Electrodynamics – PHY712

Lecture 12 – magnetostatic examples

Reference: Chap. 5 in J. D. Jackson's textbook.

Calculation of the vector potential for a confined current density

If the current density J(r) is confined in space, the vector potential in the Coulomb gauge can be calculated from

$$\mathbf{A}(\mathbf{r}) = \frac{\mu_0}{4\pi} \int d^3r' \frac{\mathbf{J}(\mathbf{r}')}{|\mathbf{r} - \mathbf{r}'|}.$$
 (1)



Simple example of current density from a rotating charged sphere

Consider the following example corresponding to a rotating charged sphere of radius a, with ρ_0 denoting the uniform charge density within the sphere and ω denoting the angular rotation of the sphere:

$$\mathbf{J}(\mathbf{r}') = \begin{cases} \rho_0 \boldsymbol{\omega} \times \mathbf{r}' & \text{for } r' \leq a \\ 0 & \text{otherwise} \end{cases}$$
 (2)

In order to evaluate the vector potential (1) for this problem, we can make use of the expansion:

$$\frac{1}{|\mathbf{r} - \mathbf{r}'|} = \sum_{lm} \frac{4\pi}{2l+1} \frac{r_{<}^l}{r_{>}^{l+1}} Y_{lm}(\hat{\mathbf{r}}) Y_{lm}^*(\hat{\mathbf{r}}'). \tag{3}$$

Noting that

$$\mathbf{r}' = r' \sqrt{\frac{4\pi}{3}} \left(Y_{1-1}(\hat{\mathbf{r}'}) \frac{\hat{\mathbf{x}} + \mathbf{i}\hat{\mathbf{y}}}{\sqrt{2}} + Y_{11}(\hat{\mathbf{r}'}) \frac{-\hat{\mathbf{x}} + \mathbf{i}\hat{\mathbf{y}}}{\sqrt{2}} + Y_{10}(\hat{\mathbf{r}'}) \hat{\mathbf{z}} \right), \tag{4}$$

we see that the angular integral in Eq. (1) can be simplified with the use of the identity:

$$\int d\Omega' \sum_{m} Y_{lm}(\hat{\mathbf{r}}) Y_{lm}^*(\hat{\mathbf{r}'}) \mathbf{r}' = \frac{r'}{r} \mathbf{r} \ \delta_{l1}.$$
 (5)



Simple example of current density from a rotating charged sphere – continued

Therefore the vector potential for this system is:

$$\mathbf{A}(\mathbf{r}) = \frac{\mu_0 \rho_0 \boldsymbol{\omega} \times \mathbf{r}}{3r} \int_0^a dr' \ r'^3 \frac{r_{<}}{r_{>}^2},\tag{6}$$

which can be evaluated as:

$$\mathbf{A}(\mathbf{r}) = \begin{cases} \frac{\mu_0 \rho_0}{3} \boldsymbol{\omega} \times \mathbf{r} \left(\frac{a^2}{2} - \frac{3r^2}{10} \right) & \text{for } r \leq a \\ \frac{\mu_0 \rho_0}{3} \boldsymbol{\omega} \times \mathbf{r} \frac{a^5}{5r^3} & \text{for } r \geq a \end{cases}$$
(7)

$$\mathbf{B}(\mathbf{r}) = \nabla \times \mathbf{A}(\mathbf{r}) = \begin{cases} \frac{\mu_0 \rho_0}{3} \left[\boldsymbol{\omega} \left(a^2 - \frac{6}{5} r^2 \right) + \frac{3}{5} \mathbf{r} (\boldsymbol{\omega} \cdot \mathbf{r}) \right] & \text{for } r \leq a \\ \frac{\mu_0 \rho_0}{3} \left[-\boldsymbol{\omega} \frac{a^5}{5r^3} + \frac{3a^5}{5r^5} \mathbf{r} (\boldsymbol{\omega} \cdot \mathbf{r}) \right] & \text{for } r \geq a \end{cases}$$
(8)



