

Solutions to Problems 1

2. **It is possible that the universe has small extra dimensions. If so, we should be able to detect them if we use particles with wavelength shorter than the scale L of the extra dimension. Having performed experiments with 4 TeV protons without seeing hints of extra dimensions, estimate the maximum size this extra dimension might be in meters.**

We simply use the relationship $\lambda p = 2\pi\hbar = 2\pi$. First we need to find the momentum, which for relativistic neutrinos is effectively the same as the energy. Hence

$$\lambda = \frac{2\pi}{p} = \frac{2\pi(0.197 \text{ GeV} \cdot \text{fm})}{4000.0 \text{ GeV}} = 3.09 \times 10^{-4} \text{ fm} = 3.09 \times 10^{-19} \text{ m}.$$

I have no idea exactly what the real limit is, but obviously it is pretty short.

5. **By July 4, 2012, approximately 5 fb^{-1} of integrated luminosity at $\sqrt{s} = 8 \text{ TeV}$ had been analyzed by the CMS and ATLAS detectors. How many Higgs particles were produced? The main signal seen was from the process $H \rightarrow \gamma\gamma$. The branching ratio for this decay is about 0.25%. How many $H \rightarrow \gamma\gamma$ events should have been seen by each experiment?**

We need the cross-section, which we get from the link in the previous problem. Looking at <http://arxiv.org/abs/1012.0530>, we see from table 3 that the cross section for Higgs production at $\sqrt{s} = 8 \text{ TeV}$ for a mass of 125 GeV is about 19.81 pb and for a mass of 130 GeV is about 18.34 pb. The numbers look vaguely linear in this region, so using a linear fit, we estimate at 126 GeV the cross-section is about 19.52 pb. We multiply this by the integrated luminosity to obtain

$$N_H = \sigma_H \int L dt = (19.52 \text{ pb})(5.00 \text{ fb}^{-1}) = 97.6 \frac{10^{-12} \text{ b}}{10^{-15} \text{ b}} \approx 97,600.$$

According to table 6, the branching ratio to photons is not 0.25%, but 0.23%, so the number of decays is

$$N(H \rightarrow \gamma\gamma) = N_H BR(H \rightarrow \gamma\gamma) = (97,600)(0.0023) = 224.$$

This is the approximate number of events expected in each of the two detectors.

- 7. Look up the total lifetime of the π^+ and K^+ mesons (summary tables, mesons). What would the rate Γ in GeV be? Then look up the branching ratios in each case to decay to $\mu^+\nu_\mu$. Find the partial rates $\Gamma(\pi^+ \rightarrow \mu^+\nu_\mu)$ and $\Gamma(K^+ \rightarrow \mu^+\nu_\mu)$ in each case, and their ratio.**

According to the particle data book, $\tau_\pi = 2.60 \times 10^{-8}$ s and $\tau_K = 1.23 \times 10^{-8}$ s. As noted in the text, $\Gamma = \tau^{-1}$, so we have

$$\Gamma_\pi = \tau_\pi^{-1} = \frac{\hbar}{\tau_\pi} = \frac{6.582 \times 10^{-16} \text{ eV} \cdot \text{s}}{2.60 \times 10^{-8} \text{ s}} = 2.53 \times 10^{-8} \text{ eV} = 2.53 \times 10^{-17} \text{ GeV},$$

$$\Gamma_K = \tau_K^{-1} = \frac{\hbar}{\tau_K} = \frac{6.582 \times 10^{-16} \text{ eV} \cdot \text{s}}{1.238 \times 10^{-8} \text{ s}} = 5.32 \times 10^{-8} \text{ eV} = 5.32 \times 10^{-17} \text{ GeV}.$$

The branching ratio to $\mu^+\nu_\mu$ is 99.99% for the pion and 63.55% for the kaon. We multiply by these numbers to get the partial decay rates:

$$\Gamma(\pi^+ \rightarrow \mu^+\nu_\mu) = \Gamma_\pi BR(\pi^+ \rightarrow \mu^+\nu_\mu) = (2.53 \times 10^{-17} \text{ GeV})(0.9999) = 2.53 \times 10^{-17} \text{ GeV},$$

$$\Gamma(K^+ \rightarrow \mu^+\nu_\mu) = \Gamma_K BR(K^+ \rightarrow \mu^+\nu_\mu) = (5.32 \times 10^{-17} \text{ GeV})(0.6355) = 3.38 \times 10^{-17} \text{ GeV},$$

$$\frac{\Gamma(\pi^+ \rightarrow \mu^+\nu_\mu)}{\Gamma(K^+ \rightarrow \mu^+\nu_\mu)} = \frac{2.53}{3.38} = 0.749.$$

As we will later realize, the denominator is larger overwhelmingly because of the higher mass of the kaon.

