

**PHY 114 A General Physics II**  
**11 AM-12:15 PM TR Olin 101**

**Plan for Lecture 25 (Chapters 44-45):**

**Some topics in nuclear physics**

- 1. Nuclear binding energies**
- 2. Radioactivity**
- 3. Nuclear reactions**

13	03/08/2012	Faraday's law	<a href="#">31.1-31.5</a>	<a href="#">31.12.31.23.31.40</a>	03/20/2012
	03/13/2012	No class (Spring Break)			
	03/15/2012	No class (Spring Break)			
14	03/20/2012	Induction and AC circuits	<a href="#">32.1-32.6</a>	<a href="#">32.4.32.20.32.43</a>	03/22/2012
15	03/22/2012	AC circuits	<a href="#">33.1-33.9</a>	<a href="#">33.8.33.24.33.71</a>	03/27/2012
16	03/27/2012	Electromagnetic waves	<a href="#">34.1-34.3</a>	<a href="#">34.3.34.10.34.13</a>	03/29/2012
17	03/29/2012	Electromagnetic waves	<a href="#">34.4-34.7</a>	<a href="#">34.22.34.46.34.57</a>	04/03/2012
18	04/03/2012	Ray optics Evening exam	<a href="#">35.1-35.8</a>	<a href="#">35.20.35.27.35.35</a>	04/10/2012
19	04/05/2012	Image formation Evening exam	<a href="#">36.1-36.4</a>	<a href="#">36.8.36.31.36.42</a>	04/10/2012
20	04/10/2012	Image formation	<a href="#">36.5-36.10</a>	<a href="#">36.52.36.54.36.64</a>	04/12/2012
21	04/12/2012	Wave interference	<a href="#">37.1-37.6</a>	<a href="#">37.2.37.19.37.29</a>	04/17/2012
22	04/17/2012	Diffraction	<a href="#">38.1-38.6</a>	<a href="#">38.24.38.30.38.37</a>	04/19/2012
23	04/19/2012	Quantum Physics	<a href="#">40.1-42.10</a>	<a href="#">40.41.41.12.42.10</a>	04/24/2012
24	04/24/2012	Molecules and solids Evening exam	<a href="#">43.1-43.8</a>	<a href="#">43.2.43.40.43.43</a>	05/01/2012
25	04/26/2012	Nuclear reactions Evening exam	<a href="#">45.1-45.4</a>	<a href="#">45.6.45.20.45.30</a>	05/01/2012
26	05/01/2012	Nuclear radiation	<a href="#">45.5-45.7</a>		
	05/08/2012	Final exam 9 AM			

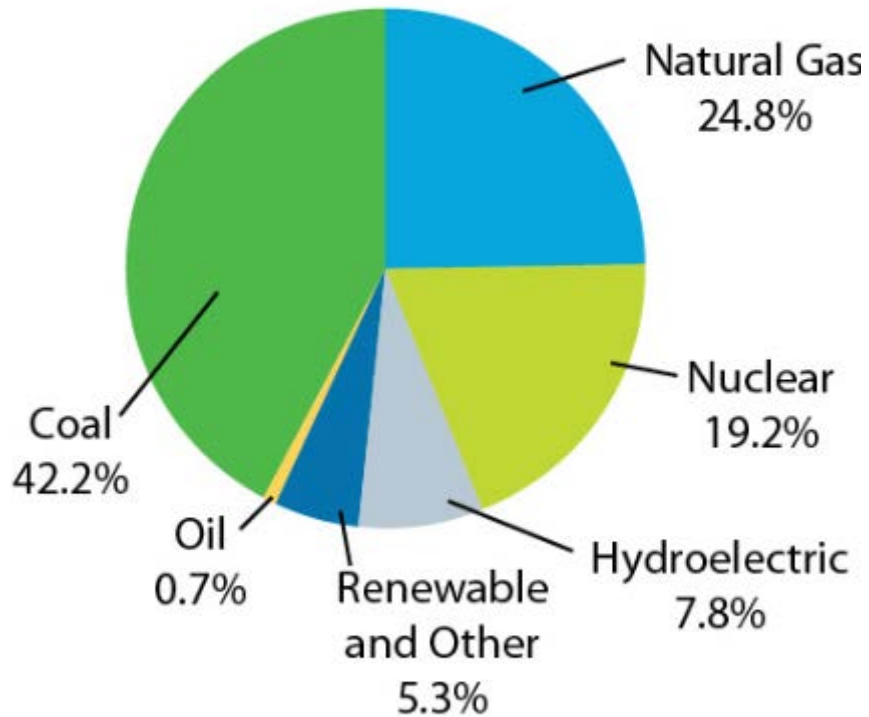


What do you think of when you hear the phrase nuclear reaction?

