

**PHY 712 Electrodynamics
9:50 AM MWF Olin 105**

Plan for Lecture 27:
Continue reading Chap. 14 –
Radiation by moving charges

- 1. Motion in a line**
- 2. Motion in a circle**
- 3. Spectral analysis of radiation**

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WFU Physics

Events

Colloquium: "Emerson Corporation-A Global Manufacturer of Industrial and Residential Products" - Wednesday, April 3, 2019 at 4:00 PM
Randall D. Ledford, PhD Wake Forest Alumni
Retired CTO of Emerson Electric Company
George P. Williams, Jr. Lecture Hall (Olin 101)
Wednesday, April 3, 2019, at 4:00 PM There
will be a reception following the talk.

News

Roy Thompson presents with science museum to create exhibit

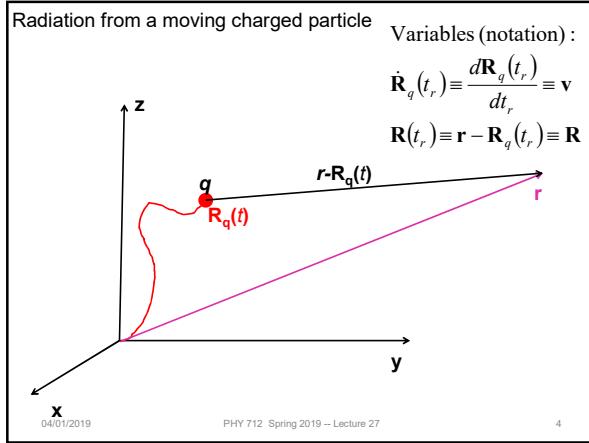
Randall Ledford receives Physics Department Outstanding Alumni Award

Nationally recognized for teaching excellence; internationally respected for research

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Fri: 03/15/2019	No class	Spring Break		
21 Mon: 03/18/2019	Chap. 9	Radiation from localized oscillating sources	#15	3/22/2019
22 Wed: 03/20/2019	Chap. 9	Radiation from oscillating sources	#16	3/25/2019
23 Fri: 03/22/2019	Chap. 9 and 10	Radiation from oscillating sources	#17	3/27/2019
24 Mon: 03/25/2019	Chap. 11	Special Theory of Relativity	Pick topic	3/29/2019
25 Wed: 03/27/2019	Chap. 11	Special Theory of Relativity	#18	4/01/2019
26 Fri: 03/29/2019	Chap. 11	Special Theory of Relativity	#19	4/03/2019
27 Mon: 04/01/2019	Chap. 14	Radiation from accelerating charged particles	#20	4/05/2019
28 Wed: 04/03/2019				
29 Fri: 04/05/2019				
30 Mon: 04/08/2019				
31 Wed: 04/10/2019				
32 Fri: 04/12/2019				
33 Mon: 04/15/2019				
34 Wed: 04/17/2019				
35 Fri: 04/19/2019	No class	Good Friday		
36 Mon: 04/22/2019				
37 Wed: 04/24/2019				
38 Fri: 04/26/2019		Presentations I		
Mon: 04/29/2019		Presentations II		
Wed: 05/01/2019		Presentations III		

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Liénard-Wiechert fields (cgs Gaussian units):

$$\mathbf{E}(\mathbf{r}, t) = \frac{q}{(R - \frac{\mathbf{v} \cdot \mathbf{R}}{c})^3} \left[\left(\mathbf{R} - \frac{\mathbf{v} R}{c} \right) \left(1 - \frac{v^2}{c^2} \right) + \left(\mathbf{R} \times \left\{ \left(\mathbf{R} - \frac{\mathbf{v} R}{c} \right) \times \frac{\dot{\mathbf{v}}}{c^2} \right\} \right) \right]. \quad (19)$$

$$\mathbf{B}(\mathbf{r}, t) = \frac{q}{c} \left[\frac{-\mathbf{R} \times \mathbf{v}}{(R - \frac{\mathbf{v} \cdot \mathbf{R}}{c})^3} \left(1 - \frac{v^2}{c^2} + \frac{\dot{\mathbf{v}} \cdot \mathbf{R}}{c^2} \right) - \frac{\mathbf{R} \times \dot{\mathbf{v}}/c}{(R - \frac{\mathbf{v} \cdot \mathbf{R}}{c})^2} \right]. \quad (20)$$

