

Physics 742 – Graduate Quantum Mechanics 2
Solutions to Chapter 15

6. [25] A hydrogen atom is in interstellar space in the 1s state, but not in the true ground state ($F = 0$), but rather in the hyperfine excited state ($F=1$), specifically in the state $|\phi_I\rangle = |n, l, j, F, m_F\rangle = |1, 0, \frac{1}{2}, 1, 0\rangle$. It is going to transition to the true ground state $|\phi_F\rangle = |n, l, j, F, m_F\rangle = |1, 0, \frac{1}{2}, 0, 0\rangle$ via a magnetic dipole interaction.
- (a) [5] Write out the initial and final states in terms of the explicit spin states of the electron and proton $|\pm, \pm\rangle$. Find all non-zero components of the matrix $\langle\phi_F|\mathbf{S}|\phi_I\rangle$, where \mathbf{S} is the electron spin operator.

The explicit form for these spin states can be found, for example, eq. (8.21). We therefore have

$$|1, 0\rangle = \frac{1}{\sqrt{2}}(|+-\rangle + |-+\rangle), \quad |0, 0\rangle = \frac{1}{\sqrt{2}}(|+-\rangle - |-+\rangle).$$

The spin operator acting on the initial state yields

$$\begin{aligned} S_x |10\rangle &= \frac{1}{2} \hbar \sigma_x \frac{1}{\sqrt{2}}(|+-\rangle + |-+\rangle) = \frac{1}{\sqrt{8}} \hbar (|--\rangle + |++\rangle), \\ S_y |10\rangle &= \frac{1}{2} \hbar \sigma_y \frac{1}{\sqrt{2}}(|+-\rangle + |-+\rangle) = \frac{1}{\sqrt{8}} \hbar (i|--\rangle - i|++\rangle), \\ S_z |10\rangle &= \frac{1}{2} \hbar \sigma_z \frac{1}{\sqrt{2}}(|+-\rangle + |-+\rangle) = \frac{1}{\sqrt{8}} \hbar (|+-\rangle - |-+\rangle) = \frac{1}{2} \hbar |00\rangle. \end{aligned}$$

Therefore, the only non-zero matrix element is S_z , so $\langle\phi_F|\mathbf{S}|\phi_I\rangle = \frac{1}{2} \hbar \hat{\mathbf{z}}$.

- (b) [8] Show that the rate for this transition for a wave going in a specific direction with a definite polarization is given by $\Gamma = 4\pi^2 m^{-2} \omega^{-2} \alpha \mathcal{I} |(\mathbf{k} \times \boldsymbol{\varepsilon}) \cdot \mathbf{S}_{FI}|^2 \delta(E_F - E_I + \hbar\omega)$.

Our starting point is the equation in the middle of page 283:

$$W_{FI} = A_0 e \left\{ -\omega_{FI} \sum_i \langle\phi_F|(\mathbf{k} \cdot \mathbf{R}_i)(\boldsymbol{\varepsilon} \cdot \mathbf{R}_i)|\phi_I\rangle + \frac{i}{m} (\mathbf{k} \times \boldsymbol{\varepsilon}) \cdot \langle\phi_F|(\frac{1}{2} \mathbf{L} + \mathbf{S})|\phi_I\rangle \right\}$$

The first term won't contribute in this case, since the operator can't connect the ground state with the ground state. The operator \mathbf{L} vanishes on an s -electron, so the only relevant term is

$$W_{FI} = A_0 e \frac{i}{m} (\mathbf{k} \times \boldsymbol{\varepsilon}) \cdot \langle\phi_F|\mathbf{S}|\phi_I\rangle = A_0 e \frac{i}{m} (\mathbf{k} \times \boldsymbol{\varepsilon}) \cdot \mathbf{S}_{FI}$$

We now technically need to use equation (15.22) (with the roles of I and F reversed) to give

$$\Gamma(I \rightarrow F) = \frac{2\pi}{\hbar} |W_{FI}^\dagger|^2 \delta(E_F - E_I + \hbar\omega) = \frac{2\pi A_0^2 e^2}{\hbar m^2} |(\mathbf{k} \times \boldsymbol{\varepsilon}) \cdot \mathbf{S}_{FI}|^2 \delta(E_F - E_I + \hbar\omega)$$

In a manner very similar to the notes, we use equation (15.28), together with $\alpha = k_e e^2 / \hbar c$ to rewrite this as

$$\Gamma(I \rightarrow F) = \frac{4\pi^2 \mathcal{I} k_e e^2}{\hbar m^2 \omega^2 c} |(\mathbf{k} \times \boldsymbol{\varepsilon}) \cdot \mathbf{S}_{FI}|^2 \delta(E_F - E_I + \hbar\omega) = \frac{4\pi^2 \alpha \mathcal{I}}{m^2 \omega^2} |(\mathbf{k} \times \boldsymbol{\varepsilon}) \cdot \mathbf{S}_{FI}|^2 \delta(E_F - E_I + \hbar\omega)$$

(c) [7] Show that for a wave going in a random direction with random polarization, this simplifies to $\Gamma(I \rightarrow F) = \frac{4}{3} \pi^2 \alpha \mathcal{I} |\mathbf{S}_{FI}|^2 \delta(E_f - E_i + \hbar\omega) / m^2 c^2$.