- A. Clean energy source
- B. Radiation danger
- C. Nuclear weapons
- D. No opinion

<http://www.nei.org>

## U.S. Electricity Generation Fuel Shares 2011



*Source: U.S. Energy Information  
Administration*

# Properties of Nuclei

Z Atomic number (number of protons) N Neutron number

A = Z+N (number of nucleons)

Group I	Group II	Transition elements										Group III	Group IV	Group V	Group VI	Group VII	Group 0						
H 1 $1s^1$		<div style="display: flex; justify-content: center; align-items: center;"> <div style="font-size: 2em; margin-right: 10px;">Periodic Table</div> <div style="font-size: 3em; margin-right: 10px;">Z</div> </div>																				H 1 $1s^1$	He 2 $1s^2$
Li 3 $2s^1$	Be 4 $2s^2$											B 5 $2p^1$	C 6 $2p^2$	N 7 $2p^3$	O 8 $2p^4$	F 9 $2p^5$	Ne 10 $2p^6$						
Na 11 $3s^1$	Mg 12 $3s^2$											Al 13 $3p^1$	Si 14 $3p^2$	P 15 $3p^3$	S 16 $3p^4$	Cl 17 $3p^5$	Ar 18 $3p^6$						
K 19 $4s^1$	Ca 20 $4s^2$	Sc 21 $3d^1 4s^2$	Ti 22 $3d^2 4s^2$	V 23 $3d^3 4s^2$	Cr 24 $3d^5 4s^1$	Mn 25 $3d^5 4s^2$	Fe 26 $3d^6 4s^2$	Co 27 $3d^7 4s^2$	Ni 28 $3d^8 4s^2$	Cu 29 $3d^{10} 4s^1$	Zn 30 $3d^{10} 4s^2$	Ga 31 $4p^1$	Ge 32 $4p^2$	As 33 $4p^3$	Se 34 $4p^4$	Br 35 $4p^5$	Kr 36 $4p^6$						
Rb 37 $5s^1$	Sr 38 $5s^2$	Y 39 $4d^1 5s^2$	Zr 40 $4d^2 5s^2$	Nb 41 $4d^4 5s^1$	Mo 42 $4d^5 5s^1$	Tc 43 $4d^5 5s^2$	Ru 44 $4d^7 5s^1$	Rh 45 $4d^8 5s^1$	Pd 46 $4d^{10}$	Ag 47 $4d^{10} 5s^1$	Cd 48 $4d^{10} 5s^2$	In 49 $5p^1$	Sn 50 $5p^2$	Sb 51 $5p^3$	Te 52 $5p^4$	I 53 $5p^5$	Xe 54 $5p^6$						
Cs 55 $6s^1$	Ba 56 $6s^2$	57-71* $6s^2$	Hf 72 $5d^2 6s^2$	Ta 73 $5d^3 6s^2$	W 74 $5d^4 6s^2$	Re 75 $5d^5 6s^2$	Os 76 $5d^6 6s^2$	Ir 77 $5d^7 6s^2$	Pt 78 $5d^9 6s^1$	Au 79 $5d^{10} 6s^1$	Hg 80 $5d^{10} 6s^2$	Tl 81 $6p^1$	Pb 82 $6p^2$	Bi 83 $6p^3$	Po 84 $6p^4$	At 85 $6p^5$	Rn 86 $6p^6$						
Fr 87 $7s^1$	Ra 88 $7s^2$	89-103** $7s^2$	Rf 104 $6d^2 7s^2$	Db 105 $6d^3 7s^2$	Sg 106 $6d^4 7s^2$	Bh 107 $6d^5 7s^2$	Hs 108 $6d^6 7s^2$	Mt 109 $6d^7 7s^2$	Ds 110 $6d^9 7s^1$	Rg 111 $6d^9 7s^1$	112 $6d^9 7s^1$		114 $6p^2$		116 $6p^4$								

\*Lanthanide series

La 57 $5d^1 6s^2$	Ce 58 $5d^1 4f^1 6s^2$	Pr 59 $4f^3 6s^2$	Nd 60 $4f^4 6s^2$	Pm 61 $4f^5 6s^2$	Sm 62 $4f^6 6s^2$	Eu 63 $4f^7 6s^2$	Gd 64 $5d^1 4f^7 6s^2$	Tb 65 $5d^1 4f^8 6s^2$	Dy 66 $4f^{10} 6s^2$	Ho 67 $4f^{11} 6s^2$	Er 68 $4f^{12} 6s^2$	Tm 69 $4f^{13} 6s^2$	Yb 70 $4f^{14} 6s^2$	Lu 71 $5d^1 4f^{14} 6s^2$
Ac 89 $6d^1 7s^2$	Th 90 $6d^2 7s^2$	Pa 91 $5f^2 6d^1 7s^2$	U 92 $5f^3 6d^1 7s^2$	Np 93 $5f^4 6d^1 7s^2$	Pu 94 $5f^6 7s^2$	Am 95 $5f^7 7s^2$	Cm 96 $5f^7 6d^1 7s^2$	Bk 97 $5f^8 6d^1 7s^2$	Cf 98 $5f^{10} 7s^2$	Es 99 $5f^{11} 7s^2$	Fm 100 $5f^{12} 7s^2$	Md 101 $5f^{13} 7s^2$	No 102 $5f^{14} 7s^2$	Lr 103 $5f^{14} 6d^1 7s^2$