In this case, the electric and magnetic fields are related according to

$$\mathbf{B}(\mathbf{r}, t) = \frac{\mathbf{R} \times \mathbf{E}(\mathbf{r}, t)}{R}. \quad (21)$$

$$\dot{\mathbf{R}}_q(t_r) \equiv \frac{d\mathbf{R}_q(t_r)}{dt_r} \equiv \mathbf{v} \quad \mathbf{R}(t_r) \equiv \mathbf{r} - \mathbf{R}_q(t_r) \equiv \mathbf{R} \quad \dot{\mathbf{v}} \equiv \frac{d^2\mathbf{R}_q(t_r)}{dt_r^2}$$

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Electric field far from source:

$$\mathbf{E}(\mathbf{r}, t) = \frac{q}{\left(R - \frac{\mathbf{v} \cdot \mathbf{R}}{c} \right)^3} \left\{ \mathbf{R} \times \left[\left(\mathbf{R} - \frac{\mathbf{v} R}{c} \right) \times \frac{\dot{\mathbf{v}}}{c^2} \right] \right\}$$

$$\mathbf{B}(\mathbf{r}, t) = \frac{\mathbf{R} \times \mathbf{E}(\mathbf{r}, t)}{R}$$

Let $\hat{\mathbf{R}} \equiv \frac{\mathbf{R}}{R}$ $\beta \equiv \frac{\mathbf{v}}{c}$ $\dot{\beta} \equiv \frac{\dot{\mathbf{v}}}{c}$

$$\mathbf{E}(\mathbf{r}, t) = \frac{q}{c R \left(1 - \beta \cdot \hat{\mathbf{R}} \right)^3} \left\{ \hat{\mathbf{R}} \times [(\hat{\mathbf{R}} - \beta) \times \dot{\beta}] \right\}$$

$$\mathbf{B}(\mathbf{r}, t) = \hat{\mathbf{R}} \times \mathbf{E}(\mathbf{r}, t)$$

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Poynting vector:

$$\mathbf{S}(\mathbf{r}, t) = \frac{c}{4\pi} (\mathbf{E} \times \mathbf{B})$$

$$\mathbf{E}(\mathbf{r}, t) = \frac{q}{cR(1 - \beta \cdot \hat{\mathbf{R}})^3} \left\{ \hat{\mathbf{R}} \times [(\hat{\mathbf{R}} - \beta) \times \dot{\beta}] \right\}$$

$$\mathbf{B}(\mathbf{r}, t) = \hat{\mathbf{R}} \times \mathbf{E}(\mathbf{r}, t)$$

$$\mathbf{S}(\mathbf{r}, t) = \frac{c}{4\pi} \hat{\mathbf{R}} |\mathbf{E}(\mathbf{r}, t)|^2 = \frac{q^2}{4\pi c R^2} \hat{\mathbf{R}} \frac{|\hat{\mathbf{R}} \times [(\hat{\mathbf{R}} - \beta) \times \dot{\beta}]|^2}{(1 - \beta \cdot \hat{\mathbf{R}})^6}$$