Another example – current associated with an electron in a spherical atom

In this case, we assume that the current density is due to an electron in a bound atomic state with quantum numbers $|nlm_l\rangle$, as described by a wavefunction $\psi_{nlm_l}(\mathbf{r})$, where the azimuthal quantum number m_l is associated with a factor of the form $e^{im_l\phi}$. For such a wavefunction the quantum mechanical current density operator can be evaluated:

$$\mathbf{J}(\mathbf{r}) = \frac{-e\hbar}{2m_e i} \left(\psi_{nlm_l}^* \nabla \psi_{nlm_l} - \psi_{nlm_l} \nabla \psi_{nlm_l}^* \right). \tag{9}$$

Since the only complex part of this wavefunction is associated with the azimuthal quantum number, this can be written:

$$\mathbf{J}(\mathbf{r}) = \frac{-e\hbar}{2m_e i r \sin \theta} \left(\psi_{nlm_l}^* \frac{\partial}{\partial \phi} \psi_{nlm_l} - \psi_{nlm_l} \frac{\partial}{\partial \phi} \psi_{nlm_l}^* \right) \hat{\phi} = \frac{-e\hbar m_l \hat{\phi}}{m_e r \sin \theta} \left| \psi_{nlm_l} \right|^2.$$
(10)

where m_e denotes the electron mass and e denotes the magnitude of the electron charge.



$$\mathbf{J}(\mathbf{r}) = \frac{-e\hbar m_l \hat{\boldsymbol{\phi}}}{m_e r \sin \theta} |\psi_{nlm_l}(\mathbf{r})|^2 = \frac{-e\hbar m_l \hat{\mathbf{z}} \times \mathbf{r}}{m_e r^2 \sin^2 \theta} |\psi_{nlm_l}(\mathbf{r})|^2.$$
(11)

$$\mathbf{A}(\mathbf{r}) = \frac{\mu_0}{4\pi} \left(\frac{-e\hbar m_l}{m_e} \right) \int d^3r' \frac{\hat{\mathbf{z}} \times \mathbf{r}'}{|\mathbf{r} - \mathbf{r}'|} \frac{|\psi_{nlm_l}(\mathbf{r}')|^2}{r'^2 \sin^2 \theta'}.$$
 (12)

Note that for some atomic wavefunctions, $\psi_{nlm_l}(\mathbf{r}')$, the evaluation of the vector potential $\mathbf{A}(\mathbf{r})$ simplifies.



For example, consider the $|nlm = 211\rangle$ state of a H atom:

$$\psi_{211}(\mathbf{r}) = -\sqrt{\frac{1}{64\pi a^3}} \frac{r}{a} e^{-r/(2a)} \sin \theta e^{i\phi}, \tag{13}$$

and

$$\mathbf{J}(\mathbf{r}') = \frac{-e\hbar}{64m_e\pi a^5} e^{-r'/a} \,\,\hat{\mathbf{z}} \times \mathbf{r}',\tag{14}$$

where a here denotes the Bohr radius. Using arguments similar to those above, we find that

$$\mathbf{A}(\mathbf{r}) = \frac{-e\hbar\mu_0\hat{\mathbf{z}} \times \mathbf{r}}{192m_e\pi a^5 r} \int_0^\infty dr' \ r'^3 \ e^{-r'/a} \ \frac{r_{<}}{r_{>}^2}.$$
 (15)

This expression can be integrated to give:

$$\mathbf{A}(\mathbf{r}) = \frac{-e\hbar\mu_0\hat{\mathbf{z}} \times \mathbf{r}}{8m_e\pi r^3} \left[1 - e^{-r/a} \left(1 + \frac{r}{a} + \frac{r^2}{2a^2} + \frac{r^3}{8a^3} \right) \right].$$
 (16)