\*\*Actinide series

# Constituents of the nucleus

**TABLE 44.1** *Masses of Selected Particles in Various Units*

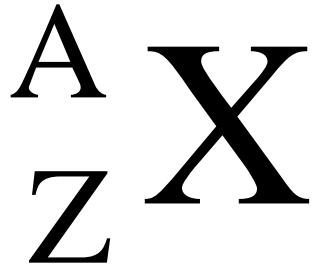
Particle	kg	Mass u	MeV/c <sup>2</sup>
Proton	$1.672\ 62 \times 10^{-27}$	1.007 276	938.27
Neutron	$1.674\ 93 \times 10^{-27}$	1.008 665	939.57
Electron	$9.109\ 38 \times 10^{-31}$	$5.485\ 79 \times 10^{-4}$	0.510 999
$^1_1\text{H}$ atom	$1.673\ 53 \times 10^{-27}$	1.007 825	938.783
$^4_2\text{He}$ nucleus	$6.644\ 66 \times 10^{-27}$	4.001 506	3 727.38
$^{12}_6\text{C}$ atom	$1.992\ 65 \times 10^{-27}$	12.000 000	11 177.9

Example for proton :

$$\begin{aligned} mc^2 &= 1.67263 \times 10^{-27} (299792458)^2 = 1.50 \times 10^{-10} \text{ J} \\ &= 938.27 \text{ MeV} \end{aligned}$$

Nuclear notation:

Z Atomic number (number of protons)  
A Nucleon number

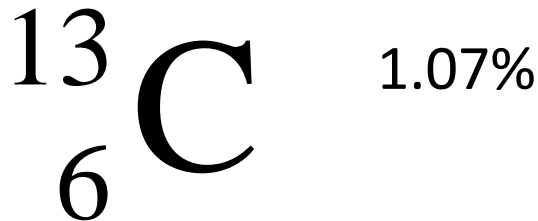
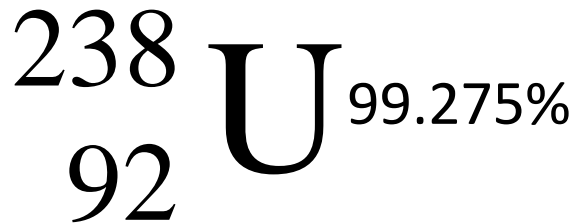
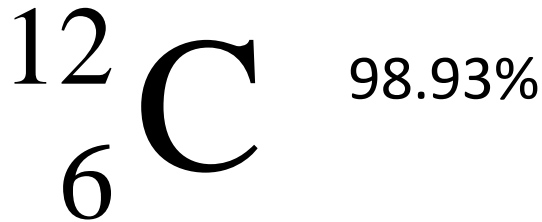
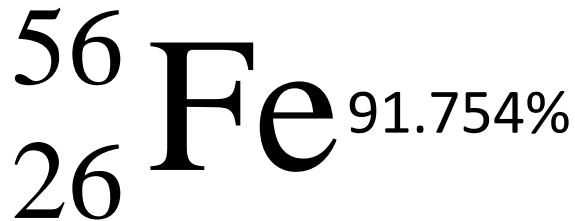


Natural abundance

Natural abundance

Some examples:

Some isotopes of C:



# Atomic Weights and Isotopic Compositions for All E

Z	A	Isotope	Relative Atomic Mass	Isotopic Composition	Standard Atomic Weight
1	H	1	1.007 825 032 1(4)	99.9885(70)% 0.0115(70)	1.007 94(7)
	D	2	2.014 101 778 0(4)		
	T	3	3.016 049 2675(11)		
2	He	3	3.016 029 309 7(9)	0.000 137(3)	4.002 602(2)
		4	4.002 603 2497(10)	99.999 863(3)	
3	Li	6	6.015 122 3(5)	7.59(4)	6.941(2)
		7	7.016 004 0(5)	92.41(4)	
26	Fe	54	53.939 6148(14)	5.845(35)	55.845(2)
		56	55.934 9421(15)	91.754(36)	
		57	56.935 3987(15)	2.119(10)	
		58	57.933 2805(15)	0.282(4)	
27	Co	59	58.933 2002(15)	100	58.933 200(9)

Ref: <http://www.physics.nist.gov/PhysRefData/Compositions/index.html>



mass of neutral atom → includes Z electrons

Z	A				
27	Co	59	58.933 2002(15)	100	58.933 200(9)

Mass of nucleus:

$$M_{\text{sum}} = Zm_p + (A-Z)m_n + Zm_e = 59.474281 \text{ u}$$

$$\Delta M = 0.5410806 \text{ u}$$

What should we do with this mass deficit?