Note: We have used the fact that

$$\hat{\mathbf{R}} \cdot \mathbf{E}(\mathbf{r}, t) = 0$$

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Power radiated

$$\mathbf{S}(\mathbf{r}, t) = \frac{c}{4\pi} \hat{\mathbf{R}} |\mathbf{E}(\mathbf{r}, t)|^2 = \frac{q^2}{4\pi c R^2} \hat{\mathbf{R}} \frac{|\hat{\mathbf{R}} \times [(\hat{\mathbf{R}} - \beta) \times \dot{\beta}]|^2}{(1 - \beta \cdot \hat{\mathbf{R}})^6}$$

$$\frac{dP}{d\Omega} = \mathbf{S} \cdot \hat{\mathbf{R}} R^2 = \frac{q^2}{4\pi c} \frac{|\hat{\mathbf{R}} \times [(\hat{\mathbf{R}} - \beta) \times \dot{\beta}]|^2}{(1 - \beta \cdot \hat{\mathbf{R}})^6}$$

In the non-relativistic limit: $\beta \ll 1$

$$\frac{dP}{d\Omega} = \frac{q^2}{4\pi c} |\hat{\mathbf{R}} \times [\hat{\mathbf{R}} \times \dot{\beta}]|^2 = \frac{q^2}{4\pi c^3} |\dot{\mathbf{v}}|^2 \sin^2 \Theta$$

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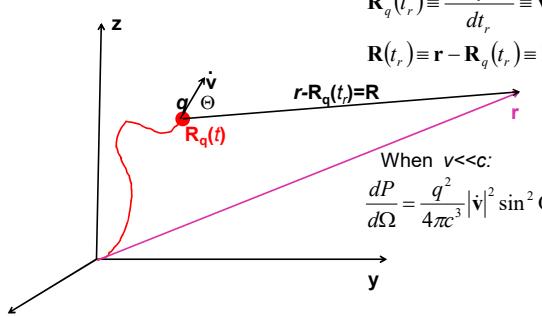
Radiation from a moving charged particle

Variables (notation) :

$$\dot{\mathbf{R}}_q(t_r) \equiv \frac{d\mathbf{R}_q(t_r)}{dt_r} \equiv \mathbf{v}$$

$$\mathbf{R}(t_r) \equiv \mathbf{r} - \mathbf{R}_q(t_r) \equiv \mathbf{R}$$

$$\text{When } v \ll c: \quad \frac{dP}{d\Omega} = \frac{q^2}{4\pi c^3} |\dot{\mathbf{v}}|^2 \sin^2 \Theta$$



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Radiation power in non-relativistic case -- continued

$$\frac{dP}{d\Omega} = \frac{q^2}{4\pi c^3} |\dot{\mathbf{v}}|^2 \sin^2 \Theta$$

$$P = \int d\Omega \frac{dP}{d\Omega} = \frac{2}{3} \frac{q^2}{c^3} |\dot{\mathbf{v}}|^2$$

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Radiation distribution in the relativistic case

$$\frac{dP}{d\Omega} = \mathbf{S} \cdot \hat{\mathbf{R}} R^2 = \frac{q^2}{4\pi c} \left| \hat{\mathbf{R}} \times \left[(\hat{\mathbf{R}} - \beta) \times \dot{\beta} \right] \right|^2 \Bigg|_{t_r=t-R/c}$$

This expression gives us the energy per unit field time t . We are often interested in the power per unit retarded time $t_r = t - R/c$:

$$\frac{dP(t)}{d\Omega} = \frac{dP_r(t_r)}{d\Omega} \frac{dt}{dt_r} = \frac{dt}{dt_r} = 1 - \beta \cdot \hat{\mathbf{R}}$$

$$\frac{dP_r(t_r)}{d\Omega} = \frac{q^2}{4\pi c} \left| \hat{\mathbf{R}} \times \left[(\hat{\mathbf{R}} - \beta) \times \dot{\beta} \right] \right|^2 \Bigg|_{t_r=t-R/c}$$

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Radiation distribution in the relativistic case -- continued

$$\frac{dP_r(t_r)}{d\Omega} = \frac{q^2}{4\pi c} \left| \hat{\mathbf{R}} \times \left[(\hat{\mathbf{R}} - \beta) \times \dot{\beta} \right] \right|^2 \Bigg|_{t_r=t-R/c}$$