Previous result:

$$\mathbf{A}(\mathbf{r}) = \frac{-e\hbar\mu_0\hat{\mathbf{z}} \times \mathbf{r}}{8m_e\pi r^3} \left[1 - e^{-r/a} \left(1 + \frac{r}{a} + \frac{r^2}{2a^2} + \frac{r^3}{8a^3} \right) \right].$$
 (17)

Note that for $r \to \infty$:

$$\mathbf{A}(\mathbf{r}) = \frac{-e\hbar\mu_0\hat{\mathbf{z}} \times \mathbf{r}}{8m_e\pi r^3} = \frac{\mu_0}{4\pi} \left(-\frac{e\hbar}{2m_e} \right) \frac{\hat{\mathbf{z}} \times \mathbf{r}}{r^3} = \frac{\mu_0}{4\pi} \frac{\mathbf{m} \times \mathbf{r}}{r^3},\tag{18}$$

where

$$\mathbf{m} = \left(-\frac{e\hbar}{2m_e}\right)\hat{\mathbf{z}}.\tag{19}$$

More generally:

$$\mathbf{m} = \frac{1}{2} \int d^3 r' \ \mathbf{r}' \times \mathbf{J}(\mathbf{r}'). \tag{20}$$



Note that the general form of the current density for a spherical atom is given by:

$$\mathbf{J}(\mathbf{r}) = \frac{-e\hbar m_l \hat{\phi}}{m_e r \sin \theta} |\psi_{nlm_l}(\mathbf{r})|^2 = \frac{-e\hbar m_l}{m_e} \frac{\hat{\mathbf{z}} \times \hat{\mathbf{r}}}{r \sin^2 \theta} |\psi_{nlm_l}(\mathbf{r})|^2.$$
(21)

In this case, the general form of the magnetic dipole moment is given by

$$\mathbf{m} = \frac{1}{2} \int d^3 r' \ \mathbf{r}' \times \mathbf{J}(\mathbf{r}') = -\frac{e\hbar m_l}{2m_e} \hat{\mathbf{z}} \int d^3 r' \left| \psi_{nlm_l}(\mathbf{r}') \right|^2 = -\frac{e\hbar}{2m_e} m_l \hat{\mathbf{z}}.$$
 (22)



Systematic multipole analysis of vector potential for a general confined current density $\mathbf{J}(\mathbf{r})$ (assuming $\nabla \cdot \mathbf{J}(\mathbf{r}) = 0$.

$$\mathbf{A}(\mathbf{r}) = \frac{\mu_0}{4\pi} \int d^3 r' \frac{\mathbf{J}(\mathbf{r}')}{|\mathbf{r} - \mathbf{r}'|}.$$
 (23)

For field point r outside of extent of current density:

$$\frac{1}{|\mathbf{r} - \mathbf{r}'|} = \frac{1}{r} + \frac{\mathbf{r} \cdot \mathbf{r}'}{r^3} \cdots . \tag{24}$$

$$\mathbf{A}(\mathbf{r}) \approx \frac{\mu_0}{4\pi} \left(\frac{1}{r} \int d^3 r' \mathbf{J}(\mathbf{r}') + \frac{\mathbf{r}}{r^3} \cdot \int d^3 r' \mathbf{r}' \mathbf{J}(\mathbf{r}') \dots \right)$$
(25)

Note that

$$\int d^3r' \mathbf{J}(\mathbf{r'}) = 0 \tag{26}$$

$$\mathbf{r} \cdot \int d^3 r' \mathbf{r'} \mathbf{J}(\mathbf{r'}) = -\frac{1}{2} \mathbf{r} \times \int d^3 r' \mathbf{r'} \times \mathbf{J}(\mathbf{r'}) \equiv \mathbf{m} \times \mathbf{r}.$$
 (27)