- (A) Chalk it up to inaccuracy of my calculator.
- (B) Figure that NIST made a mistake.
- (C) Give up on physics as a quantitative science.
- (D) Find some meaning associated with  $\Delta M$ .

mass of neutral atom

→ includes Z electrons

Z	A		mass of neutral atom		mass of neutral atom
27	Co	59	58.933 2002(15)		58.933 200(9)

Mass of nucleus:

$$M_{\text{sum}} = Zm_p + (A-Z)m_n + Zm_e = 59.474281 \text{ u}$$

$$\Delta M = 0.5410806 \text{ u} = 504.0135 \text{ MeV}/c^2$$

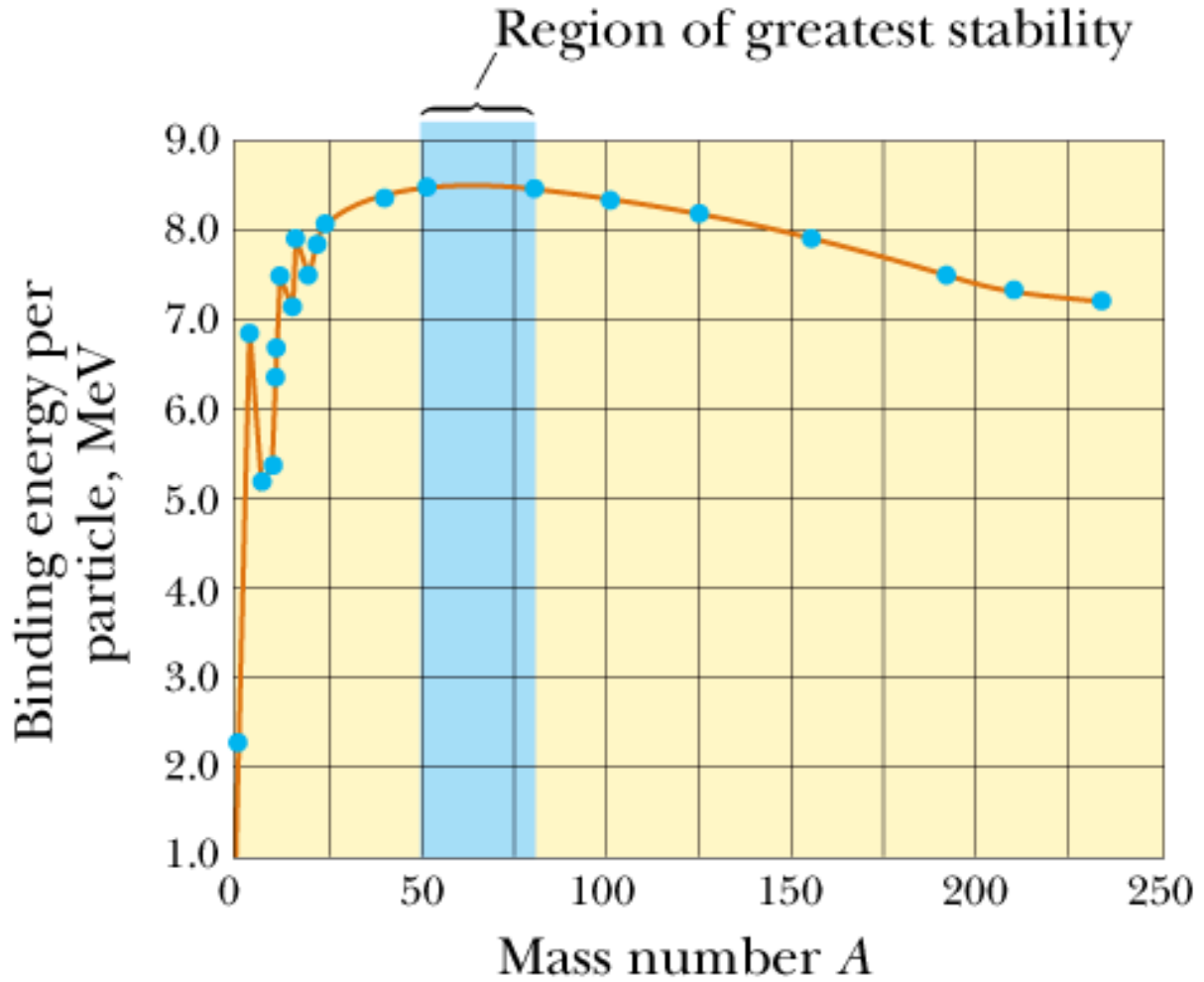
$$\Delta M/A = 8.5426 \text{ MeV}/c^2 \text{ /nucleon} \rightarrow \text{energy associated with nuclear "binding"}$$

Another example:

2	He	3	3.016 029 309 7(9)		0.000 137(3)	4.002 602(2)
		4	4.002 603 2497(10)		99.999 863(3)	

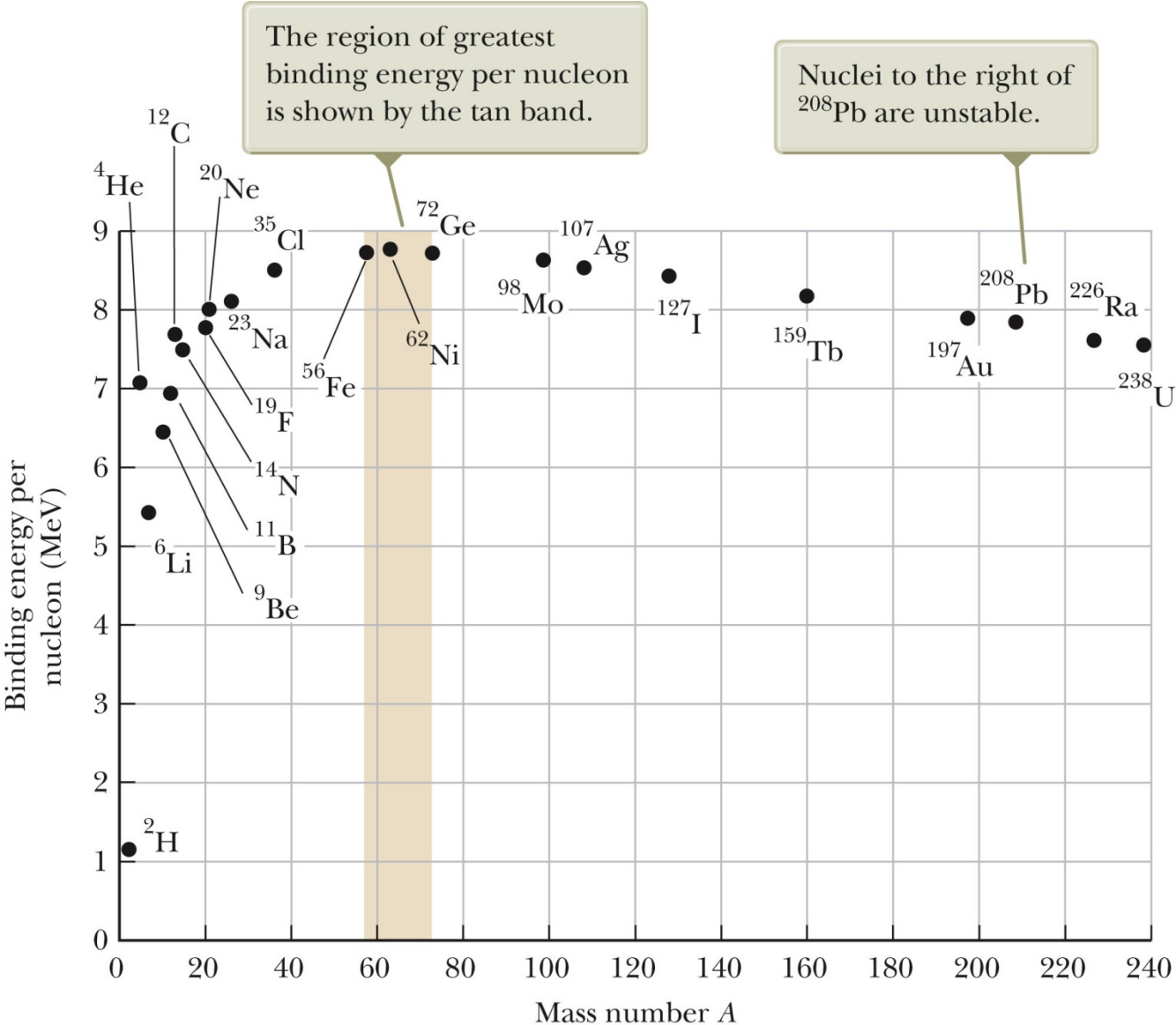
$$\Delta M/A(^4_2\text{He}) = 7.1 \text{ MeV}/c^2 \text{ /nucleon}$$

$$\Delta M/A(^3_2\text{He}) = 2.6 \text{ MeV}/c^2 \text{ /nucleon}$$

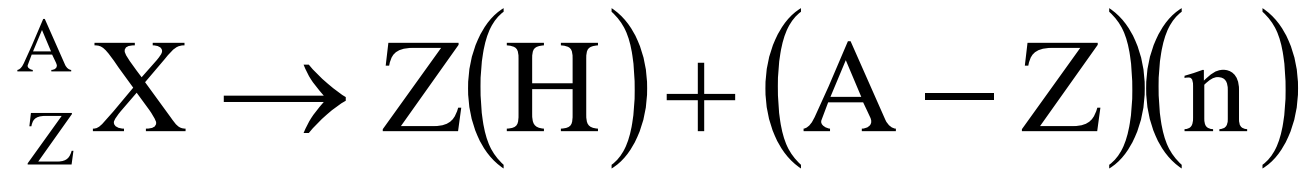


U.

# Nuclear binding energies per nucleon



# Nuclear binding energy :



Neutral atom  
with Z electrons  
and Z protons

Neutral H atom

Neutrons

Nuclear forces are so strong that within the nucleus, mass can be converted to energy and visa versa:  $E = mc^2$  .