For linear acceleration: $\beta \times \dot{\beta} = 0$

$$\frac{dP_r(t_r)}{d\Omega} = \frac{q^2}{4\pi c} \left| \hat{\mathbf{R}} \times (\hat{\mathbf{R}} \times \dot{\beta}) \right|^2 \Bigg|_{t_r=t-R/c} = \frac{q^2}{4\pi c^3} |\dot{\mathbf{v}}|^2 \frac{\sin^2 \theta}{(1 - \beta \cos \theta)^5}$$

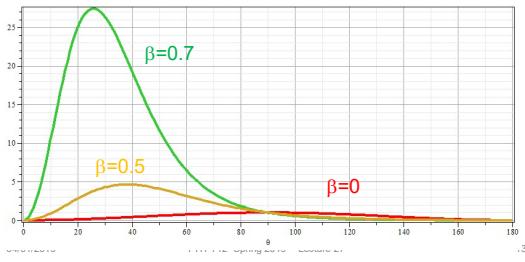
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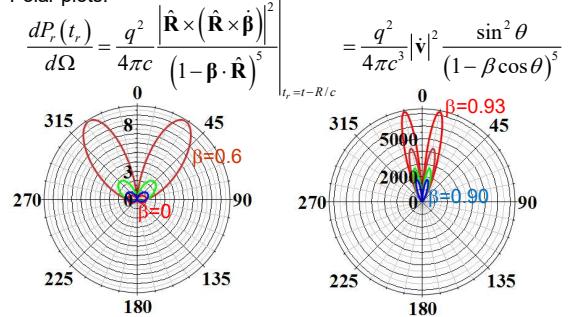
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Power from linearly accelerating particle

$$\frac{dP_r(t_r)}{d\Omega} = \frac{q^2}{4\pi c} \left| \hat{\mathbf{R}} \times (\hat{\mathbf{R}} \times \hat{\mathbf{p}}) \right|^2 \Bigg|_{t_r=t-R/c} = \frac{q^2}{4\pi c^3} |\dot{\mathbf{v}}|^2 \frac{\sin^2 \theta}{(1-\beta \cos \theta)^5}$$



Polar plots:



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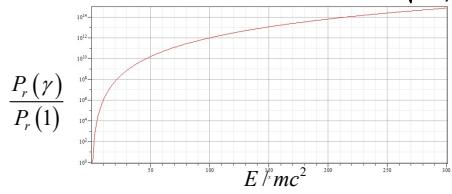
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Power from linearly accelerating particle

$$\frac{dP_r(t_r)}{d\Omega} = \frac{q^2}{4\pi c} \left| \hat{\mathbf{R}} \times (\hat{\mathbf{R}} \times \hat{\mathbf{p}}) \right|^2 \Bigg|_{t_r=t-R/c} = \frac{q^2}{4\pi c^3} |\dot{\mathbf{v}}|^2 \frac{\sin^2 \theta}{(1-\beta \cos \theta)^5}$$

$$P_r(\gamma) = \int \frac{dP_r(t_r)}{d\Omega} d\Omega = \frac{2}{3} \frac{q^2}{c^3} |\dot{\mathbf{v}}|^2 \gamma^6 \quad \text{where } \gamma = \frac{1}{\sqrt{1-\beta^2}}$$



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Power distribution for linear acceleration -- continued

$$\frac{dP_r(t_r)}{d\Omega} = \frac{q^2}{4\pi c} \left| \hat{\mathbf{R}} \times (\hat{\mathbf{R}} \times \dot{\beta}) \right|^2 \Bigg|_{t_r=t-R/c} = \frac{q^2}{4\pi c^3} |\dot{\mathbf{v}}|^2 \frac{\sin^2 \theta}{(1-\beta \cos \theta)}$$

$$P_r(t_r) = \int d\Omega \frac{dP_r(t_r)}{d\Omega} = \frac{2}{3} \frac{q^2}{c^3} |\dot{\mathbf{v}}|^2 \gamma^6 \quad \text{where } \gamma \equiv \frac{1}{\sqrt{1-\beta^2}}$$