Magnetic dipolar field

The magnetic dipole moment is defined by

$$\mathbf{m} = \frac{1}{2} \int d^3 r' \mathbf{r'} \times \mathbf{J}(\mathbf{r'}), \tag{28}$$

with the corresponding potential

$$\mathbf{A}(\mathbf{r}) = \frac{\mu_0}{4\pi} \frac{\mathbf{m} \times \hat{\mathbf{r}}}{r^2},\tag{29}$$

and magnetostatic field

$$\mathbf{B}_{\mathbf{m}}(\mathbf{r}) = \frac{\mu_0}{4\pi} \left\{ \frac{3\hat{\mathbf{r}}(\mathbf{m} \cdot \hat{\mathbf{r}}) - \mathbf{m}}{r^3} + \frac{8\pi}{3} \mathbf{m} \delta^3(\mathbf{r}) \right\}. \tag{30}$$



Magnetic dipolar field – continued

Some details:

$$\nabla \times (s\mathbf{V}) = \nabla s \times \mathbf{V} + s\nabla \times \mathbf{V}. \tag{31}$$

$$\nabla \times (\mathbf{V}_1 \times \mathbf{V}_2) = \mathbf{V}_1(\nabla \cdot \mathbf{V}_2) - \mathbf{V}_2(\nabla \cdot \mathbf{V}_1) + (\mathbf{V}_2 \cdot \nabla)\mathbf{V}_1 - (\mathbf{V}_1 \cdot \nabla)\mathbf{V}_2. \quad (32)$$

For r > 0:

$$\nabla \times \left(\frac{\mathbf{m} \times \mathbf{r}}{r^3}\right) = \frac{3\mathbf{r}(\mathbf{m} \cdot \mathbf{r}) - r^2\mathbf{m}}{r^5}.$$
 (33)



Justification for the δ function contribution at the origin of the magnetic dipole

Note: This derivation is very similar to the analogous electrostatic case.

The evaluation of the field at the origin of the dipole is poorly defined, but we make the following approximation.

$$\mathbf{B}(\mathbf{r} \approx \mathbf{0}) \approx \left(\int_{\text{sphere}} \mathbf{B}(\mathbf{r}) \mathbf{d}^3 \mathbf{r} \right) \delta^3(\mathbf{r}).$$
 (34)

First we note that

$$\int_{r \le R} \mathbf{B}(\mathbf{r}) d^3 r = R^2 \int_{r=R} \hat{\mathbf{r}} \times \mathbf{A}(\mathbf{r}) \ d\Omega.$$
 (35)

This result follows from the divergence theorm:

$$\int_{\text{vol}} \nabla \cdot \mathcal{V} \mathbf{d^3 r} = \int_{\text{surface}} \mathcal{V} \cdot \mathbf{dA}.$$
 (36)



Singular contribution to dipolar field – continued

The divergence theorem can be used to prove Eq. (35) for each cartesian coordinate of $\nabla \times \mathbf{A}$ since $\nabla \times \mathbf{A} = \hat{\mathbf{x}} (\hat{\mathbf{x}} \cdot (\nabla \times \mathbf{A})) + \hat{\mathbf{y}} (\hat{\mathbf{y}} \cdot (\nabla \times \mathbf{A})) + \hat{\mathbf{z}} (\hat{\mathbf{z}} \cdot (\nabla \times \mathbf{A}))$. Note that $\hat{\mathbf{x}} \cdot (\nabla \times \mathbf{A}) = -\nabla \cdot (\hat{\mathbf{x}} \times \mathbf{A})$ and that we can use the Divergence theorem with $\mathcal{V} \equiv \hat{\mathbf{x}} \times \mathbf{A}(\mathbf{r})$ for the x- component for example:

$$\int_{\text{vol}} \nabla \cdot (\hat{\mathbf{x}} \times \mathbf{A}) d^3 r = \int_{\text{surface}} (\hat{\mathbf{x}} \times \mathbf{A}) \cdot \hat{\mathbf{r}} dA = \int_{\text{surface}} (\mathbf{A} \times \hat{\mathbf{r}}) \cdot \hat{\mathbf{x}} dA.$$
 (37)