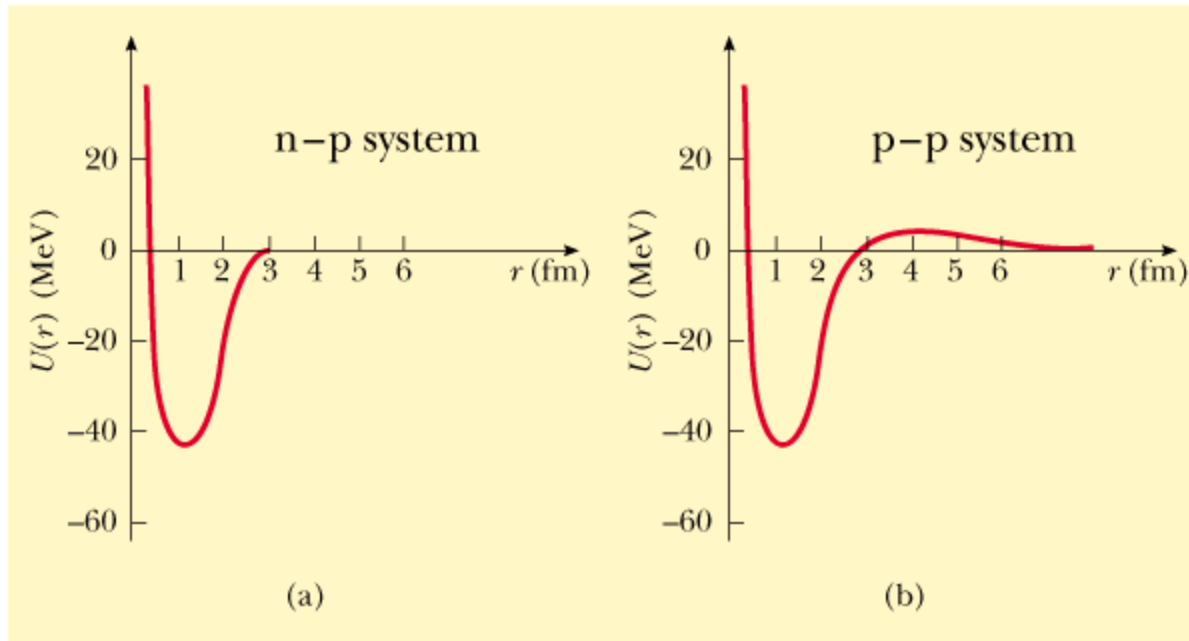
Nuclear binding energy :

$$E_b = -\left[ M\left({}_Z^A X\right) - Z \cdot M\left({}_1^1 H\right) - (A - Z)m_n \right] c^2$$

Example  ${}_1^2\text{H}$ :

$$\begin{aligned} E_b &= -\left[ M\left({}_1^2 X\right) - Z \cdot M\left({}_1^1 H\right) - (A - Z)m_n \right] c^2 \\ &= -\left[ 2.014102 - 1.007825 - 1.008665 \right] u \cdot 931.494 \text{ MeV} / u \\ &= 2.2 \text{ MeV} \end{aligned}$$

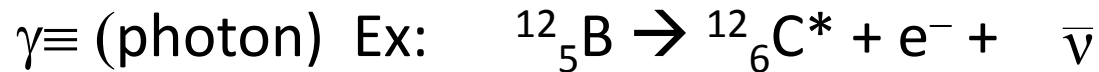
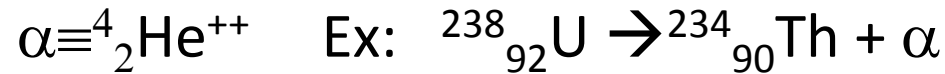
→ There must be a strong attraction between nuclear particles



$$1 \text{ fm} = 10^{-15} \text{ m}$$

Not all nuclei are stable:

Some types of nuclear decay



Measure of radioactive decays:

$$\text{Decay rate: } \frac{dN}{dt} = -\lambda N$$

$$\text{Solution: } N(t) = N_0 e^{-\lambda t}$$

$$\left| \frac{dN}{dt}(t) \right| = N_0 \lambda e^{-\lambda t}$$



Half-life:

When  $N(t) = \frac{1}{2} N_0$ :

$$N(T_{1/2}) = N_0 e^{-\lambda T_{1/2}} \equiv \frac{1}{2} N_0$$

$$\Rightarrow T_{1/2} = \frac{\ln 2}{\lambda} = \frac{0.6931472}{\lambda}$$

$$\text{Note that: } N(t) = N_0 \left(\frac{1}{2}\right)^{\left(t/T_{1/2}\right)}$$

Some units of nuclear decay:

1 Ci (Curie)  $\equiv 3.7 \times 10^{10}$  decays/s

1 Bq (Becquerel)  $\equiv 1$  decay/s

Example:

Suppose that you have a sample of  $10^{23}$   $^{14}_6\text{C}$  nuclei each of which has a half-life of 5730 years. How many Curies of radiation is this?

$$T_{1/2} = \frac{\ln 2}{\lambda} = \frac{0.6931472}{\lambda}$$

$$\lambda = \frac{0.6931472}{T_{1/2}} = 3.8 \times 10^{-12} \text{ decays/s}$$

$$\left| \frac{dN}{dt} \right| = \lambda N = 3.8 \times 10^{-12} \times 10^{23} \text{ decays/s} = 10.4 \text{ Ci}$$

Suppose you have a sample of  $6.02 \times 10^{23}$  atoms of  $^{97}\text{Tc}$  which has a half-life of  $2.6 \times 10^6$  years.