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Power distribution for circular acceleration

$$\frac{dP_r(t_r)}{d\Omega} = \frac{q^2}{4\pi c} \left| \hat{\mathbf{R}} \times [(\hat{\mathbf{R}} - \beta) \times \dot{\beta}] \right|^2 \Bigg|_{t_r=t-R/c}$$

$$= \frac{q^2}{4\pi c} \frac{|\dot{\beta}|^2 (1-\beta \cdot \hat{\mathbf{R}})^2 - (\hat{\mathbf{R}} \cdot \dot{\beta})^2 (1-\beta^2)}{(1-\beta \cdot \hat{\mathbf{R}})^5} \Bigg|_{t_r=t-R/c}$$

$$P_r(t_r) = \int d\Omega \frac{dP_r(t_r)}{d\Omega} = \frac{2}{3} \frac{q^2}{c^3} |\dot{\mathbf{v}}|^2 \gamma^4$$

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Power distribution for circular acceleration

$$\frac{dP_r(t_r)}{d\Omega} = \frac{q^2}{4\pi c} \frac{|\dot{\beta}|^2 (1-\beta \cdot \hat{\mathbf{R}})^2 - (\hat{\mathbf{R}} \cdot \dot{\beta})^2 (1-\beta^2)}{(1-\beta \cdot \hat{\mathbf{R}})^5} \Bigg|_{t_r=t-R/c}$$

$$= \frac{q^2}{4\pi c^3} \frac{|\dot{\mathbf{v}}|^2}{(1-\beta \cos(\theta))^3} \left(1 - \frac{\cos^2 \theta \sin^2 \phi}{\gamma^2 (1-\beta \cos(\theta))^2} \right)$$

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Spectral composition of electromagnetic radiation

Previously we determined the power distribution from a charged particle:

$$\frac{dP(t)}{d\Omega} = \mathbf{S} \cdot \hat{\mathbf{R}} R^2 = \frac{q^2}{4\pi c} \frac{|\hat{\mathbf{R}} \times [(\hat{\mathbf{R}} - \hat{\mathbf{p}}) \times \dot{\hat{\mathbf{p}}}]|^2}{(1 - \hat{\mathbf{p}} \cdot \hat{\mathbf{R}})^6}$$

$$\equiv |\mathbf{a}(t)|^2$$

$$\text{where } \mathbf{a}(t) \equiv \sqrt{\frac{q^2}{4\pi c} \frac{|\hat{\mathbf{R}} \times [(\hat{\mathbf{R}} - \hat{\mathbf{p}}) \times \dot{\hat{\mathbf{p}}}]|^2}{(1 - \hat{\mathbf{p}} \cdot \hat{\mathbf{R}})^3}}$$

Time integrated power per solid angle:

$$\frac{dW}{d\Omega} = \int_{-\infty}^{\infty} dt \frac{dP(t)}{d\Omega} = \int_{-\infty}^{\infty} dt |\mathbf{a}(t)|^2 = \int_{-\infty}^{\infty} d\omega |\tilde{\mathbf{a}}(\omega)|^2$$

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Spectral composition of electromagnetic radiation -- continued

Time integrated power per solid angle :

$$\frac{dW}{d\Omega} = \int_{-\infty}^{\infty} dt \frac{dP(t)}{d\Omega} = \int_{-\infty}^{\infty} dt |\mathbf{a}(t)|^2 = \int_{-\infty}^{\infty} d\omega |\tilde{\mathbf{a}}(\omega)|^2$$