Therefore,

$$\int_{r \le R} (\nabla \times \mathbf{A}) d^3 r = -\int_{r=R} (\mathbf{A} \times \hat{\mathbf{r}}) \cdot (\hat{\mathbf{x}}\hat{\mathbf{x}} + \hat{\mathbf{y}}\hat{\mathbf{y}} + \hat{\mathbf{z}}\hat{\mathbf{z}}) dA = R^2 \int_{r=R} (\hat{\mathbf{r}} \times \mathbf{A}) d\Omega$$
(38)

which is identical to Eq. (35). We can use the identity (as in electrostatic case),

$$\int d\Omega \frac{\hat{\mathbf{r}}}{|\mathbf{r} - \mathbf{r}'|} = \frac{4\pi}{3} \frac{r_{<}}{r_{>}^{2}} \hat{\mathbf{r}'}.$$
 (39)



Singular contribution to dipolar field – continued

Now, expressing the vector potential in terms of the current density:

$$\mathbf{A}(\mathbf{r}) = \frac{\mu_0}{4\pi} \int d^3r \frac{\mathbf{J}(\mathbf{r}')}{|\mathbf{r} - \mathbf{r}'|},\tag{40}$$

the integral over Ω in Eq. 35 becomes

$$R^{2} \int_{r=R} (\hat{\mathbf{r}} \times \mathbf{A}) d\Omega = \frac{4\pi R^{2}}{3} \frac{\mu_{0}}{4\pi} \int d^{3}r' \frac{r_{<}}{r_{>}^{2}} \hat{\mathbf{r}'} \times \mathbf{J}(\mathbf{r}'). \tag{41}$$

If the sphere R contains the entire current distribution, then $r_>=R$ and $r_<=r'$ so that (41) becomes

$$R^{2} \int_{r=R} (\hat{\mathbf{r}} \times \mathbf{A}) d\Omega = \frac{4\pi}{3} \frac{\mu_{0}}{4\pi} \int d^{3}r' \ \mathbf{r}' \times \mathbf{J}(\mathbf{r}') \equiv \frac{8\pi}{3} \frac{\mu_{0}}{4\pi} \mathbf{m}, \tag{42}$$

which thus justifies the delta-function contribution in Eq. 30 and results so-called "Fermi contact" contribution in the "hyperfine" interaction.



Magnetic field due to electrons in the vicinity of a nucleus

Contribution due to "orbital" magnetism in a spherical atom

The current density associated with an electron in a bound state of an atom as described by a quantum mechanical wavefunction $\psi_{nlm_l}(\mathbf{r})$ can be written:

$$\mathbf{J}(\mathbf{r}) = \frac{-e\hbar m_l \hat{\phi}}{m_e r \sin \theta} \left| \psi_{nlm_l}(\mathbf{r}) \right|^2. \tag{43}$$

In the following, it will be convenient to represent the azimuthal unit vector $\hat{\phi}$ in terms of cartesian coordinates:

$$\hat{\phi} = -\sin\phi\hat{\mathbf{x}} + \cos\phi\hat{\mathbf{y}} = \frac{\hat{\mathbf{z}} \times \mathbf{r}}{r\sin\theta}.$$
 (44)

The vector potential for this current density can be written

$$\mathbf{A}(\mathbf{r}) = -\frac{\mu_0}{4\pi} \frac{e\hbar}{m_e} m_l \int d^3 r' \frac{\hat{\mathbf{z}} \times \mathbf{r'}}{|\mathbf{r} - \mathbf{r'}|} \frac{|\psi_{nlm_l}(\mathbf{r'})|^2}{r'^2 \sin^2 \theta'}$$
(45)



Contribution due to "orbital" magnetism in a spherical atom – continued

We want to evaluate the magnetic field $B = \nabla \times A$ in the vicinity of the nucleus $(\mathbf{r} \to 0)$. Taking the curl of the Eq. 45, we obtain

$$\mathbf{B_o}(\mathbf{r}) = \frac{\mu_0}{4\pi} \frac{e\hbar}{m_e} m_l \int d^3r' \frac{(\mathbf{r} - \mathbf{r}') \times (\hat{\mathbf{z}} \times \mathbf{r}')}{|\mathbf{r} - \mathbf{r}'|^3} \frac{|\psi_{nlm_l}(\mathbf{r}')|^2}{r'^2 \sin^2 \theta'}$$
(46)