- 8. Find a dimensionless combination of e , ε_0 , \hbar and c . Then, setting $\varepsilon_0 = \hbar = c = 1$, find the dimensionless value of the fundamental charge e .**

The values of these constants are

$$e = 1.602 \times 10^{-19} \text{ C},$$

$$\varepsilon_0 = 8.854 \times 10^{-12} \text{ C}^2 \text{s}^2 / \text{m}^3 / \text{kg},$$

$$\hbar = 1.055 \times 10^{-34} \text{ kg} \cdot \text{m}^2 / \text{s},$$

$$c = 2.998 \times 10^8 \text{ m/s}.$$

We note that e has C in it, and ε_0 has C^2 , so if we square the former and divide by the latter, the C will cancel out. This will leave kg in the numerator, but if we divide by \hbar , that will go away, and it isn't hard to see that if you then divide by c , you get something dimensionless. So we have

$$\frac{e^2}{\epsilon_0 \hbar c} = \frac{(1.602 \times 10^{-19} \text{ C})^2}{(8.854 \times 10^{-12} \text{ C}^2 \text{ s}^2 / \text{m}^3 / \text{kg})(1.055 \times 10^{-34} \text{ kg} \cdot \text{m}^2 / \text{s})(2.998 \times 10^8 \text{ m/s})} = 0.09164.$$

The quantity on the right is true in any units, but in particle physics units, the factors in the denominator on the left are 1, and this formula simplifies to $e^2 = 0.09245$, or taking the square root,

$$e = \sqrt{0.09164} = 0.3027.$$

We could have gotten the same answer starting from the equation on the inside front cover $e^2/4\pi = 1/137.04$

- 9. By using a suitable combination of \hbar and c , write Newton's constant in the form $G_N = M_p^n$, where M_p has units of mass or energy, and is called the *Planck mass*. Determine the integer n and the value of M_p in GeV. If a proton collider were operating at $E = M_p$ and used $B = 10$ T magnets, what would be its radius in light-years?**

The relevant constants are

$$\begin{aligned} G_N &= 6.674 \times 10^{-11} \text{ m}^3 / \text{kg} / \text{s}^2, \\ \hbar &= 1.055 \times 10^{-34} \text{ kg} \cdot \text{m}^2 / \text{s}, \\ c &= 2.998 \times 10^8 \text{ m/s}. \end{aligned}$$

Let's start by writing it as a mass, which we can do if we get rid of all the meters and seconds. It is clear that G_N/\hbar will have only one factor of m in the numerator and one factor of s in the denominator, so if we then divide by c we get

$$\frac{G_N}{\hbar c} = \frac{6.674 \times 10^{-11} \text{ m}^3 / \text{kg} / \text{s}^2}{(1.055 \times 10^{-34} \text{ kg} \cdot \text{m}^2 / \text{s})(2.998 \times 10^8 \text{ m/s})} = 2.11 \times 10^{15} \text{ kg}^{-2}.$$

Since it is mass to the minus two, we write it in the form M_p^{-2} , so that by definition,

$$\begin{aligned} M_p^{-2} &\equiv \frac{G_N}{\hbar c} = 2.11 \times 10^{15} \text{ kg}^{-2}, \\ M_p &= (2.11 \times 10^{15} \text{ kg}^{-2})^{-1/2} = 2.177 \times 10^{-8} \text{ kg}. \end{aligned}$$

In particle physics units, we would write this as $G = M_p^{-2}$. We then convert this to an energy by using the conversion $1 \text{ kg} = 5.6 \times 10^{26} \text{ GeV}$, so

$$M_p = (2.177 \times 10^{-8} \text{ kg})(5.6 \times 10^{26} \text{ GeV/kg}) = 1.22 \times 10^{19} \text{ GeV}.$$

We can then find how large a collider we need to reach this energy using eq. (1.5). The momentum is effectively the same as the energy, and we are working in units where $c = 1$, so

$$p = \left(\frac{B}{T}\right) \left(\frac{R}{\text{km}}\right) (299.8 \text{ GeV}),$$

$$R = \frac{1.22 \times 10^{19}}{299.8 \cdot 10} \text{ km} = 4.066 \times 10^{15} \text{ km} = \frac{(4.066 \times 10^{15} \text{ km})c}{(2.998 \times 10^5 \text{ km/s})(3.156 \times 10^7 \text{ s/y})} \approx 430 \text{ ly}.$$

It is inconceivable we will ever reach this energy without new technology.

14. Perform the following integrals:

(a) $\int_0^\infty E^n \delta(E^2 - \mathbf{p}^2 - m^2) dE$ for arbitrary \mathbf{n} .