Fourier amplitude :

$$\tilde{\mathbf{a}}(\omega) \equiv \frac{1}{\sqrt{2\pi}} \int_{-\infty}^{\infty} dt \mathbf{a}(t) e^{i\omega t} \quad \mathbf{a}(t) = \frac{1}{\sqrt{2\pi}} \int_{-\infty}^{\infty} d\omega \tilde{\mathbf{a}}(\omega) e^{-i\omega t}$$

Parseval's theorem

Marc-Antoine Parseval des Chênes 1755-1836

<http://www-history.mcs.st-andrews.ac.uk/Biographies/Parseval.html>

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Spectral composition of electromagnetic radiation -- continued

Consequences of Parseval's analysis :

$$\frac{dW}{d\Omega} = \int_{-\infty}^{\infty} dt \frac{dP(t)}{d\Omega} = \int_{-\infty}^{\infty} dt |\mathbf{a}(t)|^2 = \int_{-\infty}^{\infty} d\omega |\tilde{\mathbf{a}}(\omega)|^2$$

Note that : $\tilde{\mathbf{a}}(\omega) = \tilde{\mathbf{a}}^*(-\omega)$

$$\frac{dW}{d\Omega} = \int_{-\infty}^{\infty} d\omega |\tilde{\mathbf{a}}(\omega)|^2 = \int_0^{\infty} d\omega \left(|\tilde{\mathbf{a}}(\omega)|^2 + |\tilde{\mathbf{a}}(-\omega)|^2 \right) \equiv \int_0^{\infty} d\omega \frac{\partial^2 I}{\partial \Omega \partial \omega}$$

$$\frac{\partial^2 I}{\partial \Omega \partial \omega} \equiv 2 |\tilde{\mathbf{a}}(\omega)|^2$$

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Spectral composition of electromagnetic radiation -- continued

For our case: $\boldsymbol{a}(t) = \sqrt{\frac{q^2}{4\pi c}} \frac{\left| \hat{\mathbf{R}} \times [(\hat{\mathbf{R}} - \boldsymbol{\beta}) \times \dot{\boldsymbol{\beta}}] \right|}{(1 - \boldsymbol{\beta} \cdot \hat{\mathbf{R}})^3} \Big|_{t_r = t - R/c}$

Fourier amplitude:

$$\begin{aligned} \tilde{\boldsymbol{a}}(\omega) &\equiv \frac{1}{\sqrt{2\pi}} \int_{-\infty}^{\infty} dt e^{i\omega t} \boldsymbol{a}(t) \\ &= \sqrt{\frac{q^2}{8\pi^2 c}} \int_{-\infty}^{\infty} dt e^{i\omega t} \frac{\left| \hat{\mathbf{R}} \times [(\hat{\mathbf{R}} - \boldsymbol{\beta}) \times \dot{\boldsymbol{\beta}}] \right|}{(1 - \boldsymbol{\beta} \cdot \hat{\mathbf{R}})^3} \Big|_{t_r = t - R/c} \end{aligned}$$

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Spectral composition of electromagnetic radiation -- continued

Fourier amplitude :

$$\begin{aligned} \tilde{\boldsymbol{a}}(\omega) &\equiv \frac{1}{\sqrt{2\pi}} \int_{-\infty}^{\infty} dt \boldsymbol{a}(t) e^{i\omega t} \\ &= \sqrt{\frac{q^2}{8\pi^2 c}} \int_{-\infty}^{\infty} dt \frac{\left| \hat{\mathbf{R}} \times [(\hat{\mathbf{R}} - \boldsymbol{\beta}) \times \dot{\boldsymbol{\beta}}] \right|}{(1 - \boldsymbol{\beta} \cdot \hat{\mathbf{R}})^3} e^{i\omega t} \Big|_{t_r = t - R/c} \\ &= \sqrt{\frac{q^2}{8\pi^2 c}} \int_{-\infty}^{\infty} dt_r \frac{dt}{dt_r} \frac{\left| \hat{\mathbf{R}} \times [(\hat{\mathbf{R}} - \boldsymbol{\beta}) \times \dot{\boldsymbol{\beta}}] \right|}{(1 - \boldsymbol{\beta} \cdot \hat{\mathbf{R}})^3} e^{i\omega(t_r + R(t_r)/c)} \\ &= \sqrt{\frac{q^2}{8\pi^2 c}} \int_{-\infty}^{\infty} dt_r \frac{\left| \hat{\mathbf{R}} \times [(\hat{\mathbf{R}} - \boldsymbol{\beta}) \times \dot{\boldsymbol{\beta}}] \right|}{(1 - \boldsymbol{\beta} \cdot \hat{\mathbf{R}})^2} e^{i\omega(t_r + R(t_r)/c)} \Big|_{t_r = t - R/c} \end{aligned}$$