Evaluating this expression with $(\mathbf{r} \to 0)$, we obtain

$$\mathbf{B_o(0)} = -\frac{\mu_0}{4\pi} \frac{e\hbar}{m_e} m_l \int d^3r' \frac{\mathbf{r'} \times (\hat{\mathbf{z}} \times \mathbf{r'})}{r'^3} \frac{|\psi_{nlm_l}(\mathbf{r'})|^2}{r'^2 \sin^2 \theta'}$$
(47)



Contribution due to "orbital" magnetism in a spherical atom – continued

$$\mathbf{B_o(0)} = -\frac{\mu_0}{4\pi} \frac{e\hbar}{m_e} m_l \int d^3r' \frac{\mathbf{r'} \times (\hat{\mathbf{z}} \times \mathbf{r'})}{r'^3} \frac{|\psi_{nlm_l}(\mathbf{r'})|^2}{r'^2 \sin^2 \theta'}$$
(48)

Expanding the cross product and expressing the result in spherical polar coordinates, we obtain in the numerator

$$\mathbf{\hat{r}}' \times (\mathbf{\hat{z}} \times \mathbf{\hat{r}}') = \mathbf{\hat{z}}(\mathbf{1} - \cos^2 \theta') - \mathbf{\hat{x}} \cos \theta' \sin \theta' \cos \phi' - \mathbf{\hat{y}} \cos \theta' \sin \theta' \sin \phi').$$

In evaluating the integration over the azimuthal variable ϕ' , the $\hat{\mathbf{x}}$ and $\hat{\mathbf{y}}$ components vanish which reduces to

$$\mathbf{B_{o}}(\mathbf{0}) = -\frac{\mu_0}{4\pi} \frac{e\hbar}{m_e} m_l \int d^3 r' \frac{\hat{\mathbf{z}} r'^2 \sin^2 \theta'}{r'^3} \frac{|\psi_{nlm_l}(\mathbf{r}')|^2}{r'^2 \sin^2 \theta'}$$
(49)

and

$$\mathbf{B_o(0)} = -\frac{\mu_0 e \hbar m_l \hat{\mathbf{z}}}{4\pi m_e} \int d^3 r' \left| \psi_{nlm_l} \right|^2 \frac{1}{r'^3} \equiv -\frac{\mu_0 e}{4\pi m_e} L_z \hat{\mathbf{z}} \left\langle \frac{1}{r'^3} \right\rangle. \tag{50}$$



"Hyperfine" interaction

The so-called "hyperfine" interaction results from the magnetic dipole moment of a nucleus $\mu_{\mathbf{N}}$ responding to the magnetic field formed by the magnetic dipole of the electron spin ($\mu_{\mathbf{e}}$) as well as the electron orbital current contribution.

$$\mathcal{H}_{HF} = -\mu_{\mathbf{N}} \cdot (\mathbf{B}_{\mu_e} + \mathbf{B}_o(0)). \tag{51}$$

$$\mathcal{H}_{HF} = -\frac{\mu_0}{4\pi} \left(\frac{3(\mu_{\mathbf{N}} \cdot \hat{\mathbf{r}})(\mu_{\mathbf{e}} \cdot \hat{\mathbf{r}}) - \mu_{\mathbf{N}} \cdot \mu_{\mathbf{e}}}{r^3} + \frac{8\pi}{3} \mu_{\mathbf{N}} \cdot \mu_{\mathbf{e}} \delta^3(\mathbf{r}) + \frac{e}{m_e} \left\langle \frac{\mathbf{L} \cdot \mu_{\mathbf{N}}}{r^3} \right\rangle \right). \tag{52}$$