To simplify, let's assume that \mathbf{S}_{FI} is in the z -direction, and the direction of \mathbf{k} is at an angle θ compared to \mathbf{S}_{FI} . The simplest way to tackle this is to first rewrite the triple product as

$$(\mathbf{k} \times \boldsymbol{\varepsilon}) \cdot \mathbf{S}_{FI} = (\mathbf{S}_{FI} \times \mathbf{k}) \cdot \boldsymbol{\varepsilon}.$$

Now, the cross product is perpendicular to both \mathbf{S}_{FI} and \mathbf{k} , and has magnitude

$$|\mathbf{S}_{FI} \times \mathbf{k}| = |\mathbf{S}_{FI}| k \sin \theta$$

The polarization must be perpendicular to \mathbf{k} . One polarization is in the same plane as \mathbf{S}_{FI} and \mathbf{k} , and this will give no contribution to $(\mathbf{S}_{FI} \times \mathbf{k}) \cdot \boldsymbol{\varepsilon}$. The other will be in the same direction as $\mathbf{S}_{FI} \times \mathbf{k}$, so $(\mathbf{S}_{FI} \times \mathbf{k}) \cdot \boldsymbol{\varepsilon} = |\mathbf{S}_{FI}| k \sin \theta$ (up to an arbitrary sign). Hence we have

$$\Gamma_{\text{unpol}} = \frac{1}{2} \sin^2 \theta \frac{4\pi^2 \alpha \mathcal{I} k^2}{m^2 \omega^2} |\mathbf{S}_{FI}|^2 \delta(E_f - E_i + \hbar\omega) = \frac{2\pi^2 \alpha \mathcal{I}}{m^2 c^2} |\mathbf{S}_{FI}|^2 \delta(E_f - E_i + \hbar\omega) \sin^2 \theta.$$

We now must average this over all angles, which gives us a factor of

$$\langle \sin^2 \theta \rangle = \int \frac{d\Omega}{4\pi} \sin^2 \theta = \frac{1}{4\pi} \int_0^{2\pi} d\phi \int_{-1}^1 (1 - \cos^2 \theta) d \cos \theta = \frac{1}{2} \left(\cos \theta - \frac{1}{3} \cos^3 \theta \right) \Big|_{-1}^1 = \frac{2}{3}.$$

for a final answer of

$$\Gamma_{\text{random}} = \frac{4\pi^2 \alpha \mathcal{I}}{3m^2 c^2} |\mathbf{S}_{FI}|^2 \delta(E_f - E_i + \hbar\omega).$$

(d) [5] For low frequencies, the cosmic microwave background intensity is

$$\mathcal{I}(\omega) = k_B T \omega^2 / \pi^2 c^2 \text{ where } k_B \text{ is Boltzman's constant and } T \text{ is the temperature.}$$

Integrate the flipping rate over frequency. Find the mean time Γ^{-1} for a hydrogen atom to reverse itself in a background temperature $T = 2.73 \text{ K}$ for

$$\omega_{FI} = -2\pi(1.420 \text{ GHz}).$$

First, we take a factor of \hbar out of the delta function to yield

$$\Gamma_{\text{random}} = \frac{4\pi^2 \alpha \mathcal{I}}{3m^2 c^2 \hbar} |\mathbf{S}_{FI}|^2 \delta(\omega_{FI} + \omega).$$

Obviously, the frequency we need is $\omega = -\omega_{FI} = 2\pi(1.420 \text{ GHz})$. If we have a continuous distribution of intensities $\mathcal{I}(\omega)$, this changes to

$$\Gamma_{\text{random}} = \frac{4\pi^2\alpha}{3m^2c^2\hbar} |\mathbf{S}_{FI}|^2 \mathcal{I}(|\omega_{FI}|).$$

Substituting our various formulas in, and then making numerical substitutions, we find

$$\begin{aligned} \Gamma_{\text{random}} &= \frac{4\pi^2\alpha}{3m^2c^2\hbar} \left(\frac{\hbar}{2}\right)^2 \frac{k_B T \omega^2}{\pi^2 c^2} = \frac{k_B T \alpha \hbar \omega^2}{3m^2 c^4} \\ &= \frac{(1.381 \times 10^{-23} \text{ J/K})(2.73 \text{ K})(1.054 \times 10^{-34} \text{ J}\cdot\text{s})(2\pi(1.420 \times 10^9 \text{ s}^{-1}))^2}{3.137(9.1094 \times 10^{-31} \text{ kg})^2 (2.998 \times 10^8 \text{ m/s})^4} \\ &= 1.148 \times 10^{-13} \text{ s}^{-1} = 3.623 \times 10^{-6} \text{ y}^{-1}. \end{aligned}$$

Taking the reciprocal, the mean flipping time is 276,000 years.