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Spectral composition of electromagnetic radiation -- continued

Exact expression :

$$\tilde{\boldsymbol{a}}(\omega) = \sqrt{\frac{q^2}{8\pi^2 c}} \int_{-\infty}^{\infty} dt_r \frac{\left| \hat{\mathbf{R}} \times [(\hat{\mathbf{R}} - \boldsymbol{\beta}) \times \dot{\boldsymbol{\beta}}] \right|}{(1 - \boldsymbol{\beta} \cdot \hat{\mathbf{R}})^2} e^{i\omega(t_r + R(t_r)/c)} \Big|_{t_r = t - R/c}$$

Recall: $\dot{\mathbf{R}}_q(t_r) \equiv \frac{d\mathbf{R}_q(t_r)}{dt_r} \equiv \mathbf{v}$ $\mathbf{R}(t_r) \equiv \mathbf{r} - \mathbf{R}_q(t_r) \equiv \mathbf{R}$

For $r \gg R_q(t_r)$ $R(t_r) \approx r - \hat{\mathbf{r}} \cdot \mathbf{R}_q(t_r)$ where $\hat{\mathbf{r}} \equiv \frac{\mathbf{r}}{r}$

At the same level of approximation : $\hat{\mathbf{R}} \approx \hat{\mathbf{r}}$

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Spectral composition of electromagnetic radiation -- continued

Exact expression:

$$\tilde{\alpha}(\omega) = \sqrt{\frac{q^2}{8\pi^2 c}} \int_{-\infty}^{\infty} dt_r \frac{|\hat{\mathbf{R}} \times [(\hat{\mathbf{R}} - \hat{\beta}) \times \dot{\hat{\beta}}]|}{(1 - \hat{\beta} \cdot \hat{\mathbf{R}})^2} e^{i\omega(t_r + R(t_r)/c)} \Big|_{t_r = t - R/c}$$

Approximate expression:

$$\tilde{\alpha}(\omega) = \sqrt{\frac{q^2}{8\pi^2 c}} e^{i\omega(r/c)} \int_{-\infty}^{\infty} dt_r \frac{|\hat{\mathbf{r}} \times [(\hat{\mathbf{r}} - \hat{\beta}) \times \dot{\hat{\beta}}]|}{(1 - \hat{\beta} \cdot \hat{\mathbf{r}})^2} e^{i\omega(t_r - \hat{\mathbf{r}} \cdot \mathbf{R}_q(t_r)/c)} \Big|_{t_r = t - R/c}$$

Resulting spectral intensity expression:

$$\frac{\partial^2 I}{\partial \omega \partial \Omega} = \frac{q^2}{4\pi^2 c} \left| \int_{-\infty}^{\infty} dt_r \frac{|\hat{\mathbf{r}} \times [(\hat{\mathbf{r}} - \hat{\beta}) \times \dot{\hat{\beta}}]|}{(1 - \hat{\beta} \cdot \hat{\mathbf{r}})^2} e^{i\omega(t_r - \hat{\mathbf{r}} \cdot \mathbf{R}_q(t_r)/c)} \right|^2$$