1. How many Curies does this sample have?

$$\begin{aligned}\frac{dN}{dt} &= \frac{\ln 2}{T_{1/2}} N_0 \left(\frac{1}{2}\right)^{\frac{t}{T_{1/2}}} = \frac{\ln 2}{T_{1/2}} N \\ &= \frac{\ln 2 \cdot 6.02 \times 10^{23}}{2.6 \times 10^6 \times 3.16 \times 10^7 \text{ s}} \frac{1 \text{ Ci}}{3.7 \times 10^{10} / \text{s}} \\ &= 0.137 \text{ Ci}\end{aligned}$$

2. How long will it take before the sample has only  $3.01 \times 10^{23}$   $^{97}\text{Tc}$  atoms left?

- (a)  $1.3 \times 10^6$  years
- (b)  $2.6 \times 10^6$  years
- (c)  $5.2 \times 10^6$  years
- (d)  $26 \times 10^6$  years
- (e) None of these.

## Effects of radiation in matter

Release of energetic particles –

$\alpha$ , n, p  $\rightarrow$  move atoms in materials

$e^-$ ,  $e^+$ ,  $\gamma$   $\rightarrow$  remove or excite electrons  
 $\rightarrow$  cause chemical reactions

## Quantitative measures of radiation dose

**rad** (“radiation absorbed dose”): amount of radiation that increases the energy of 1kg of absorbed materials by 0.01 J.

**gray** (Gy): 1 Gy = 100 rad

**RBE** (“relative biological effectiveness” factor): ratio of biological damage of radiation type to that of  $\gamma$  rays

**rem** (“radiation equivalent in man”): rad x RBE

**sievert** (Sv): 1 Sv = 100 rem

## Some RBE factors

<b>Radiation type</b>	<b>RBE</b>
$\gamma$ rays	1
$\beta$ particles	1-1.7
n (slow)	4-5
n & p (fast)	10
$\alpha$ particles	10-20
heavy ions	20

## Some typical values of dose

Source	Dose
Background radiation	0.13 rem/year
Recommended limit	0.5 rem/year
Diagnostic chest X-ray	0.01 rem
Mammogram	0.1 rem
Fatal dose	400-500 rem

## Summary of radiation units:

Radiation dose: Amount of radiation absorbed/unit mass

$$1 \text{ Gy} = 1 \text{ J/kg of absorbed radiation} = 100 \text{ rad}$$

$$1 \text{ Sv} = \text{RBE} \times (\text{dose in Gy units}) = 100 \text{ rem}$$

← relative biological effectiveness

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5. HRW6 43.P.069. [53088] A typical chest x-ray radiation dose is **248**  $\mu\text{Sv}$ , delivered by x rays with an RBE factor of 0.85. Assuming that the mass of the exposed tissue is one-half the patient's mass of **90** kg, calculate the energy absorbed in joules.

mJ

---

$$248 \times 10^{-6} \times 0.85 \times 45 = 9.5 \times 10^{-3} \text{ J}$$

Example:

A sealed capsule containing the radiopharmaceutical  $^{32}_{15}\text{P}$  with an initial radioactivity of  $5.22 \times 10^6$  Bq, is implanted into a 0.1 kg tumor. Each decay produces  $e^-$  particles at an energy of  $7 \times 10^5$  eV. Determine the absorbed dose in a 10 day period.  $T_{1/2} = 14.26$  days. Assume all emitted particles are absorbed.

dose = RBE x (number of decays) x (energy released)/mass

$$1 \nearrow N_0 \left( 1 - \left( \frac{1}{2} \right)^{10/14.26} \right)$$

$$N_0 = \left. \frac{dN}{dt} \right|_0 / \frac{\ln 2}{T_{1/2}} = 5.22 \times 10^6 / \frac{\ln 2}{14.26 \times 86400} = 9.2785 \times 10^{12}$$

$$\text{dose} = 4 \text{ J/kg} = 400 \text{ rem} = 4 \text{ Sv}$$

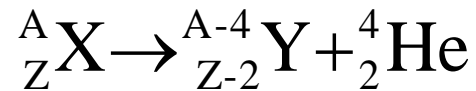


## Other facts about nuclei

- Nuclei are confined to a very small region of space – typical nuclear radii are  $10^{-15}\text{m} = \text{fm}$  (compare with atomic radius of  $10^{-10}\text{m}$ )
- For some nuclei, there are stable forms with the same  $Z$ , but different  $A$  (isotopes)
- Some nuclei are meta-stable; they transform (decay) into other forms

Decay processes :

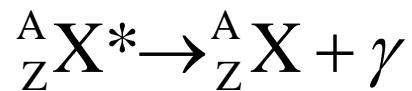
$\alpha$  decay



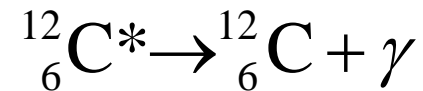
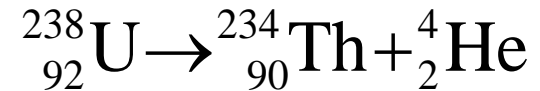
$\beta$  decay



$\gamma$  decay



Examples :



## Summary of some ideas about nuclear physics

- The basic forces that describe nuclei are stronger, shorter range, and more complicated than electromagnetic forces. Nuclear binding energies are typically 8 MeV/nucleon.