The argument of the delta function vanishes at $E = \sqrt{\mathbf{p}^2 + m^2}$, so

$$\int_0^\infty E^n \delta(E^2 - \mathbf{p}^2 - m^2) dE = \frac{E^n}{2E} \Big|_{E=\sqrt{\mathbf{p}^2+m^2}} = \frac{1}{2} (\mathbf{p}^2 + m^2)^{\frac{1}{2}(n-1)}.$$

(b) $\int_0^\infty dE_1 \int_0^\infty dE_2 \theta(\frac{1}{2}m - E_1) \theta(\frac{1}{2}m - E_2) \theta(E_1 + E_2 - \frac{1}{2}m) (\frac{1}{2}m^2 E_1 - mE_1^2)$

As we argued in class, the first two Heaviside functions restrict the upper limit on the energy integrals to each be $\frac{1}{2}m$, and the third one demands that $E_1 + E_2 > \frac{1}{2}m$. If we let the inner integral be the E_2 integral, then we have

$$\begin{aligned} & \int_0^\infty dE_1 \int_0^\infty dE_2 \theta(\frac{1}{2}m - E_1) \theta(\frac{1}{2}m - E_2) \theta(E_1 + E_2 - \frac{1}{2}m) (\frac{1}{2}m^2 E_1 - mE_1^2) \\ &= \int_0^{\frac{1}{2}m} dE_1 \int_{\frac{1}{2}m - E_1}^{\frac{1}{2}m} dE_2 (\frac{1}{2}m^2 E_1 - mE_1^2) = \int_0^{\frac{1}{2}m} dE_1 (\frac{1}{2}m^2 E_1 - mE_1^2) E_2 \Big|_{\frac{1}{2}m - E_1}^{\frac{1}{2}m} \\ &= \int_0^{\frac{1}{2}m} dE_1 (\frac{1}{2}m^2 E_1^2 - mE_1^3) = \left(\frac{1}{6}m^2 E_1^3 - \frac{1}{4}mE_1^4\right) \Big|_0^{\frac{1}{2}m} = \frac{1}{48}m^5 - \frac{1}{64}m^5 = \frac{1}{192}m^5. \end{aligned}$$

$$(c) \int \left[(1 - 2 \sin^2 \theta_w)^2 E^4 + \sin^4 \theta_w E^4 (1 + \cos \theta)^2 \right] d\Omega \quad (\text{note: } \theta_w \text{ is a constant})$$

There is an integral $\int_0^{2\pi} d\phi = 2\pi$, which is easy, and an integral over θ , which I like to do by changing to $z = \cos \theta$

$$\begin{aligned} & \int \left[(1 - 2 \sin^2 \theta_w)^2 E^4 + \sin^4 \theta_w E^4 (1 + \cos \theta)^2 \right] d\Omega \\ &= 2\pi E^4 \int_{-1}^1 dz \left[(1 - 2 \sin^2 \theta_w)^2 + \sin^4 \theta_w (1 + 2z + z^2) \right] \\ &= 2\pi E^4 \left[(1 - 2 \sin^2 \theta_w)^2 z + \sin^4 \theta_w \left(z + z^2 + \frac{1}{3} z^3 \right) \right]_{-1}^1 \\ &= 2\pi E^4 \left[2(1 - 2 \sin^2 \theta_w)^2 + \sin^4 \theta_w \left(2 + 0 + \frac{2}{3} \right) \right] = 4\pi E^4 \left(1 - 4 \sin^2 \theta_w + \frac{16}{3} \sin^4 \theta_w \right). \end{aligned}$$

It is unclear what further simplification is desirable.

$$(d) \int \frac{g^4 p \cos^2 \theta}{128\pi^2 E (E^2 - p^2 \cos^2 \theta)} d\Omega$$

$$\begin{aligned} \int \frac{g^4 p \cos^2 \theta}{128\pi^2 E (E^2 - p^2 \cos^2 \theta)} d\Omega &= \frac{2\pi g^4}{128\pi^2 E p} \int_{-1}^1 \frac{p^2 \cos^2 \theta d \cos \theta}{E^2 - p^2 \cos^2 \theta} \\ &= \frac{g^4}{32\pi E p} \int_0^1 \left[-1 + \frac{E^2}{E^2 - p^2 \cos^2 \theta} \right] d \cos \theta = \frac{g^4}{32\pi E p} \left[-1 + \frac{1}{2} \int_0^1 \left(\frac{E d \cos \theta}{E + p \cos \theta} + \frac{E d \cos \theta}{E - p \cos \theta} \right) \right] \\ &= \frac{g^4}{32\pi E p} \left\{ -1 + \frac{E}{2p} \left[\ln(E + p \cos \theta) - \ln(E - p \cos \theta) \right]_0^1 \right\} = \frac{g^4}{32\pi p^2} \left\{ \frac{1}{2} \ln \left(\frac{E + p}{E - p} \right) - \frac{p}{E} \right\}. \end{aligned}$$

The logarithm term can be written more succinctly as $\tanh^{-1}(p/E)$ if we want.