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Example – radiation from a collinear acceleration burst

$$\frac{\partial^2 I}{\partial \omega \partial \Omega} = \frac{q^2}{4\pi^2 c} \left| \int_{-\infty}^{\infty} dt_r \frac{|\hat{\mathbf{r}} \times [(\hat{\mathbf{r}} - \hat{\beta}) \times \dot{\hat{\beta}}]|}{(1 - \hat{\beta} \cdot \hat{\mathbf{r}})^2} e^{i\omega(t_r - \hat{\mathbf{r}} \cdot \mathbf{R}_q(t_r)/c)} \right|^2$$

Suppose that $\dot{\beta} = \begin{cases} \hat{\beta} \Delta v & 0 < t_r < \tau \\ \frac{c\tau}{ct} & \text{otherwise} \end{cases}$

$$\frac{\partial^2 I}{\partial \omega \partial \Omega} = \frac{q^2}{4\pi^2 c^3} \left| \frac{|\hat{\mathbf{r}} \times [\hat{\mathbf{r}} \times \hat{\beta}] \Delta v|}{(1 - \hat{\beta} \cdot \hat{\mathbf{r}})^2 \tau} \int_0^\tau dt_r e^{i\omega(t_r - \hat{\mathbf{r}} \cdot \mathbf{R}_q(t_r))} \right|^2 \quad \text{Let } \hat{\beta} \cdot \hat{\mathbf{r}} = \beta \cos \theta$$

$$\frac{\partial^2 I}{\partial \omega \partial \Omega} = \frac{q^2}{4\pi^2 c^3} \left(\frac{\Delta v \sin \theta}{(1 - \beta \cos \theta)^2} \frac{\sin(\omega \tau (1 - \beta \cos \theta)/2)}{(\omega \tau (1 - \beta \cos \theta)/2)} \right)^2$$

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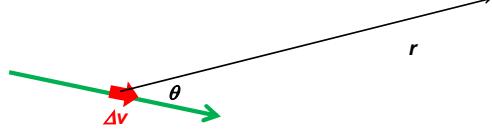
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Example:

Suppose that $\dot{\beta} = \begin{cases} \hat{\beta} \Delta v & 0 < t_r < \tau \\ \frac{c\tau}{ct} & \text{otherwise} \end{cases}$

$$\frac{\partial^2 I}{\partial \omega \partial \Omega} = \frac{q^2}{4\pi^2 c^3} \left(\frac{\Delta v \sin \theta}{(1 - \beta \cos \theta)^2} \frac{\sin(\omega \tau (1 - \beta \cos \theta)/2)}{(\omega \tau (1 - \beta \cos \theta)/2)} \right)^2$$



Example: "Bremsstrahlung" radiation

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Spectral composition of electromagnetic radiation -- continued

Alternative expression --

It can be shown that:

$$\frac{\hat{\mathbf{r}} \times [(\hat{\mathbf{r}} - \hat{\mathbf{p}}) \times \hat{\mathbf{p}}]}{(1 - \hat{\mathbf{p}} \cdot \hat{\mathbf{r}})^2} = \frac{d}{dt_r} \left(\frac{\hat{\mathbf{r}} \times (\hat{\mathbf{r}} \times \hat{\mathbf{p}})}{(1 - \hat{\mathbf{p}} \cdot \hat{\mathbf{r}})} \right)$$

Integration by parts and assumptions about the integration limit behaviors shows that the spectral intensity depends on the following integral:

$$\frac{\partial^2 I}{\partial \omega \partial \Omega} = \frac{q^2 \omega^2}{4\pi^2 c} \left| \int_{-\infty}^{\infty} dt_r [\hat{\mathbf{r}} \times (\hat{\mathbf{r}} \times \hat{\mathbf{p}}(t_r))] e^{i\omega(t_r - \hat{\mathbf{r}} \cdot \mathbf{R}_q(t_r)/c)} \right|^2$$

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