- Most nuclear decay processes are described by a simple rate equation

$$\frac{dN}{dt} = -\lambda N \Rightarrow N(t) = N_0 e^{-\lambda t} = N_0 \left(\frac{1}{2}\right)^{t/T_{1/2}}$$

For example:  $^{238}\text{U} \rightarrow ^{234}\text{Th} + ^4\text{He}$       $T_{1/2} = 4.5 \times 10^9 \text{ yr.}$

- Radiation effects on humans and other sensitive forms of life and devices are due to the light decay products –  $\alpha, \beta, \gamma$  particles

Radiation dose: Amount of radiation absorbed/unit mass

1 Gy = 1 J/kg of absorbed radiation = 100 rad

1 Sv = RBE  $\times$  (dose in Gy units) = 100 rem

 relative biological effectiveness

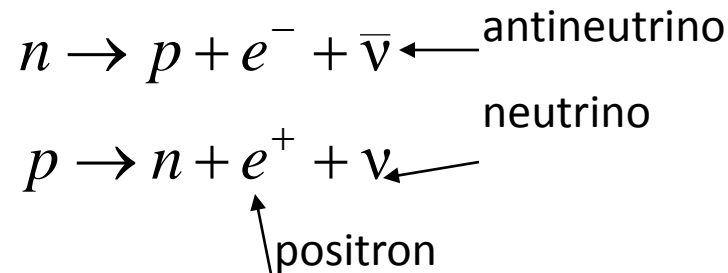
# Guide to nuclear reactions

## General rules

Total  $A$  (number of nucleons ( $Z+N$ )) is conserved

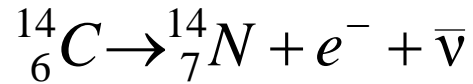
Total charge is conserved

protons and neutrons can convert to each other



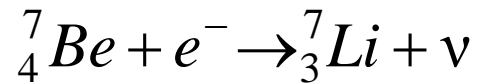
Neutrinos were first detected in 1956 by Fred Reines and George Cowan, who showed that a nucleus undergoing beta decay emits a neutrino with the electron. Neutrinos are VERY weakly interacting and recent evidence suggests they have a mass of  $< 0.1$  eV.

## Examples of nuclear reactions

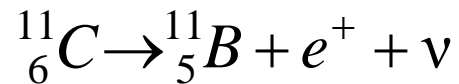


Used for carbon dating

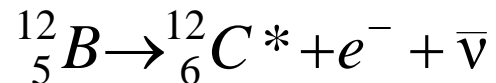
$$T_{1/2}({}^{14}_6\text{C}) = 5730 \text{ yrs.}$$



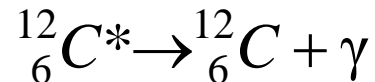
Nucleus captures an electron



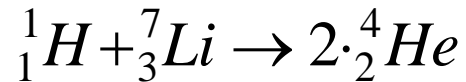
Positron emission



Beta decay followed by gamma decay



## More nuclear reactions



Energy accounting :

for a general reaction of the type  $a+X \rightarrow Y+b$

energy released as kinetic energy of the products  
can be calculated:

$$Q = (M_a + M_X - M_Y - M_\beta) c^2$$

$$\begin{aligned} \text{In this case, } Q &= (1.007825 + 7.016003 - 2 \cdot 4.002602) c^2 \\ &= 17.348 \text{ MeV} \end{aligned}$$

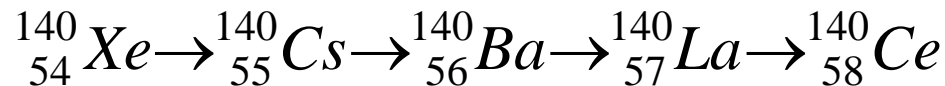
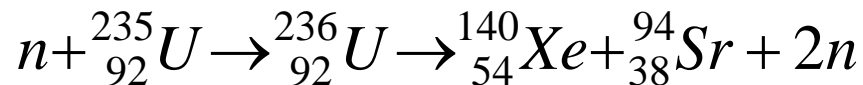
Fission: Some history –

1932 James Chadwick (England) discovered neutron

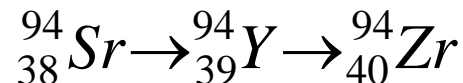
Enrico Fermi (Italy) discovered that neutrons  
could be absorbed by nuclei to form new  
elements

Lise Meitner, Otto Hahn, Fritz Strassmann, Otto  
Frisch (Germany) discovered fission of U

Example:



(14s) (64s) (13d) (40h) (stable)



(75s)(19min)(stable)

$Q \approx 200 \text{ MeV}$

Other decay products for

