right) \nabla (\nabla \cdot \mathbf{v})}_{\text{Viscosity contributions}}$$

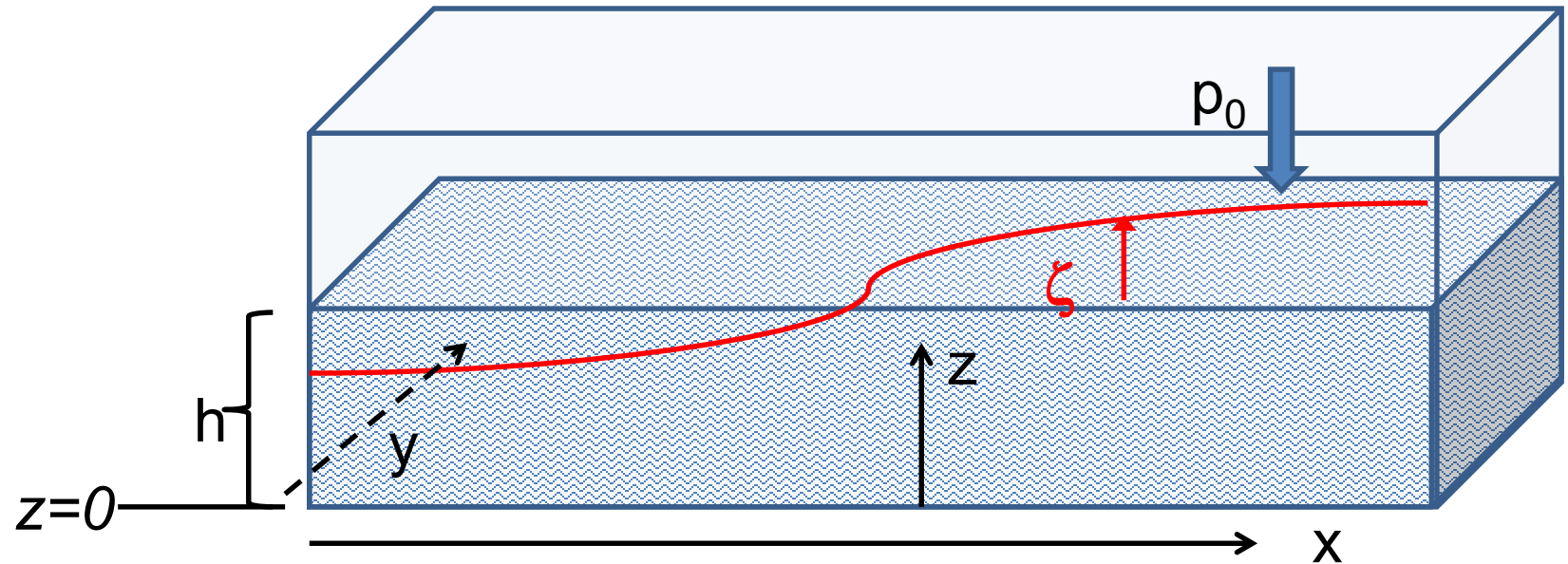
Continuity condition

$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \mathbf{v}) = 0$$

Viscosity contributions

Fluid mechanics of incompressible fluid plus surface

Non-linear effects in surface waves:



Dominant non-linear effects $\Rightarrow$ soliton solutions

$$\zeta(x, t) = \eta_0 \operatorname{sech}^2 \left(\sqrt{\frac{3\eta_0}{h}} \frac{x - ct}{2h} \right) \quad \eta_0 = \text{constant}$$

$$\text{where } c = \sqrt{\frac{gh}{1 - \eta_0/h}} \approx \sqrt{gh} \left(1 + \frac{\eta_0}{2h} \right)$$

To be continued Mon. Nov. 30, 2020

Topics ?